

Quiz 2

Name: (Last, First) _____ (please print).

Recitation number (circle one): 1 2 3

- Record all answers and show all work in this exam booklet. If you need extra space, use the back of the page.
- All scratch paper must be handed in with the exam, but will not be graded.
- This exam is closed book. You may use your handwritten notes if they are clearly labeled with your name and you hand them in with your exam.
- Whenever possible, try to solve problems using general analytic expressions. Plug in numbers only as a last step.
- Please make sure to answer all sub-questions.
- Good Luck!

Problem	Max	Grade	Grader
1	25		
2	25		
3	25		
4	25		
Total	100		

- (a). (5 pts) Consider the following reactions:
- (A) _____ A 100 MeV photon decays into an electron-positron pair. *answer:E*
 - (B) _____ A neutron decays into an electron-positron pair and a photon. *answer:B*
 - (C) _____ A neutron decays into a proton, an electron and a neutrino. *answer:L*
 - (D) _____ A proton decays into a neutron, a positron and a neutrino. *answer:E*
 - (E) _____ A neutron decays into a proton and a photon. *answer:Q*
- For each one, write **one** of the letters from the option list below.
- L violates lepton number conservation
 - B violates baryon number conservation
 - P violates parity conservation
 - E violates energy-momentum conservation
 - Q violates charge conservation
 - N violates none of the above conservation laws
- (b). (9 pts) Give each of the following quantities to the nearest power of 10 (don't show calculations, being off by one power of 10 is OK):
- (A) _____ Age of our universe when most He nuclei were formed *answer:1 min*
 - (B) _____ Age of our universe when hydrogen atoms formed. *answer:400000 yrs*
 - (C) _____ Age of our universe today. *answer:10 Gyr*
 - (D) _____ Number of stars in our Galaxy. *answer:1e11*
 - (E) _____ Light travel time to closest star (Sun!:) in minutes. *answer:8*
 - (F) _____ Hydrogen binding energy in eV/c^2 . *answer:13.6*
 - (G) _____ Electron mass in eV/c^2 . *answer:500000*
 - (H) _____ Neutron mass in eV/c^2 . *answer:10⁹*
 - (I) _____ Light travel time to 2nd closest star in years. *answer:3*
- (c). (9 pts) Indicate whether each of the following statements are true or false.
- (A) TRUE / FALSE If our Universe is only X billion years old, then we can only see objects that are now less than X billion light years away *answer:F*
 - (B) TRUE / FALSE Space must be infinite, because it cannot end with a boundary without more space on the other side. *answer:F*
 - (C) TRUE / FALSE Leptons do not feel the strong interaction. *answer:T*
 - (D) TRUE / FALSE No experiment inside an isolated sealed lab in space can distinguish between whether it is uniformly accelerating or in a uniform gravitational field. *answer:T*
 - (E) TRUE / FALSE A clock by the ceiling runs faster than one by the floor. *answer:T*
 - (F) TRUE / FALSE Hubble's law implies that the Big Bang was an explosion localized near the comoving position of our Galaxy.
 - (G) TRUE / FALSE The expansion of our galaxy is governed by the Friedmann equation. *answer:F*
 - (H) TRUE / FALSE Two galaxies can recede from each other faster than the speed of light. *answer:T*
 - (I) TRUE / FALSE We know that our entire observable universe was once at infinite density *answer:F*
- (d). (2 pts) A tritium (H^3) nucleus contains _____ up quarks and _____ down quarks. *answer:p+2n = 4 + 5*

In the Sun, one of the processes in the He fusion chain is $p + p + e^- \rightarrow d + \nu$, where d is a deuteron. Make the approximations that the deuteron rest mass is $2m_p$, and that $m_e \approx 0$ and $m_\nu \approx 0$, since both the electron and the neutrino have negligible rest mass compared with the proton rest mass m_p .

- (a). For the arrangement shown in the figure, where (in the lab frame) the two protons have the same energy γm_p and impact angle θ , and the electron is at rest, calculate the energy E_ν of the neutrino in the rest frame of the deuteron in terms of θ , m_p and γ .

answer: Use the fact that the quantity $E^2 - p^2c^2$ is invariant. In the deuteron's rest frame, after the collision:

$$E^2 - p^2c^2 = (2m_p c^2 + E_\nu)^2 - E_\nu^2 \quad (2.1)$$

$$= 4m_p^2 c^4 + 4m_p c^2 E_\nu = 4m_p c^2 (m_p c^2 + E_\nu) \quad (2.2)$$

In the lab frame, before collision:

$$E^2 - p^2c^2 = (2E_p)^2 - (2p_p \cos \theta c)^2 \quad (2.3)$$

$$= (2\gamma m_p c^2)^2 - (2\gamma \beta m_p c \cos \theta)^2 \quad (2.4)$$

Use $\gamma^2 \beta^2 = (\gamma^2 - 1)$ in the second term and simplify the algebra to find

$$E^2 - p^2c^2 = 4m_p^2 c^4 (\gamma^2 - (\gamma^2 - 1) \cos^2 \theta) \quad (2.5)$$

Equating the invariants in the two frames, we have

$$4m_p c^2 (m_p c^2 + E_\nu) = 4m_p^2 c^4 (\gamma^2 - (\gamma^2 - 1) \cos^2 \theta) \quad (2.6)$$

$$\Rightarrow E_\nu = m_p c^2 (\gamma^2 - 1) \sin^2 \theta \quad (2.7)$$

- (b). For the special case where the deuteron remains at rest in the lab frame and $\theta = 30^\circ$, solve for γ and calculate the energy of all particles (the deuteron, the neutrino, one of the protons) in terms of the proton rest mass m_p .

answer: The deuteron's rest frame is the lab frame. Also, $\theta = 30^\circ$. Use conservation of energy, along with the result from the previous part to find:

$$2\nu m_p c^2 = 2m_p c^2 + E_\nu \quad (2.8)$$

$$= 2m_p c^2 + m_p c^2 (\gamma^2 - 1)/4 \quad (2.9)$$

$$\Rightarrow 2\gamma = 2 + \gamma^2/4 - 1/4 \quad (2.10)$$

$$\Rightarrow \gamma = 7, 1 \quad (2.11)$$

$\gamma = 1$ is obviously not the solution. Thus, $\gamma = 7$ and the energies are: $E_p = 7m_p$, $E_\nu = 12m_p$, $E_d = 2m_p$

Question 3: Coulomb's Law generalized

[25 Points]

In an inertial frame S , the position \mathbf{r}_q of a point charge q moves according to $\mathbf{r}_q(t) = v\hat{\mathbf{z}}t$, i.e. with velocity v in the $\hat{\mathbf{z}}$ -direction, passing the origin at $t = 0$. In the moving frame S' where the charge is at rest at the origin, Coulomb's law states that the electric field is

$$\mathbf{E}' = A \frac{\mathbf{r}'}{r'^3},$$

where $A = q/4\pi\epsilon_0$. Show that in the frame S , the electric field at $t = 0$ is

$$\mathbf{E} = A \frac{(1 - \beta^2)}{(1 - \beta^2 \sin^2 \theta)^{3/2}} \frac{\mathbf{r}}{r^3},$$

where θ is the usual polar angle ($z = r \cos \theta$, $x^2 + y^2 = r^2 \sin^2 \theta$).

answer: Let us convert all quantities to the cartesian coordinates. In the frame S' , the components of the electric field are:

$$\vec{E}' = E'_x \hat{x}' + E'_y \hat{y}' + E'_z \hat{z}' \quad (3.1)$$

$$E'_x = \frac{Ax'}{(x'^2 + y'^2 + z'^2)^{3/2}} \quad (3.2)$$

$$E'_y = \frac{Ay'}{(x'^2 + y'^2 + z'^2)^{3/2}} \quad (3.3)$$

$$E'_z = \frac{Az'}{(x'^2 + y'^2 + z'^2)^{3/2}} \quad (3.4)$$

We can now Lorentz transform the fields and coordinates from S' to S . First the coordinates,

$$x = x' \quad (3.5)$$

$$y = y' \quad (3.6)$$

$$\gamma z = z' \quad (3.7)$$

and then the fields,

$$E_x = \gamma E'_x = \frac{A\gamma x}{(x^2 + y^2 + \gamma^2 z^2)^{3/2}} \quad (3.8)$$

$$E_y = \gamma E'_y = \frac{A\gamma y}{(x^2 + y^2 + \gamma^2 z^2)^{3/2}} \quad (3.9)$$

$$E_z = E'_z = \frac{A\gamma z}{(x^2 + y^2 + \gamma^2 z^2)^{3/2}} \quad (3.10)$$

Note that the primed coordinates have been converted to the unprimed ones using the coordinate transformation. The total magnitude for the electric field in the S frame can be obtained from

$$E^2 = E_x^2 + E_y^2 + E_z^2 \quad (3.11)$$

$$= \frac{A^2 \gamma^2 r^2}{(x^2 + y^2 + \gamma^2 z^2)^3} \quad (3.12)$$

$$= \frac{A^2 r^2}{(1 - \beta^2)(x^2 + y^2 + z^2/(1 - \beta^2))^3} \quad (3.13)$$

$$\Rightarrow E = \frac{A(1 - \beta^2)}{r^2(1 - \beta^2 \sin^2 \theta)^{3/2}} \quad (3.14)$$

Since the electric field always has to be radial, $\vec{E} = |E|\hat{r}$.

- (a). **(10 pts)** Consider a particle coasting in the r -direction (i.e., with constant θ and ϕ) in a flat FRW metric, with no non-gravitational forces acting on it. Use variational calculus to prove that $p \propto 1/a$ (here $p = m_0\gamma u$ is its momentum and $u \equiv a\dot{r}$ is its velocity relative to nearby comoving observers). *answer:*The particle only has a radial motion $\Rightarrow d\theta = d\phi = 0$. Also, the universe is flat $\Rightarrow k = 0$. Thus, the FRW metric becomes:

$$d\tau^2 = dt^2 - a^2 dr^2 \tag{4.1}$$

$$\Rightarrow \Delta\tau = \int d\tau = \int \sqrt{dt^2 - a^2 dr^2} \tag{4.2}$$

$$= \int dt \sqrt{1 - a^2 \dot{r}^2} = \int dt f(t, r, \dot{r}) \tag{4.3}$$

The interval $\Delta\tau$ has to be extremized. The Euler lagrange equations give:

$$\frac{\partial f}{\partial r} - \frac{d}{dt} \left[\frac{\partial f}{\partial \dot{r}} \right] = 0 \tag{4.4}$$

$$\Rightarrow \frac{d}{dt} \left[\frac{a^2 \dot{r}}{\sqrt{1 - a^2 \dot{r}^2}} \right] = 0 \tag{4.5}$$

$$\Rightarrow \frac{a^2 \dot{r}}{\sqrt{1 - a^2 \dot{r}^2}} = \text{constant} \tag{4.6}$$

Identifying $a\dot{r} = u$ leave us with $\gamma u \propto 1/a \Rightarrow p = m_0\gamma u \propto 1/a$.

- (b). **(2 pts)** Given a), the value of u in the limit $a \rightarrow 0$ is _____. *answer:*1(c)
 (c). **(2 pts)** Given a), the value of u in the limit $a \rightarrow \infty$ is _____. *answer:*0
 (d). **(2 pts)** Thus relative to comoving observers, your results show that an object without external forces in an expanding universe **(circle one)** REMAINS IN UNIFORM MOTION / SLOWS DOWN / ACCELERATES. *answer:*slows down
 (e). **(3 pts)** Starting with the answer from a), derive how the wavelength λ of a photon depends on a . Your answer should be of the form $\lambda \propto$ (function of a). *answer:*For a photon, $p = hk/2\pi \propto 1/\lambda$. So $\lambda \propto 1/p \propto a$.
 (f). **(6 pts)** Solve the Friedmann equation

$$H^2 = \frac{8\pi G}{3} \rho - \frac{kc^2}{a^2}$$

to obtain a solution of the form $a(t) \propto$ (function of t) for the case where space is flat and the density is dominated by photons, and compute the age of the universe at the time when $H^{-1} = 30$ seconds.

*answer:*For a flat universe dominated by photons,

$$\left(\frac{\dot{a}}{a} \right)^2 = H^2 \propto a^{-4} \tag{4.7}$$

$$\Rightarrow \dot{a} \propto a^{-1} \tag{4.8}$$

$$\Rightarrow a = At^{1/2} \tag{4.9}$$

$$\Rightarrow H = \dot{a}/a = 1/2t \tag{4.10}$$

Thus, $t = H^{-1}/2 = 15\text{seconds}$.