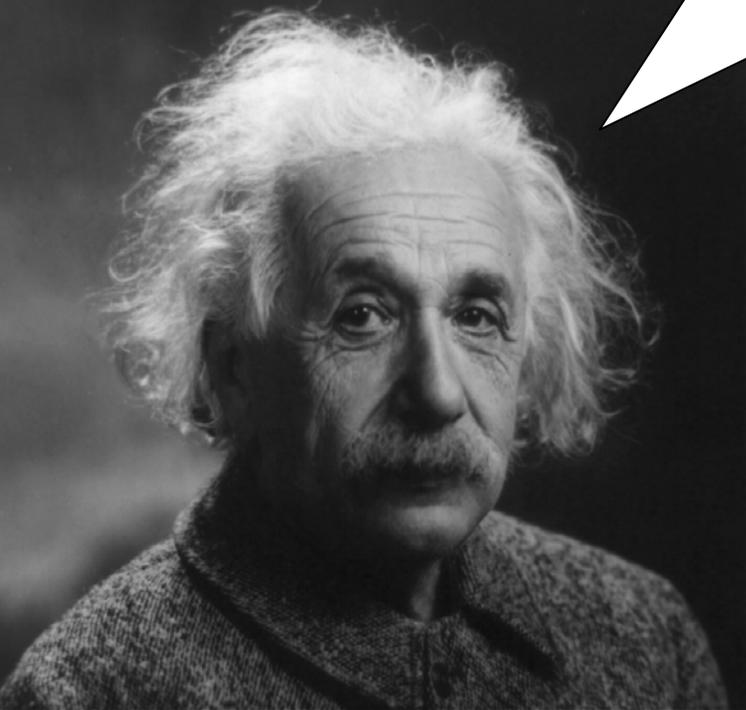


Welcome
back
to 8.033!



Relativistic dynamics summary:

Dynamics toolbox: formula summary

- Mass-energy unification:

$$E = mc^2 = m_0\gamma c^2$$

- Momentum 4-vector:

$$\mathbf{P} \equiv m_0\mathbf{U} = \begin{pmatrix} p_x \\ p_y \\ p_z \\ E/c \end{pmatrix}$$

- Energy formula:

$$E = \sqrt{(m_0c^2)^2 + (cp)^2}$$

- Velocity formula:

$$\beta = \frac{cp}{E}$$

MIT Course 8.033, Fall 2006, Lecture 12

Max Tegmark

Today's topics:

- Atomic, nuclear & particle physics
- Parallel & transverse acceleration & force
- Particle accelerators

FOCUS OF PARTICLE PHYSICS COMPONENT OF 8.033:

- 1) Give you basic overview of atomic, nuclear & particle physics so that you can
 - 1) Better see the big picture
 - 2) Get more out of popular talks and articles
 - 3) Pass the “cocktail party test”
- 2) Apply two core relativity results:
 - 1) Mass-energy unification
 - 2) Energy-momentum conservation

Atomic physics

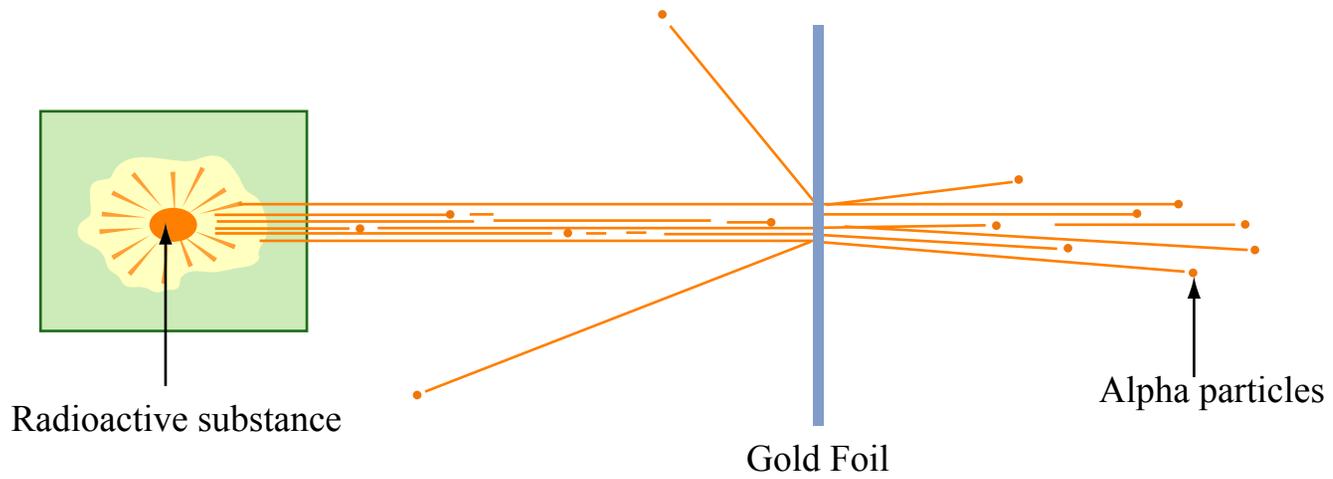


Figure by MIT OCW.

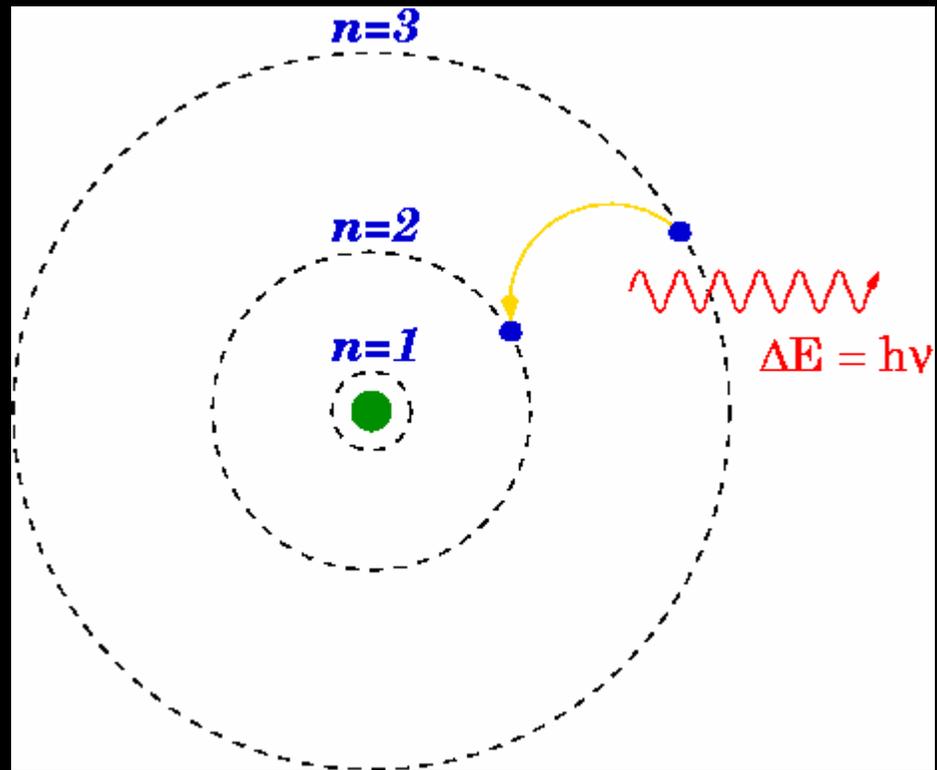


Image courtesy of Wikipedia.

1																	2
H																	He
3	4											5	6	7	8	9	10
Li	Be											B	C	N	O	F	Ne
11	12											13	14	15	16	17	18
Na	Mg											Al	Si	P	S	Cl	Ar
19	20	21	22	23	24	25	26	27	28	29	30	31	32	33	34	35	36
K	Ca	Sc	Ti	V	Cr	Mn	Fe	Co	Ni	Cu	Zn	Ga	Ge	As	Se	Br	Kr
37	38	39	40	41	42	43	44	45	46	47	48	49	50	51	52	53	54
Rb	Sr	Y	Zr	Nb	Mo	Tc	Ru	Rh	Pd	Ag	Cd	In	Sn	Sb	Te	I	Xe
55	56	71	72	73	74	75	76	77	78	79	80	81	82	83	84	85	86
Cs	Ba	Lu	Hf	Ta	W	Re	Os	Ir	Pt	Au	Hg	Tl	Pb	Bi	Po	At	Rn
87	88	103	104	105	106	107	108	109	110	111	112						
Fr	Ra	Lr	Rf	Db	Sg	Bh	Hs	Mt	Uun	Uuu	Uub						

#118 just confirmed! → ◆

57	58	59	60	61	62	63	64	65	66	67	68	69	70
La	Ce	Pr	Nd	Pm	Sm	Eu	Gd	Tb	Dy	Ho	Er	Tm	Yb
89	90	91	92	93	94	95	96	97	98	99	100	101	102
Ac	Th	Pa	U	Np	Pu	Am	Cm	Bk	Cf	Es	Fm	Md	No



Spectral line experiment

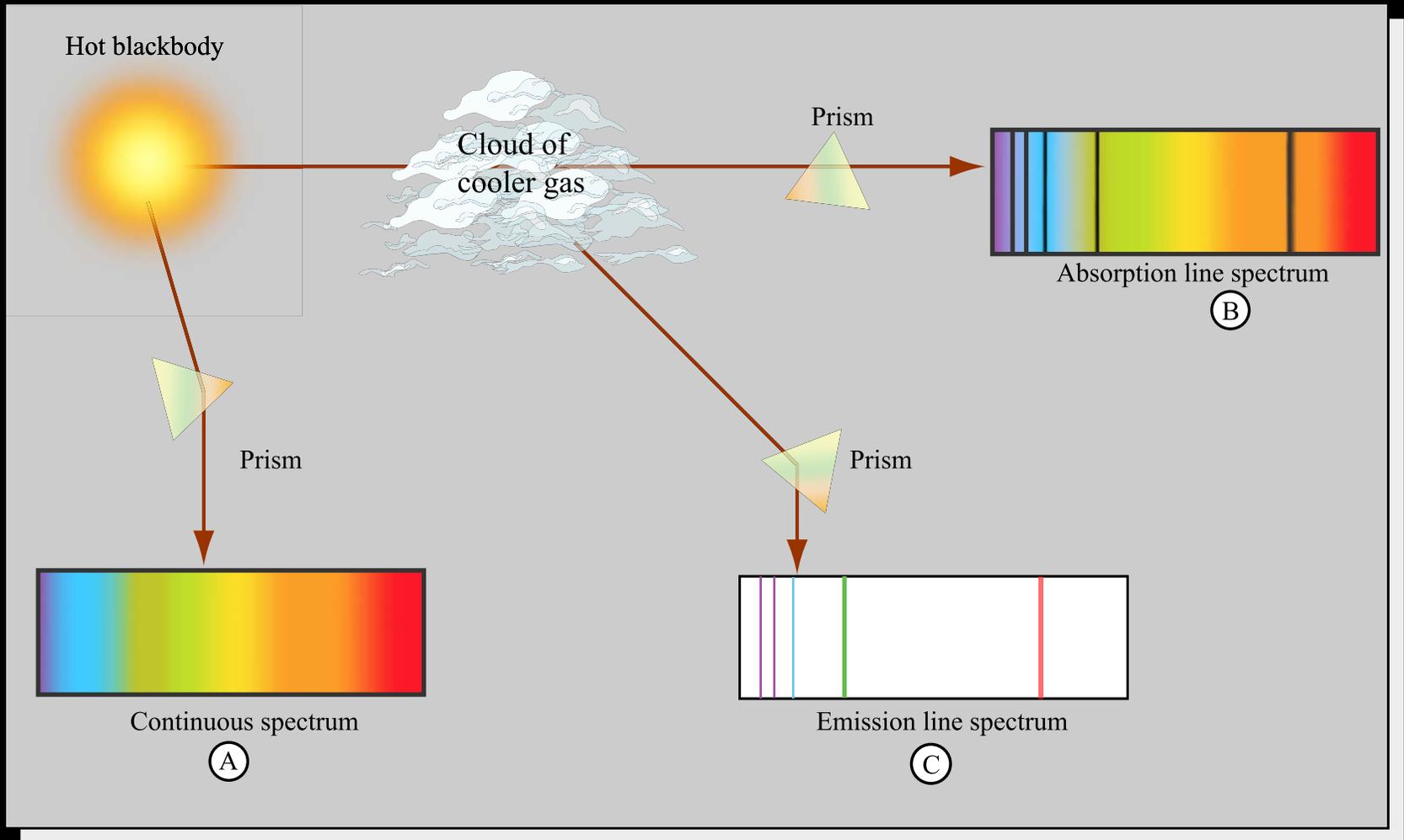


Figure by MIT OCW.

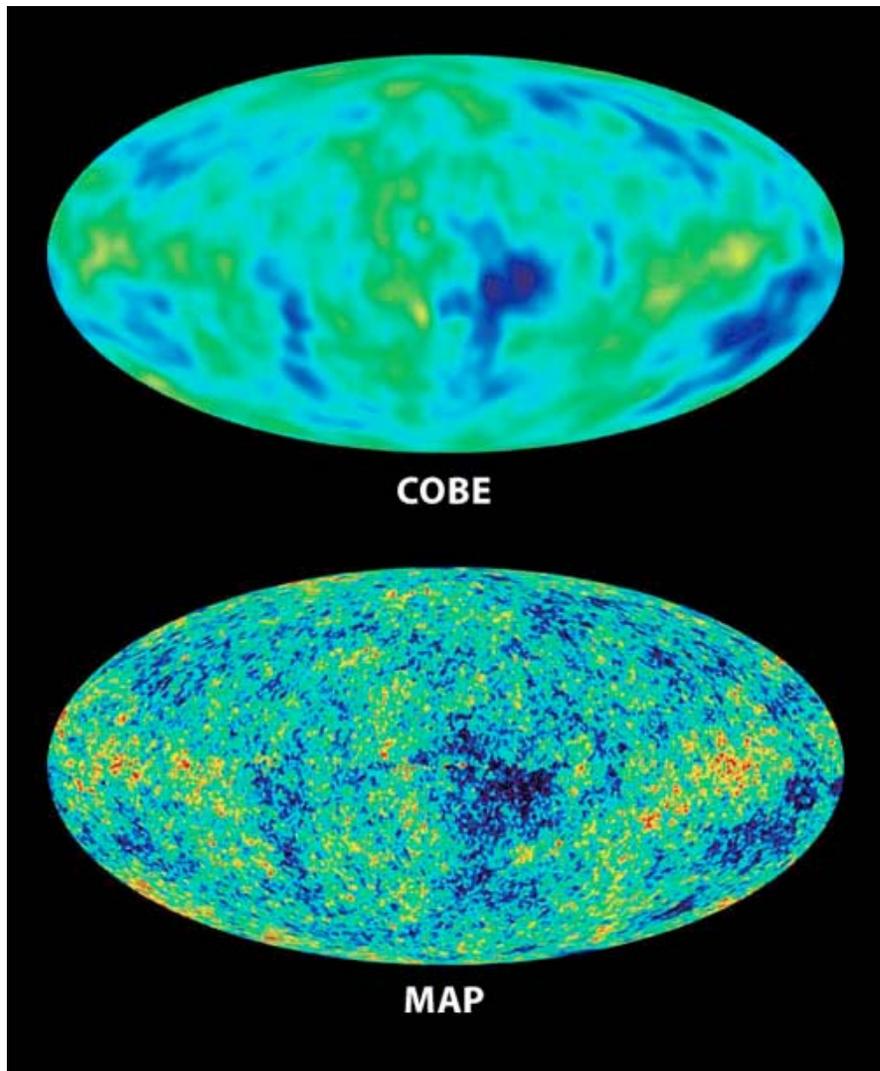


Image courtesy of NASA.

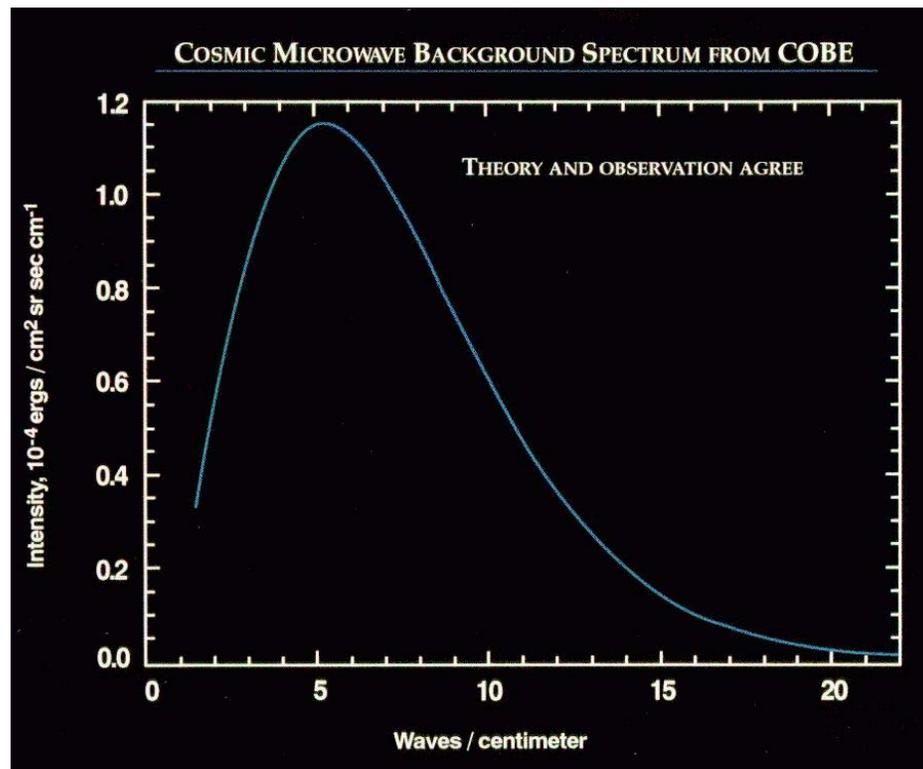


Image courtesy of NASA.

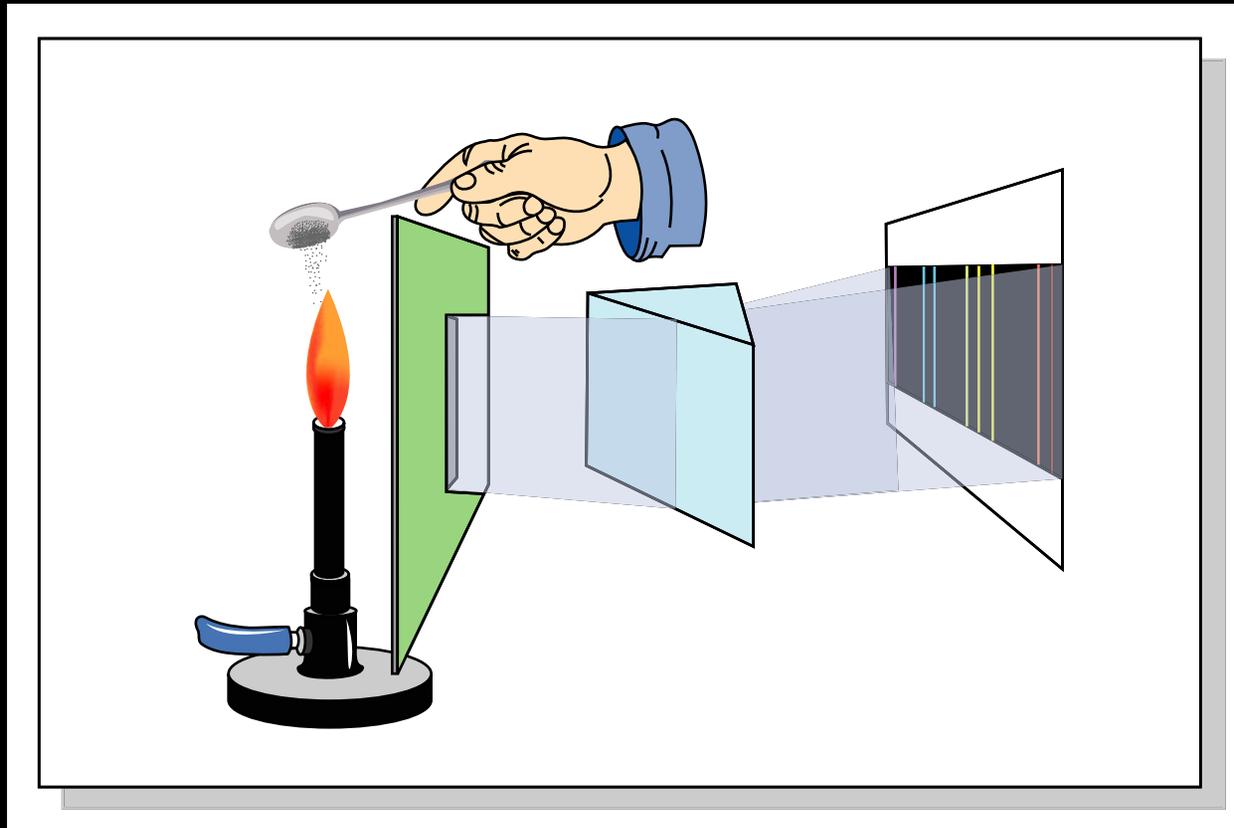


Figure by MIT OCW.

Emission spectrum of hydrogen:

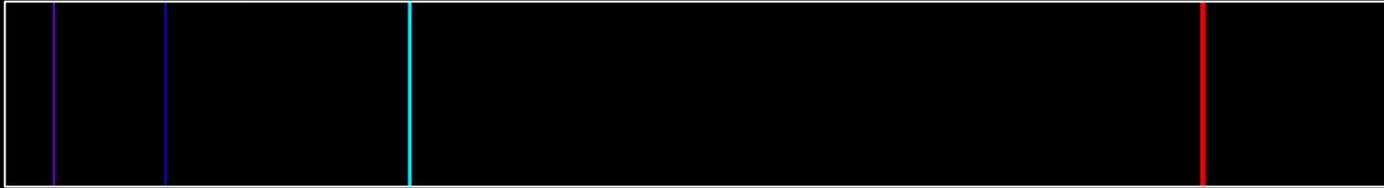


Image courtesy of Wikipedia.

Emission spectrum of iron:

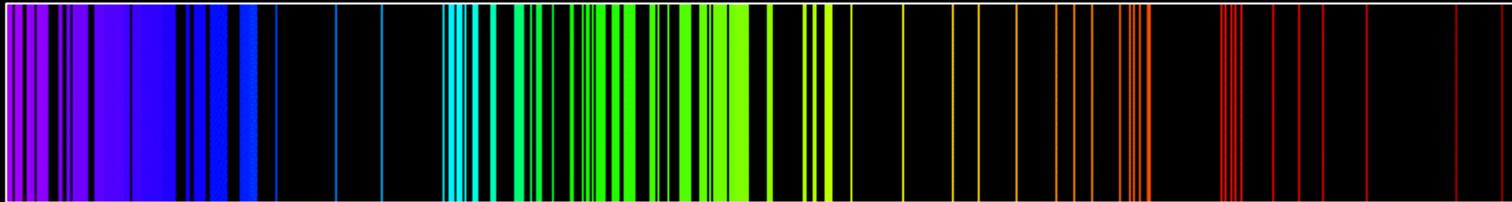


Image courtesy of Wikipedia.

How do we
know what stars
are made of?

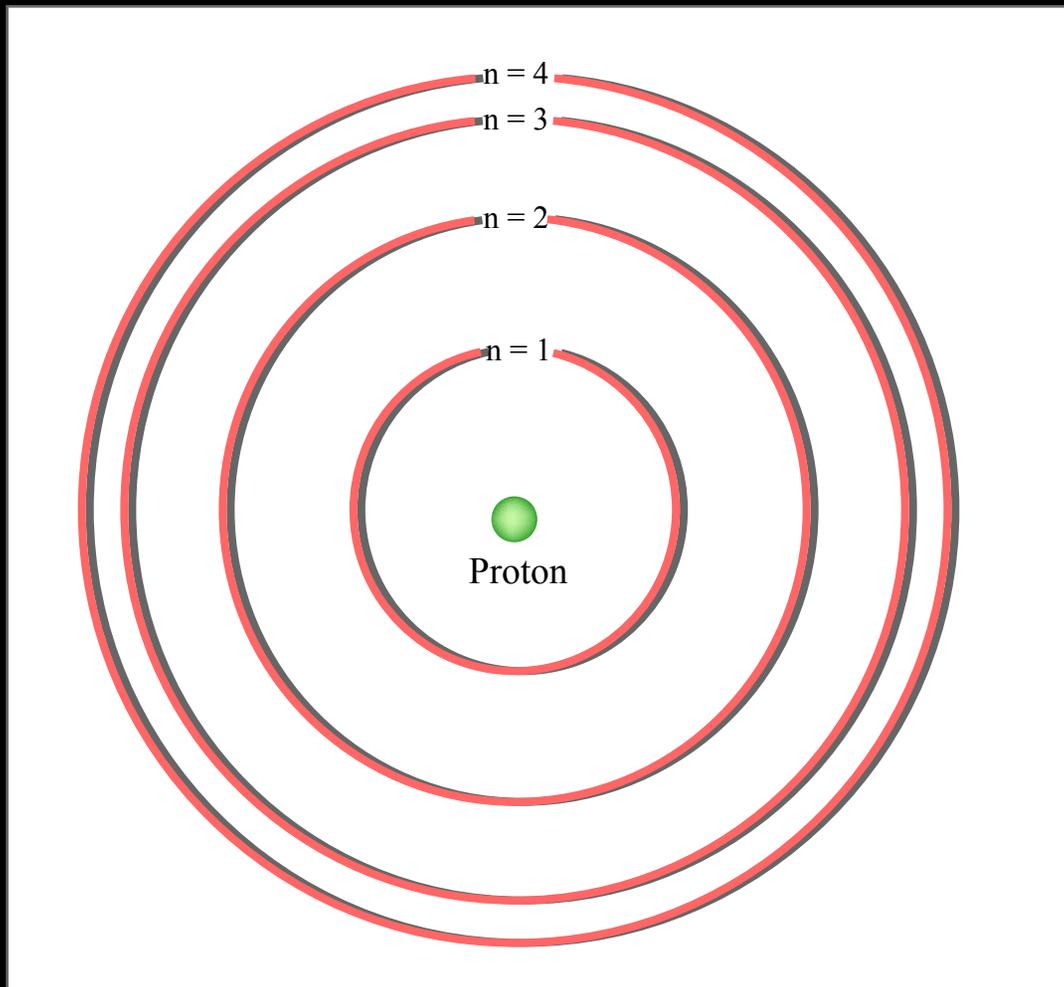
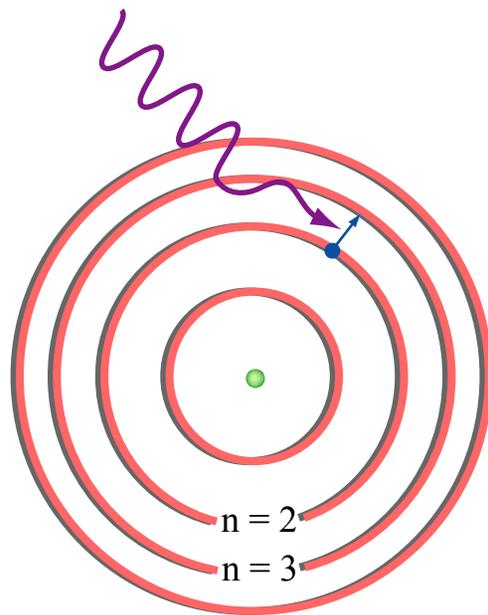
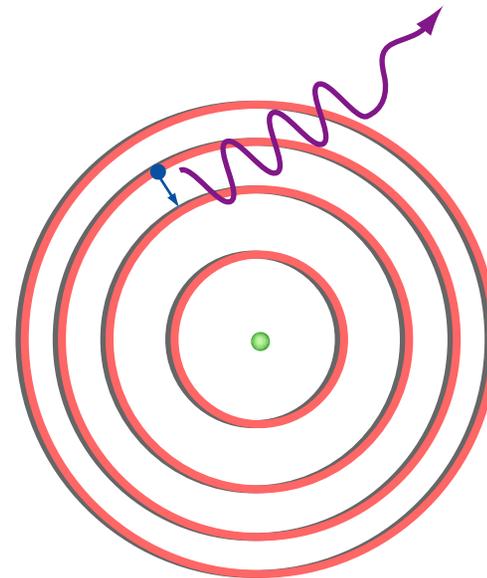


Figure by MIT OCW.



Ⓐ Absorption



Ⓑ Emission

Figure by MIT OCW.

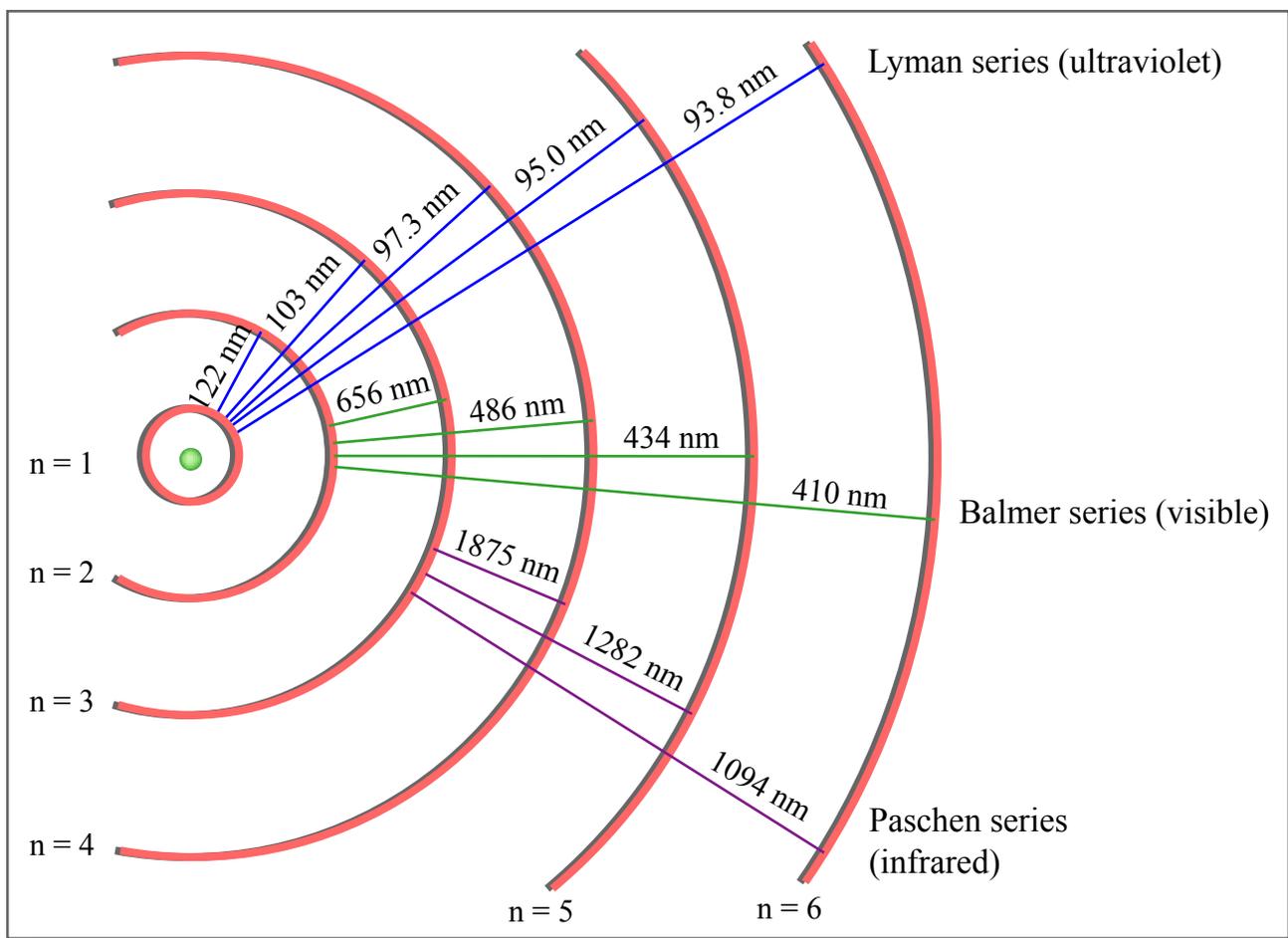
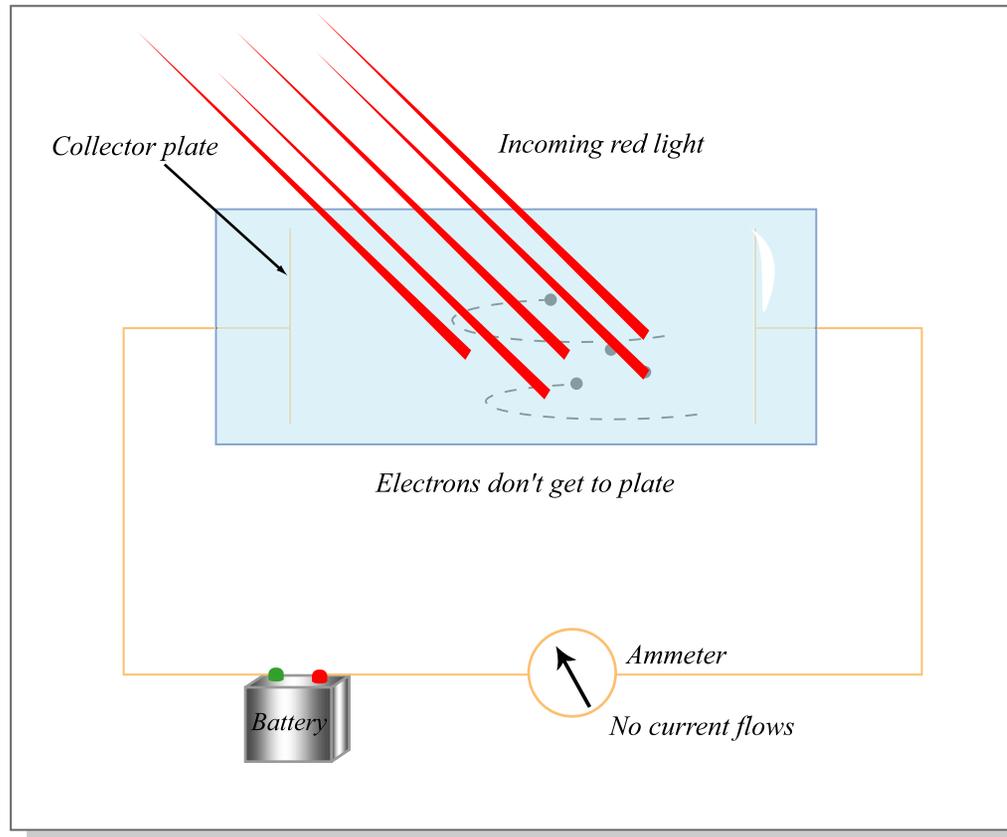


Figure by MIT OCW.

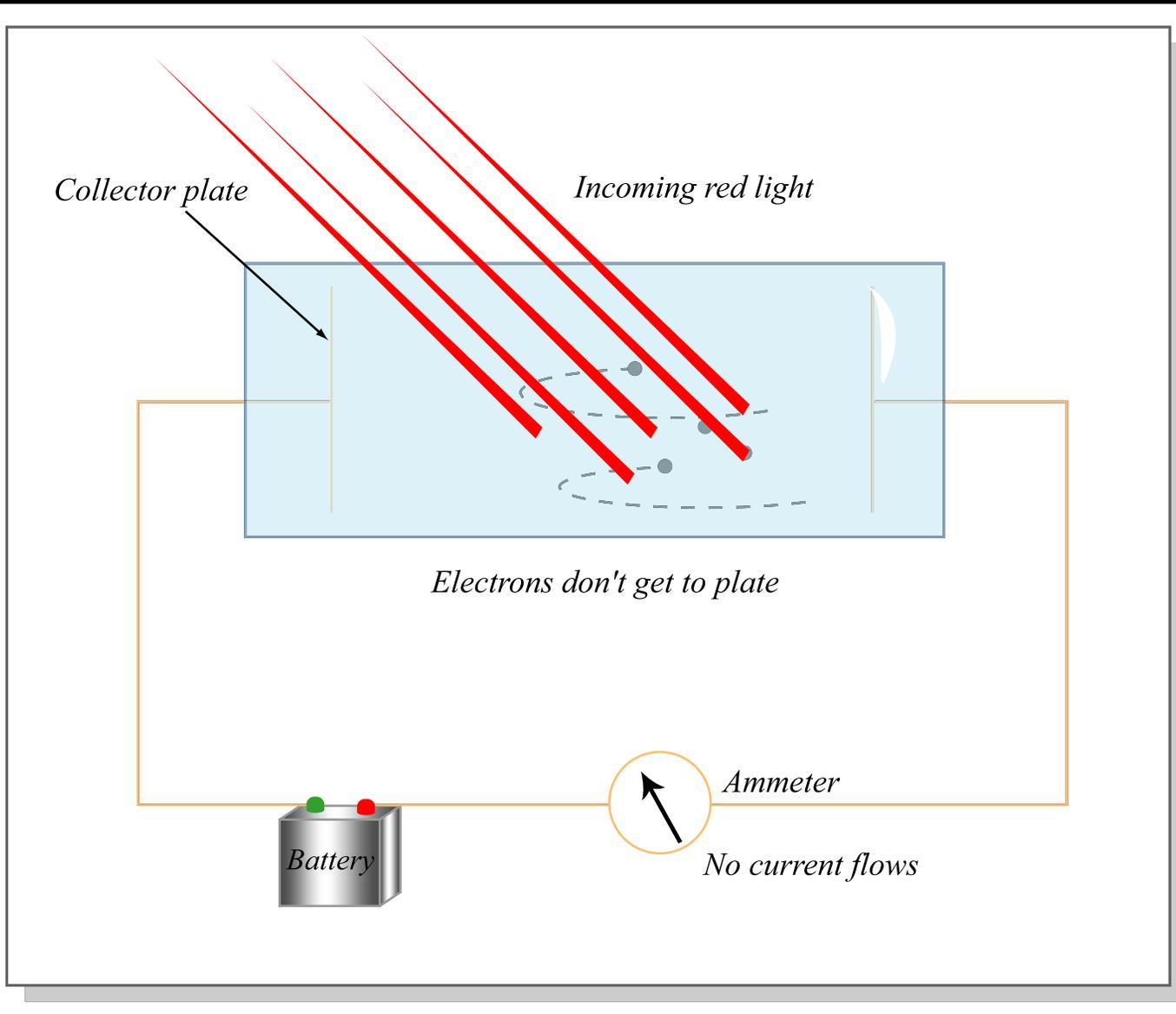
Photons

Photoelectric effect

- Einstein's model was that
 - the photon carries energy $h\nu$
 - a certain work W_e is required to liberate an electron from the metal



Photoelectric effect



Photoelectric effect

Lenard 1902

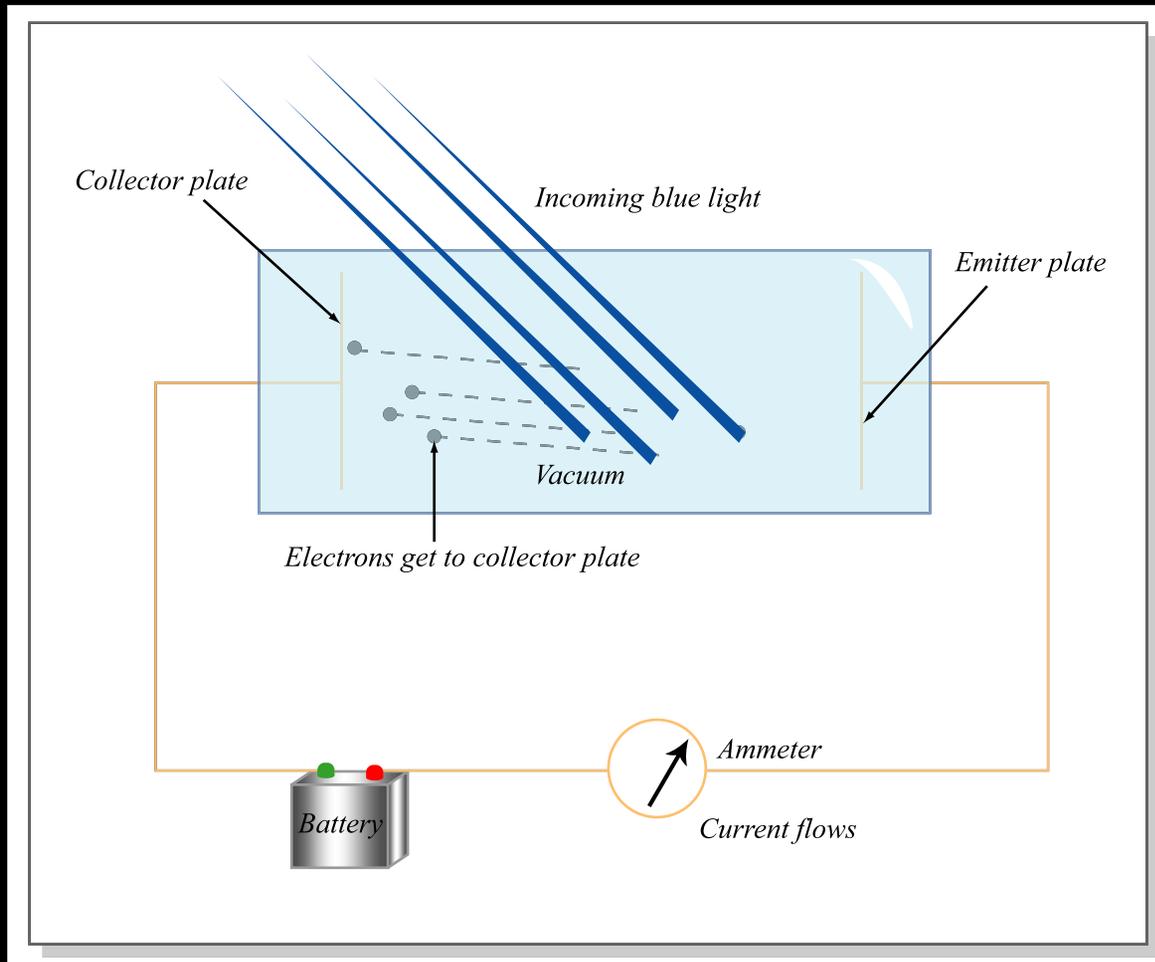


Figure by MIT OCW.

Photoelectric effect

- Einstein's model was that
 - the photon carries energy $h\nu$
 - a certain work W_e is required to liberate an electron from the metal
- This explained both of Lenard's 1902 observations:
 - light with frequency $h\nu < W_e$ liberates no electrons at all
 - light with frequency $h\nu > W_e$ liberates electrons with kinetic energy $h\nu - W_e$.
 - increasing the *intensity* of the light (the photon flux) didn't affect the existence of liberated electrons or their kinetic energy
- Bottom line: we can treat the photon as just another particle.

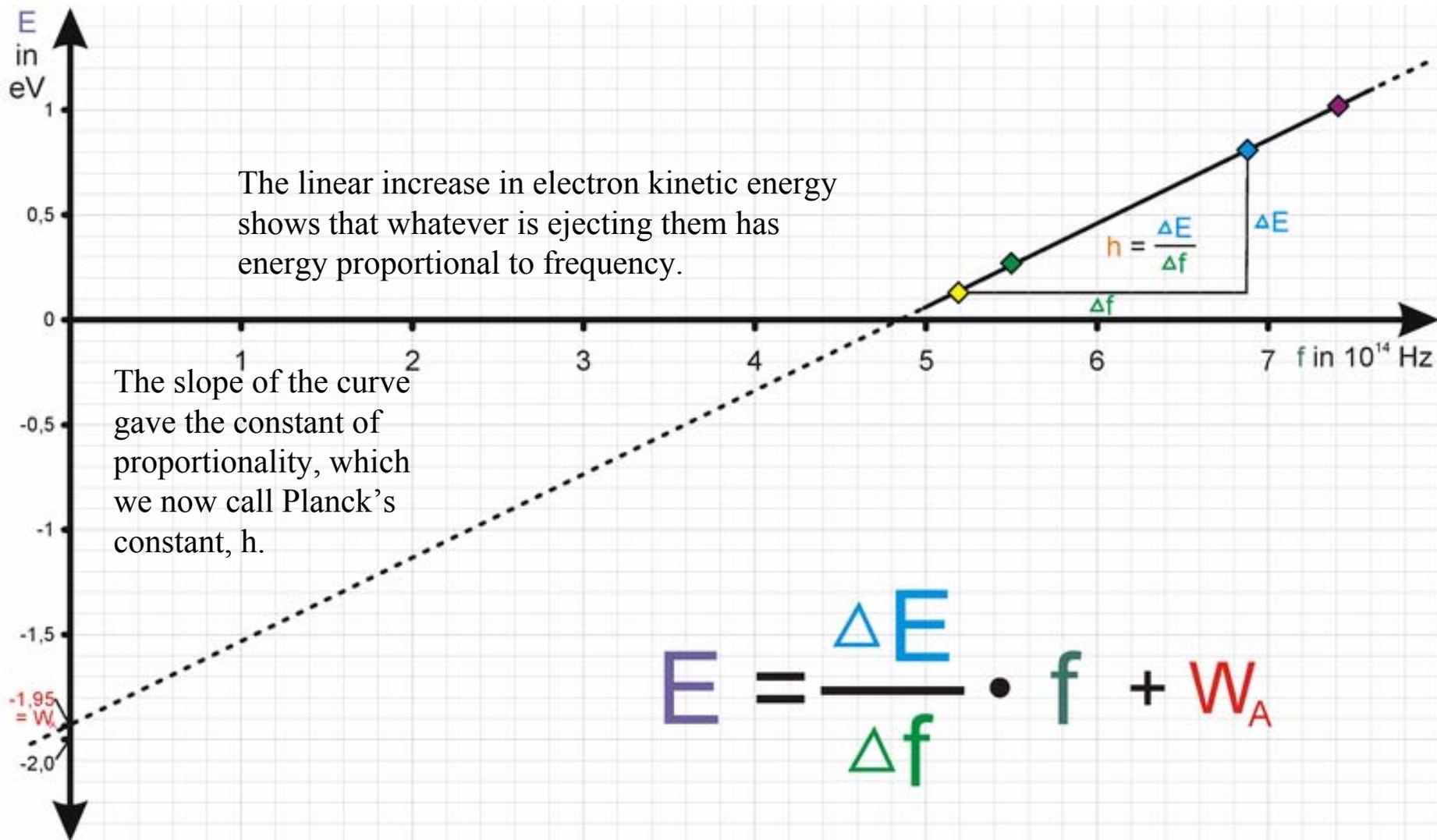


Image courtesy of Wikipedia.

(PS6, problem 6)

Working with photons: (PS6, problem 4)

- Photon 4-vector:

$$\mathbf{P} = \hbar \begin{pmatrix} \mathbf{k} \\ k \end{pmatrix},$$

where $k = \omega/c$.

- So $p = E/c$ for photons.
- Comparing \mathbf{P} with the wave 4-vector \mathbf{K} shows that

$$\mathbf{P} = \hbar \mathbf{K}.$$

This relation in fact holds for *all* particles, even massive ones — as you'll see when you get to wave-particle duality in quantum mechanics. If you take a field theory course, you'll see this pop right out of the so-called Klein-Gordon equation.

- Doppler effect is just special case of \mathbf{P} -transformation for zero rest mass — show on PS6.

Nuclear physics

Nuclear physics terminology

- The *atomic number* Z of a nucleus is its number of protons.
- The *atomic weight* A of a nucleus is its number of nucleons (protons + neutrons).
- Z determines the name of the element (its order in the periodic table).
- Nuclei with same Z and different A are said to be different *isotopes* of the same element.
- Notation example: Fe^{56} means $Z = 26$ (iron) and $A = 56$.
- The *mass excess* for a nucleus is $m_0 - A$ amu, *i.e.*, its rest mass minus the number of nucleons times amu.
- By this definition, the mass excess of C^{12} is zero.

(PS6, problem 1)

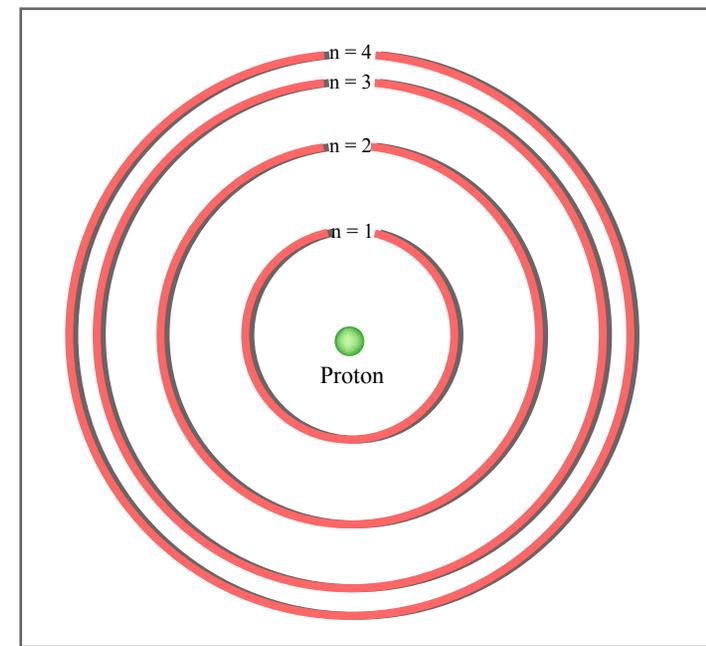


Figure by MIT OCW.

Rest mass and binding energy

- The rest energy of an object is its energy in the frame where it has zero momentum.
- This rest energy is the sum of *all* energy contributions, both positive (like rest masses and kinetic energies of its constituent particles) and negative (like potential energy from force holding constituents together).
- The *binding energy* of a nucleus is rest energy of its neutrons and protons free minus rest energy of the nucleus.

- $m_n \approx 939.6 \text{ MeV}$ (udd)
- $m_p \approx 938.3 \text{ MeV}$ (uud)
- $m_u \approx 4 \text{ MeV}$
- $m_d \approx 7 \text{ MeV}$
- Electromagnetic $\approx 1.7 \text{ MeV}$
- So mostly glue!

- Electric repulsion between protons *increases* mass of nucleus.
- Attraction between nucleons (strong force) *decreases* mass of nucleus.
- Only nuclei whose (Z, A) give positive binding energy can exist
- Semi-empirical relationship (von Weizsäcker 1935):

$$\frac{E_{\text{binding}}}{c^2} \approx \left[15.8A - 18.3A^{2/3} - 0.714\frac{Z^2}{A^{1/3}} - 23.2\frac{(A - 2Z)^2}{A} + (-1)^Z \frac{12}{A^{1/2}} \right] \text{ MeV}$$

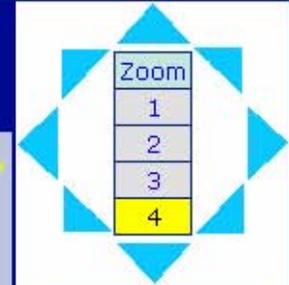
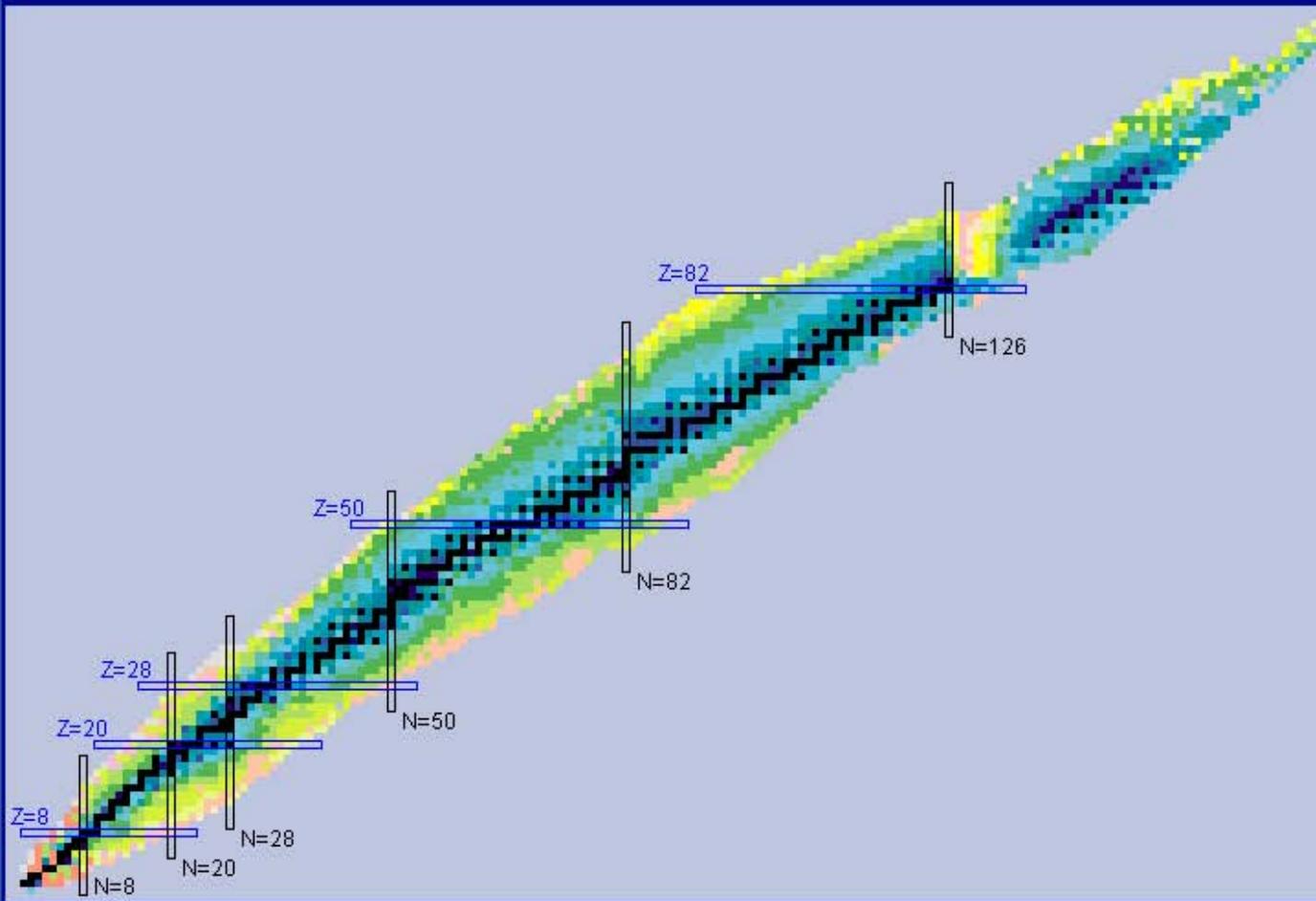
The last term is omitted if A is an odd number.

- Much work remains to be done in this field!



Chart of Nuclides

Click on a nucleus for information



Find a Nucleus:

Color code	Tooltips
Half-life	On
Decay Mode	Off

> 10+15 s	10-01 s
10+10 s	10-02 s
10+07 s	10-03 s
10+05 s	10-04 s
10+04 s	10-05 s
10+03 s	10-06 s
10+02 s	10-07 s
10+01 s	10-15 s
10+00 s	< 10-15 s
unknown	

Search options:

Levels and Gammas
Nuclear Wallet Cards
Decay Radiation

[Help - Glossary](#)

Image courtesy of the National Nuclear Data Center.

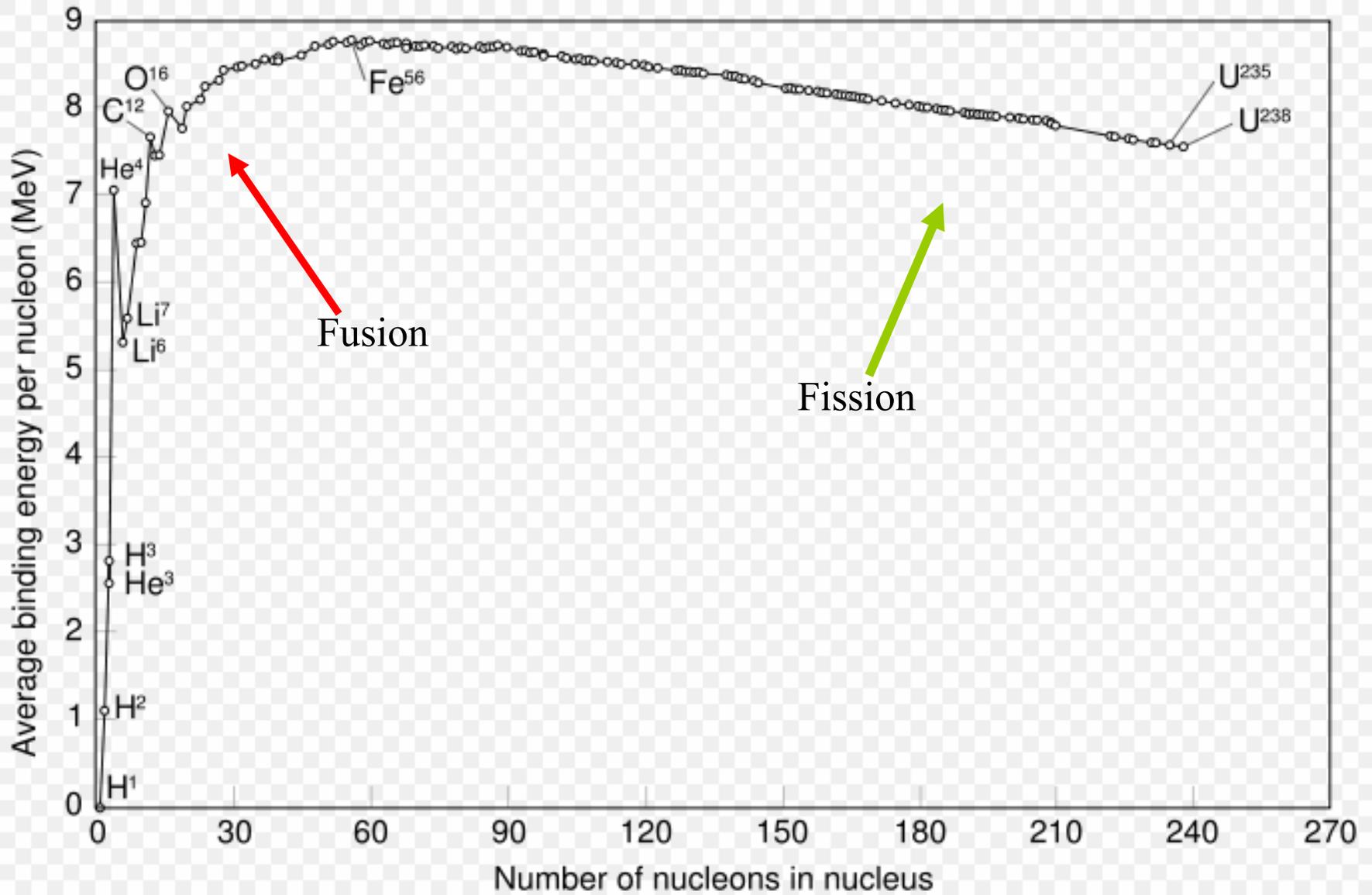


Image courtesy of Wikipedia.

