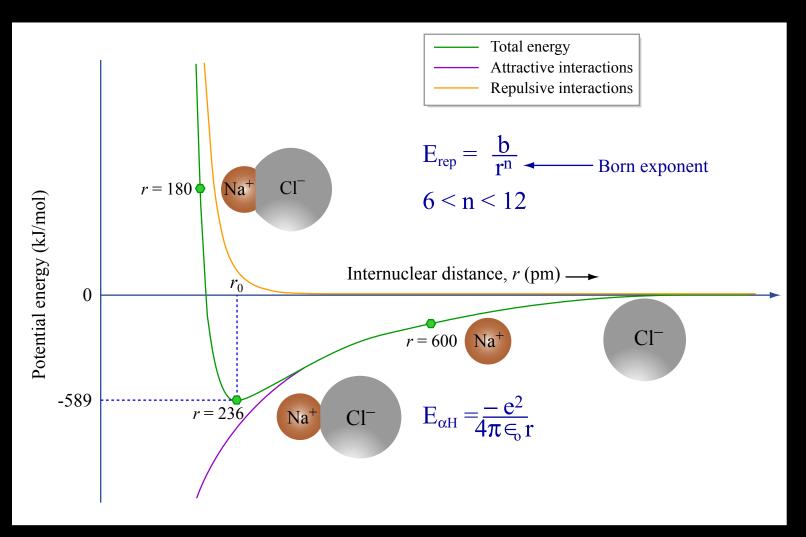
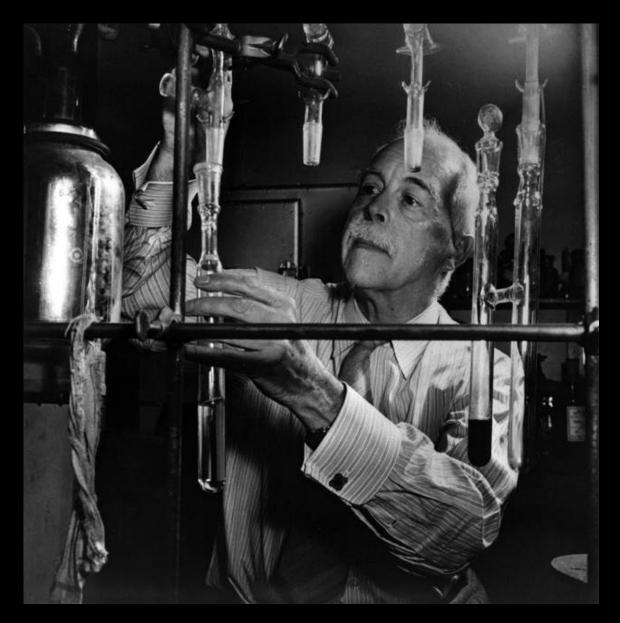
Welcome to 3.091

Lecture 9

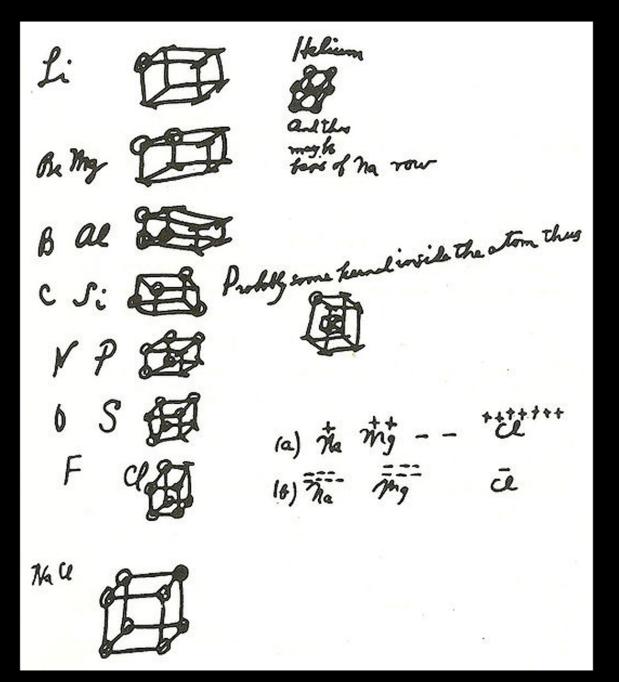
September 28, 2009

Drawing Lewis Structures





Public domain image from Wikipedia.



Public domain image from Wikipedia.

Lewis Notation: Electron - Symbols (1916)

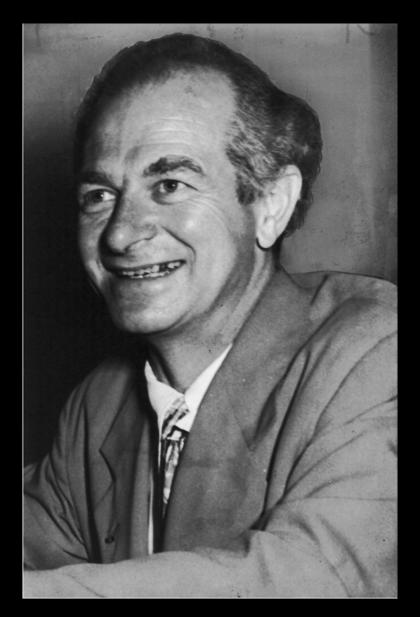
		1A(1)	2A(2)				
		ns ¹	ns ²				
Period	2	• Li	• Be •				
Pel	3	• Na	• Mg•				

3A(13)	4A(14)	5A(15)	6A(16)	7A(17)	8A(18)
ns ² np ¹	ns ² np ²	ns ² np ³	ns ² np ⁴	ns ² np ⁵	ns ² np ⁶
• B •	• C •	• • • • • • • • • • • • • • • • • • •	• • •	• • F • • • • • • • • • • • • • • • • •	• • • • • • • • • • • • • • • • • • •
• Al •	• Si •	• • • • • • • • • • • • • • • • • • •	• • • • • • • • • • • • • • • • • • •	: C1 :	• • Ar • • • • • • • • • • • • • • • • •

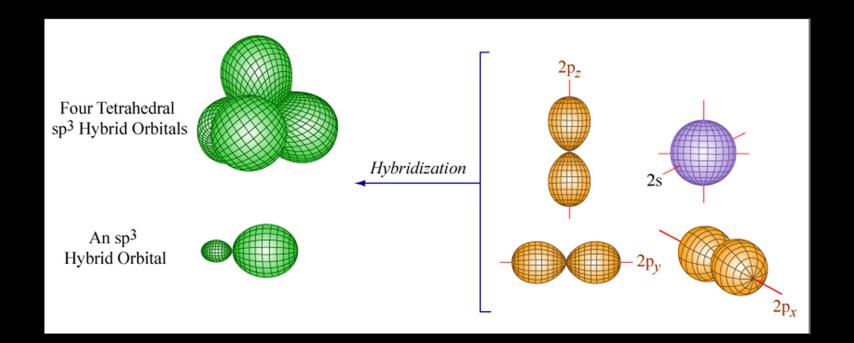
- Element symbol \equiv Nucleus & inner $e\bar{s}$
- $-Dots \equiv Valence \ e\bar{s}$

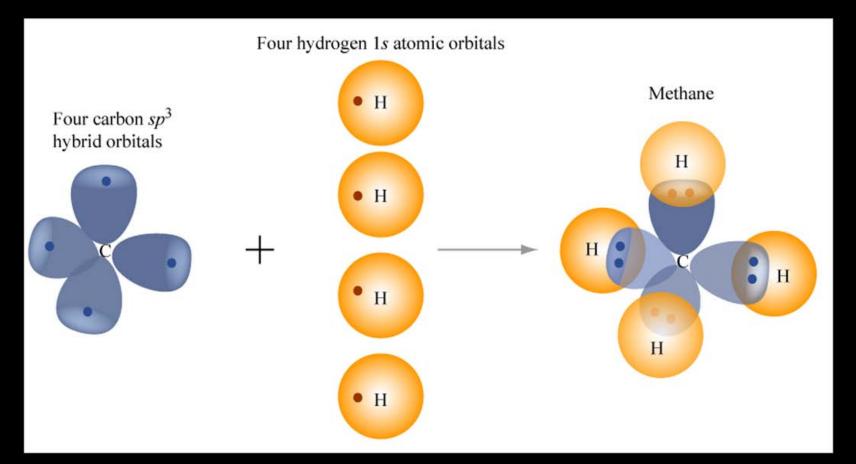
Drawing Lewis Structures

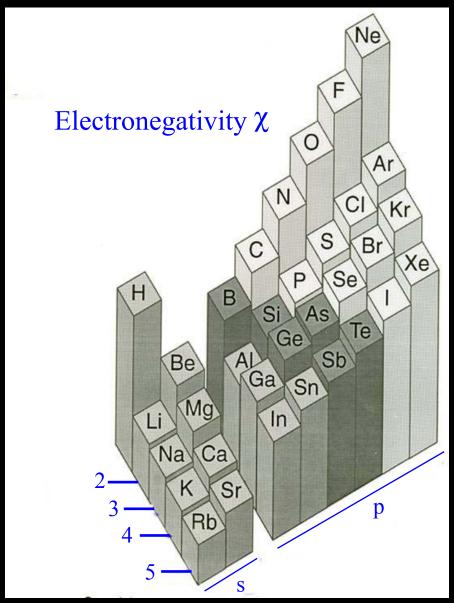
- 1. center the element with lowest AVEE
- 2. count all valence electrons
- 3. draw a single bond from each surrounding atom to the central atom; subtract 2 valence e s for each bond
- 4. distribute remaining e⁻s in pairs until each atom has 8 e⁻s in total; place NB pairs on peripheral atoms starting with highest AVEE



Public domain image from Wikipedia.





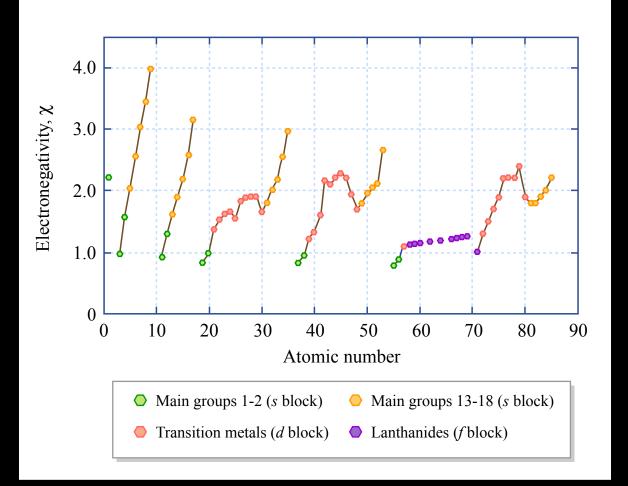


Nonmetals have high χ

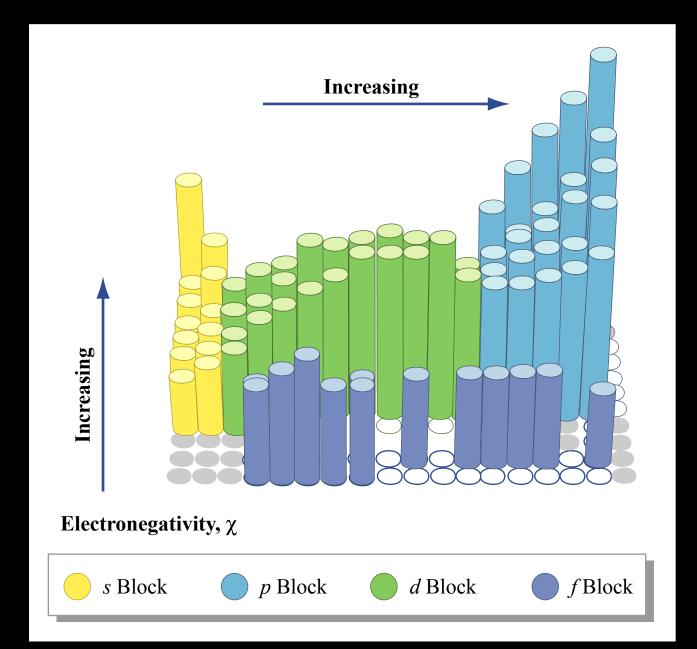
Metal have low χ

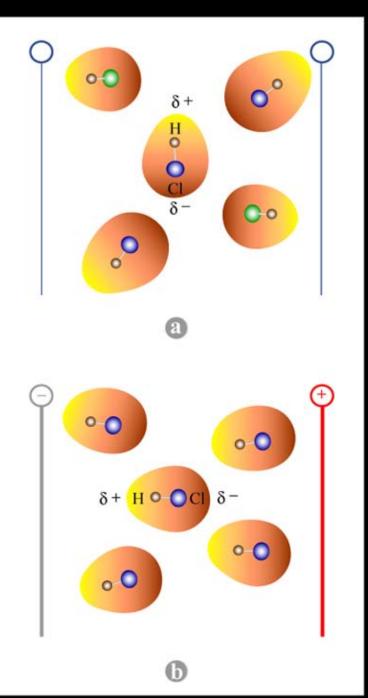
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	1																	18
1	Н																	Не
•	2.20	2											13	14	15	16	17	
2	Li	Be			0.7					3.98			В	C	N	O	F	Ne
۷	0.98	1.57			0.7					3.96			2.04	2.55	3.04	3.44	3.98	
3	Na	Mg											Al	Si	P	S	Cl	Ar
3	0.93	1.31	3	4	5	6	7	8	9	10	11	12	1.61	1.90	2.19	2.58	3.16	
4	K	Ca	Sc	Ti	V	Cr	Mn	Fe	Co	Ni	Cu	Zn	Ga	Ge	As	Se	Br	Kr
	0.82	1.00	1.36	1.54	1.63	1.66	1.55	1.83	1.88	1.91	1.90	1.65	1.81	2.01	2.18	2.55	2.96	
5	Rb	Sr	Y	Zr	Nb	Mo	Tc	Ru	Rh	Pd	Ag	Cd	In	Sn	Sb	Te	I	Xe
	0.82	0.95	1.22	1.33	1.6	2.16	2.10	2.2	2.28	2.20	1.93	1.69	1.78	1.96	2.05	2.1	2.66	
6	Cs	Ba	La	Hf	Ta	W	Re	Os	Ir	Pt	Au	Hg	T1	Pb	Bi	Po	At	Rn
	0.79	0.89	1.10	1.3	1.5	1.7	1.9	2.2	2.2	2.2	2.4	1.9	1.8	1.8	1.9	2.0	2.2	
7	Fr	Ra	Ac	Rf	Db	Sg	Bh	Hs	Mt	Ds	Rg	Uub	Uut	Uuq	Uup			
	0.7	0.9	1.1															
		Lantha		Ce	Pr	Nd	Pm	Sm	Eu	Gd	Tb	Dy	Но	Er	Tm	Yb	Lu	
			6	1.12	1.13	1.14		1.17		1.20		1.22	1.23	1.24	1.25		1.0	
		Acti	nides	Th	Pa	U	Np	Pu	Am	Cm	Bk	Cf	Es	Fm	Md	No	Lr	
			7	1.3	1.5	1.7	1.3	1.3										

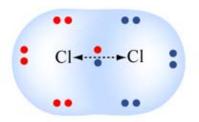




Nonpolar covalent bond

Bonding electrons shared *equally* between two atoms.

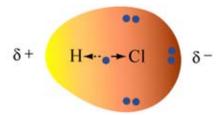
No charges on atoms.



Polar covalent bond

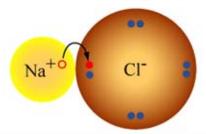
Bonding electrons shared *unequally* between two atoms.

Partial charges on atoms.



Ionic bond

Complete transfer of one or more valence electrons. Full charges on resulting ions.



% ionic character = $\left\{1 - \exp\left(-\frac{1}{4}(\Delta X)^2\right)\right\} \times 100$

▼ % Ionic Character of a Single Chemical Bond

▼ % ionic Character of a Single Chemical Bond									
Difference in Electronegativity	%IC (by L. Pauling)	%IC (by Hannay & Smyth)							
0.1	0.2	1.6							
0.2	1.0	3.3							
0.3	2.2	5.1							
0.4	3.9	7.0							
0.5	6.1	8.9							
0.6	8.6	11							
0.7	12	13							
0.8	15	15							
0.9	18	17							
1.0	22	20							
1.1	26	22							
1.2	30	24							
1.3	34	27							
1.4	39	29							
1.5	43	32							
1.6	47	35							
1.7	51	37							
1.8	56	40							
1.9	59	43							
2.0	63	46							
2.1	67	49							
2.2	70	52							
2.3	73	55							
2.4	76	59							
2.5	79	62							
2.6	82	65							
2.7	84	69							
2.8	86	72							
2.9	88	76							
3.0	89	80							
3.1	91	83							
3.2	92	87							

Thomas "sp3", Midgley

Freon 12: designer molecule, tailored chemical

- early refrigerants were toxic or flammable,
 e.g., ammonia, methyl chloride, sulfur dioxide
- in late 1920s Midgley discovered CCl₂F₂ with properties of a refrigerant and propellant: perfect!



- but in the upper atmosphere, u.v. light breaks the C – Cl bond, and atomic Cl attacks ozone

 $Cl + O_3 \Rightarrow ClO + O_2$

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but in the upper atmosphere, u.v. light breaks
 the C – Cl bond, and atomic Cl attacks ozone

$$C1 + O_3 \Rightarrow C1O + O_2$$

CFCs => ozone dipletion

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Figure removed due to copyright restrictions. Chandler, David L. "MIT scientist shares Nobel for identifying ozone damage." *The Boston Globe*, October 12, 1995.

193. Acia, 292, 236-270 (1973).

Stratospheric sink for chlorofluoromethanes: chlorine atomc-atalysed destruction of ozone

Mario J. Molina & F. S. Rowland

Department of Chemistry, University of California, Irvine, California 92664

Nature, **249** June 28, 1974

Chlorofluoromethanes are being added to the environment in steadily increasing amounts. These compounds are chemically inert and may remain in the atmosphere for 40-150 years, and concentrations can be expected to reach 10 to 30 times present levels. Photodissociation of the chlorofluoromethanes in the stratosphere produces significant amounts of chlorine atoms, and leads to the destruction of atmospheric ozone.

HALOGENATED aliphatic hydrocarbons have been added to the natural environment in steadily increasing amounts over several decades as a consequence of their growing use, chiefly as aerosol propellants and as refrigerants1.2. Two chlorofluoromethanes, CF2Cl2 and CFCl3, have been detected throughout the troposphere in amounts (about 10 and 6 parts per 1011 by volume, respectively) roughly corresponding to the integrated world industrial production to date3-5,31. The chemical inertness and high volatility which make these materials suitable for technological use also mean that they remain in the atmosphere for a long time. There are no obvious rapid sinks for their removal, and they may be useful as inert tracers of atmospheric motions4-6. We have attempted to calculate the probable sinks and lifetimes for these molecules. The most important sink for atmospheric CFCl3 and CF2Cl2 seems to be stratospheric

photolytic dissociation to CFCl2 + Cl and to CF2Cl + Cl, respectively, at altitudes of 20-40 km. Each of the reactions creates two odd-electron species-one Cl atom and one free radical. The dissociated chlorofluoromethanes can be traced to their ultimate sinks. An extensive catalytic chain reaction leading to the net destruction of O3 and O occurs in the stratosphere:

$$Cl + O_3 \rightarrow ClO + O_2$$

 $ClO + O \rightarrow Cl + O_2$ (1)

This has important chemical consequences. Under most conditions in the Earth's atmospheric ozone layer, (2) is the slower of the reactions because there is a much lower concentration of O than of O3. The odd chlorine chain (CI, CIO) can be compared with the odd nitrogen chain (NO, NO2) which is believed to be intimately involved in the regulation of the present level of O3 in the atmosphere7-10. At stratospheric temperatures, ClO reacts with O six times faster than NO2 reacts with O (refs 11, 12). Consequently, the Cl-ClO chain can be considerably more efficient than the NO-NO2 chain in the catalytic conversion of O3 + O-2O2 per unit time per reacting chain13.

Photolytic sink

Both CFCl3 and CF2Cl2 absorb radiation in the far ultraviolet14, and stratospheric photolysis will occur mainly in the 'window' at 1,750-2,200 Å between the more intense absorptions of the Reprinted by permission from Macmillan Publishers Ltd: *Nature* © 1974. Source: Molina, Mario J., and Rowland, F. S. "Stratospheric sink for

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3.091SC Introduction to Solid State Chemistry Fall 2009

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