APERTURE ANTENNAS

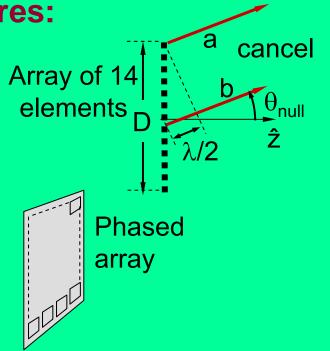
Dense phased arrays ~apertures:

Pairs of radiators cancel (e.g., a,b) Sum of pairs therefore = 0 First null at $\theta = \sim \lambda/D$

Aperture antennas:







Derivation of far-field aperture radiation:

Find surface current \overline{J}_s that could produce the same aperture fields Find the far fields radiated by an infinitesimal element of \overline{J}_s Integrate the contributions from the entire aperture Simplify the expressions by using small-angle approximations Note the Fourier relationship between aperture and far fields

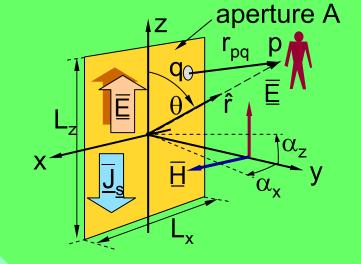
APERTURE ANTENNAS

Assume UPW in aperture:

$$\underline{E}_{+} = \hat{z} \, \underline{E}_{o} e^{-jky} \implies \underline{H}_{+} = \hat{x} \, \frac{\underline{E}_{o}}{\eta_{o}} e^{-jky}$$

$$\underline{\underline{E}}_{-} = \hat{z} \, \underline{\underline{E}}_{o} e^{+jky} \Rightarrow \underline{\underline{H}}_{-} = -\hat{x} \, \frac{\underline{\underline{E}}_{o}}{\eta_{o}} e^{+jky}$$

Satisfies $E_{//}$ continuous at y = 0



Equivalent surface current:

$$\underline{\overline{J}}_{s} = \hat{y} \times [\underline{\overline{H}}(y=0_{+}) - \underline{\overline{H}}(y=0_{-})]$$

$$= \hat{y} \times \hat{x}[\underline{\underline{E}}_{o}/\eta_{o} + \underline{\underline{E}}_{o}/\eta_{o}]$$

$$\overline{\underline{J}}_{s} = -\hat{z} \ 2\underline{\underline{E}}_{o}/\eta_{o} \ [A \ m^{-1}]$$

 $\hat{z} \underline{I} dz = \underline{J}_s dx dz$ (Hertzian dipole radiator)

Radiated far fields:

$$\begin{split} & \underline{\mathbb{E}}_{ff}(\theta,\phi) = \hat{\theta} \frac{j\eta_o}{2\lambda r} \sin\theta \, \iint_A \, \underline{J}_{zs}(x,z) \, e^{-jkr_{pq}} dx \, dz \\ & = -\hat{\theta} \frac{j}{\lambda r} \, \sin\theta \, \iint_A \, \underline{\mathbb{E}}_{oz}(x,z) \, e^{-jkr_{pq}} dx \, dz \end{split}$$

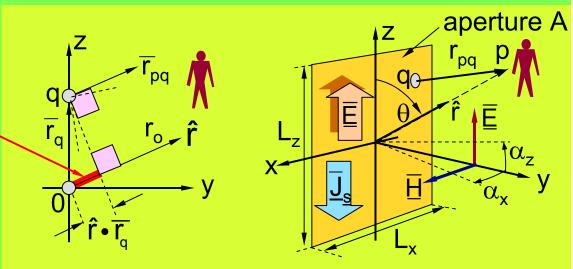
FAR-FIELD RADIATION

Assume UPW in aperture:

$$\overline{\mathbb{E}}_{ff}(\theta,\phi) = -\hat{\theta} \frac{j}{\lambda r} \sin\theta \iint_{A} \underline{\mathbb{E}}_{oz}(x,z) e^{-jkr_{pq}} dx dz$$

$$e^{-jkr} \cong e^{-jkr_0 + jk\hat{r} \cdot \overline{r}_q}$$

Fraunhofer (approximation (else, Fresnel)



$$\overline{\mathbb{E}}_{\mathrm{ff}}(\alpha_{\mathrm{X}},\alpha_{z}) \cong -\hat{\theta}\frac{\mathsf{j}}{\lambda r} \mathrm{e}^{-\mathsf{j}\mathsf{k}\mathsf{r}_{0}} \iint_{\mathsf{A}} \mathsf{E}_{\mathrm{oz}}(\mathsf{x},\mathsf{z}) \mathrm{e}^{+\mathsf{j}\frac{2\pi}{\lambda}}(\underline{\mathsf{x}}\alpha_{\mathsf{X}} + z\alpha_{z}) \, \mathrm{d}\mathsf{x} \, \mathrm{d}\mathsf{z}$$
 ~Fourier (angle \Leftrightarrow space) Close to the y axis

Fourier relationship:

$$\underline{S}(f) = \int_{-\infty}^{\infty} s(t)e^{-j2\pi ft}dt$$
 $s(t) = \int_{-\infty}^{\infty} \underline{S}(f)e^{+j2\pi ft}df$ (frequency \Leftrightarrow time)

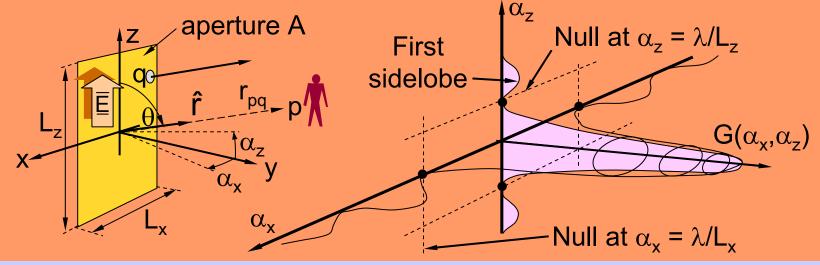
EXAMPLE: RECTANGULAR APERTURE

Assume UPW in aperture: (Observer close to the y axis)

$$\begin{split} & \underline{\overline{E}}_{ff}(\alpha_{x},\alpha_{z}) \cong -\hat{\theta} \frac{j}{\lambda r} e^{-jkr_{0}} \int_{-L_{z}/2}^{L_{z}/2} e^{+j2\pi\alpha_{z}\frac{z}{\lambda}} \int_{-L_{x}/2}^{Lx/2} \underline{E}_{oz}(x,z) e^{+j2\pi\alpha_{x}\frac{x}{\lambda}} \, dx \, dz \\ & \int_{-L_{x}/2}^{L_{x}/2} \underline{E}_{oz}(x,z) e^{+j2\pi\alpha_{x}\frac{x}{\lambda}} \, dx \, = \, \underline{E}_{oz} \frac{\lambda}{j2\pi\alpha_{x}} (e^{+j\pi\alpha_{x}\frac{L_{x}}{\lambda}} - e^{-j\pi\alpha_{x}\frac{L_{x}}{\lambda}}) = \, \underline{E}_{oz} L_{x} \frac{\sin \frac{\pi\alpha_{x}L_{x}}{\lambda}}{\frac{\pi\alpha_{x}L_{x}}{\lambda}} \\ & \triangleq \, \underline{E}_{oz} L_{x} sinc \frac{\pi\alpha_{x}L_{x}}{\lambda} \end{split}$$

$$\overline{\underline{E}}_{ff}(\alpha_{x},\alpha_{z}) \cong -\hat{\theta}\frac{j}{\lambda r}\underline{E}_{oz}e^{-jkr_{o}} L_{x}L_{z} \operatorname{sinc}\frac{\pi\alpha_{x}L_{x}}{\lambda} \operatorname{sinc}\frac{\pi\alpha_{z}L_{z}}{\lambda}$$

Antenna Gain: $G(\alpha_x, \alpha_z) \propto |\underline{E}_{ff}|^2 \propto sinc^2(\pi \alpha_x L_x/\lambda) sinc^2(\pi \alpha_z L_z/\lambda)$



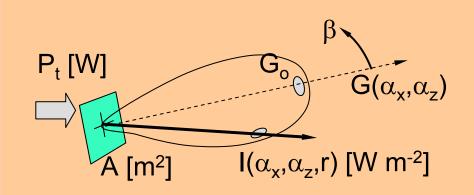
*sinc $\theta = (\sin \theta)/\theta = [1 \text{ for } \theta = 0]$

RECTANGULAR APERTURE (2)

Antenna Gain $G(\alpha_x, \alpha_z)$:

$$G(\alpha_x, \alpha_z) = \frac{I(\alpha_x, \alpha_z, r)}{(P_t / 4\pi r^2)}$$

$$P_{t} = A \frac{|\underline{E}_{oz}|^{2}}{2\eta_{o}} [W] (A=L_{x}L_{z})$$



$$I(\alpha_x, \alpha_z, r) = \frac{|\underline{\mathsf{E}}_{\mathsf{ff}}|^2}{2\eta_o} [\mathsf{W/m}^2] = \frac{1}{2\eta_o} (\frac{1}{\lambda r})^2 |\underline{\mathsf{E}}_{\mathsf{oz}}|^2 \mathsf{A}^2 \mathsf{sinc}^2 (\frac{\pi \alpha_x \mathsf{L}_x}{\lambda}) \mathsf{sinc}^2 (\frac{\pi \alpha_z \mathsf{L}_z}{\lambda})$$

$$G(\alpha_x, \alpha_z) = A(4\pi/\lambda^2) \operatorname{sinc}^2(\pi\alpha_x L_x/\lambda) \operatorname{sinc}^2(\pi\alpha_z L_z/\lambda)$$

$$G_o = A \frac{4\pi}{\lambda^2} \Rightarrow A = A_e = G_o \frac{\lambda^2}{4\pi}^* \qquad \left[\frac{\sin\theta}{\theta} \to 1 \text{ as } \theta \to 0\right]$$

 η_A = "aperture efficiency" $\equiv A_e/A \approx 0.65$ typically

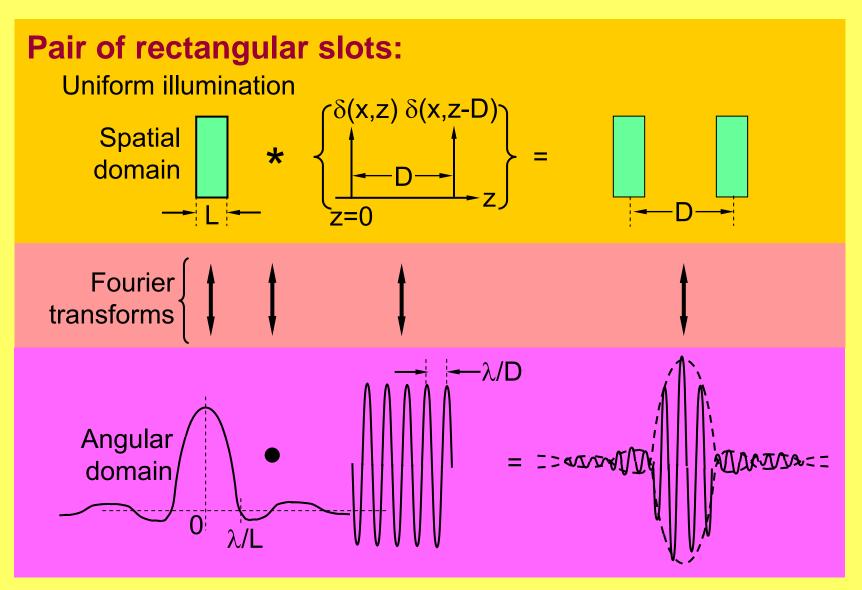
(non-uniform aperture amplitude and/or phase $\Rightarrow A_e < A$)

Huygen's approximation:

$$\overline{\sqsubseteq}_{ff} \cong \hat{\theta} (j/2\lambda r) (1+\cos\beta) \oiint_{A} \sqsubseteq_{oz} (x,z) e^{-jkr_{pq}} dx dz$$

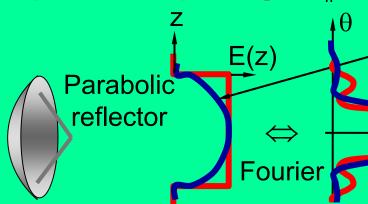
*True for any uniformly illuminated aperture

EXAMPLE: SLIT DIFFRACTION



DIFFRACTION

Aperture tapering: $\underline{\mathbb{E}}_{ff} \cong \hat{\theta} (j/2\lambda r) (1+\cos\beta) \oiint_{A} \underline{\mathbb{E}}_{oz}(x,z) e^{-jkr_{pq}} dx dz$



Tapered illumination reduces sidelobes, broadens beamwidth

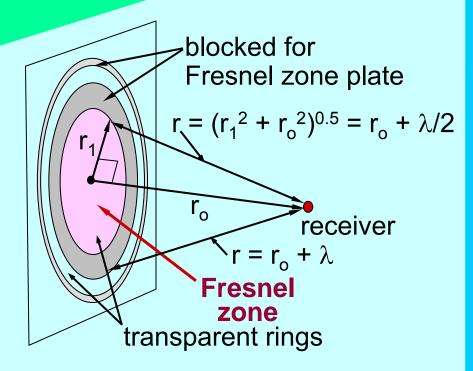
 $G(\theta)$

Fresnel zone plate:

Blocks rings of negative phase, maximizing Huygen's integral

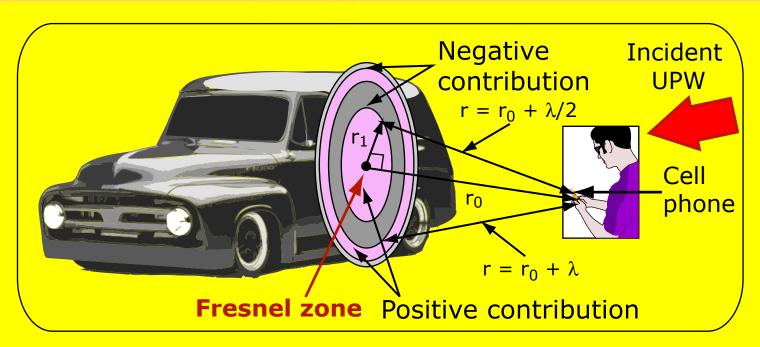
Aperture radius = r_1 maximizes received signal. Additional transparent rings increase it further, yielding lense behavior

Adjacent rings cancel if $\beta \approx 1$



FRESNEL ZONE SIZE, REFLECTIONS

Fresnel zone radius
$$r_1 = \sqrt{(r_o + \frac{\lambda}{2})2 - r_o^2} \cong \sqrt{r_o \lambda} \ \ {\rm for} \ r_o \gg \lambda$$



When the truck is much larger than the Fresnel zone, it acts like a large mirror. When the truck is smaller, the power reflected declines with illuminated zone area on the truck. Thus the reflecting power of a mirror depends on its size and distance relative to a wavelength.

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