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Lecture 19 - Electromagnetic Radiation

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I. Retarded Potentials

A. Maxwell's Equations

$$\begin{aligned}\nabla \times \bar{E} &= -\frac{\partial \bar{B}}{\partial t} \\ \nabla \times \bar{H} &= \bar{J}_f + \frac{\partial \bar{D}}{\partial t} \\ \nabla \cdot \bar{B} &= 0 \\ \nabla \cdot \bar{D} &= \rho_f \\ \bar{B} &= \mu \bar{H}, \bar{D} = \epsilon \bar{E}\end{aligned}$$

B. Vector Identities

$$\begin{aligned}\nabla \times (\nabla \Phi) &= 0 \\ \nabla \cdot (\nabla \times \bar{A}) &= 0 \\ \nabla \times (\nabla \times \bar{A}) &= \nabla(\nabla \cdot \bar{A}) - \nabla^2 \bar{A}\end{aligned}$$

C. Potentials

$$\begin{aligned}\nabla \cdot \bar{B} &= 0 \Rightarrow \bar{B} = \nabla \times \bar{A} \\ \nabla \times \bar{E} &= -\frac{\partial \bar{B}}{\partial t} \Rightarrow \nabla \times \left(\bar{E} + \frac{\partial \bar{A}}{\partial t} \right) = 0 \\ &\Rightarrow \bar{E} + \frac{\partial \bar{A}}{\partial t} = -\nabla \Phi \\ \nabla \times \bar{H} &= \bar{J}_f + \epsilon \frac{\partial \bar{E}}{\partial t} \Rightarrow \frac{1}{\mu} \nabla \times (\nabla \times \bar{A}) = \bar{J}_f + \epsilon \frac{\partial}{\partial t} \left(-\frac{\partial \bar{A}}{\partial t} - \nabla \Phi \right) \\ \nabla \times (\nabla \times \bar{A}) &= \nabla(\nabla \cdot \bar{A}) - \nabla^2 \bar{A} = \mu \bar{J}_f - \epsilon \mu \frac{\partial^2 \bar{A}}{\partial t^2} - \epsilon \mu \nabla \frac{\partial \Phi}{\partial t} \\ \nabla^2 \bar{A} - \nabla \left[\nabla \cdot \bar{A} + \frac{1}{c^2} \frac{\partial \Phi}{\partial t} \right] - \frac{1}{c^2} \frac{\partial^2 \bar{A}}{\partial t^2} &= -\mu \bar{J}_f, c^2 = \frac{1}{\epsilon \mu}\end{aligned}$$

Setting the gauge (Lorentz gauge):

$$\begin{aligned}\nabla \cdot \bar{A} &= -\frac{1}{c^2} \frac{\partial \Phi}{\partial t} \\ \nabla^2 \bar{A} - \frac{1}{c^2} \frac{\partial^2 \bar{A}}{\partial t^2} &= -\mu \bar{J}_f \\ \nabla \cdot \bar{E} = -\frac{\partial}{\partial t} (\nabla \cdot \bar{A}) - \nabla^2 \Phi &= \frac{\rho_f}{\epsilon} \Rightarrow \nabla^2 \Phi - \frac{1}{c^2} \frac{\partial^2 \Phi}{\partial t^2} = -\frac{\rho_f}{\epsilon}\end{aligned}$$

D. Solutions to the Wave Equations

1. Solution in Time Domain

In spherical coordinates:

$$\nabla^2 \Phi = \frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial \Phi}{\partial r} \right) + \frac{1}{r^2 \sin(\theta)} \frac{\partial}{\partial \theta} \left(\sin(\theta) \frac{\partial \Phi}{\partial \theta} \right) + \frac{1}{r^2 \sin^2(\theta)} \frac{\partial^2 \Phi}{\partial \phi^2}$$

Consider a point charge $Q(t)$ at $r = 0$. Then $\frac{\partial}{\partial \theta} = \frac{\partial}{\partial \phi} = 0$, $\Phi = \Phi(r, t)$:

$$\nabla^2 \Phi = \frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial \Phi}{\partial r} \right) = \frac{1}{r} \frac{\partial^2}{\partial r^2} (r \Phi)$$

Then

$$\nabla^2 \Phi - \frac{1}{c^2} \frac{\partial^2 \Phi}{\partial t^2} = 0 \text{ for } r > 0 \Rightarrow \frac{\partial^2}{\partial r^2} (r \Phi) - \frac{1}{c^2} \frac{\partial^2 (r \Phi)}{\partial t^2} = 0$$

This is a wave equation in $r \Phi(r, t)$

$$r \Phi(r, t) = f_+ \left(t - \frac{r}{c} \right) + \underbrace{f_- \left(t + \frac{r}{c} \right)}_{0}$$

No sources for $r > 0$
so only waves
emanating radially
outward from point
charge at $r = 0$

$$\lim_{r \rightarrow 0} \Phi(r, t) = \frac{f_+(t)}{r} = \frac{Q(t)}{4\pi\epsilon r} \Rightarrow f_+(t) = \frac{Q(t)}{4\pi\epsilon}$$

$$\Phi(r, t) = \frac{Q(t - \frac{r}{c})}{4\pi\epsilon r}$$

2. Solution in Frequency Domain

$$\Phi(r, t) = \operatorname{Re} [\hat{\Phi}(r) e^{j\omega t}], Q(t) = \operatorname{Re} [\hat{Q} e^{j\omega t}]$$

$$\frac{d^2}{dr^2} \left(r \hat{\Phi}(r) \right) + \frac{\omega^2}{c^2} r \hat{\Phi}(r) = 0$$

$$r \hat{\Phi}(r) = A_1 e^{-jkr} + \underbrace{A_2 e^{+jkr}}_0, k = \frac{\omega}{c}$$

No $-r$ traveling wave

$$\lim_{r \rightarrow 0} \hat{\Phi}(r) = \frac{A_1}{r} = \frac{\hat{Q}}{4\pi\epsilon r} \Rightarrow A_1 = \frac{\hat{Q}}{4\pi\epsilon}$$

$$\hat{\Phi}(r) = \frac{\hat{Q}}{4\pi\epsilon r} e^{-jkr} \Rightarrow \Phi(r, t) = \operatorname{Re} \left[\underbrace{\frac{\hat{Q}}{4\pi\epsilon r} e^{j\omega(t - \frac{r}{c})}}_{\text{Spherical Wave}} \right]$$

Vector Potential

$$\nabla \cdot \bar{A} = -\frac{1}{c^2} \frac{\partial \Phi}{\partial t} \Rightarrow \nabla \cdot \hat{A} = -\frac{j\omega}{c^2} \hat{\Phi}$$

$$\nabla \cdot \hat{A} = \frac{1}{r^2} \frac{\partial}{\partial r} (r^2 \hat{A}_r) + \frac{1}{r \sin(\theta)} \frac{\partial}{\partial \theta} (\sin(\theta) \hat{A}_\theta) + \frac{1}{r \sin(\theta)} \frac{\partial \hat{A}_\phi}{\partial \phi}^0$$

$$\frac{\partial}{\partial \theta} = \frac{\partial}{\partial \phi} = 0$$

$$\begin{aligned} \frac{1}{r^2} \frac{\partial}{\partial r} (r^2 \hat{A}_r) &= -\frac{j\omega}{c^2} \hat{\Phi} = -j\omega \mu \frac{\hat{Q}}{4\pi \epsilon r} e^{-jkr} \\ \frac{\partial}{\partial r} (r^2 \hat{A}_r) &= \frac{d}{dr} (r^2 \hat{A}_r) = -\frac{j\omega \mu}{4\pi} \hat{Q} r e^{-jkr} \\ r^2 \hat{A}_r &= -\frac{j\omega \mu}{4\pi} \hat{Q} \int dr r e^{-jkr} \\ &= -\frac{j\omega \mu \hat{Q}}{4\pi} \frac{e^{-jkr}}{(k^2)} (+jkr + 1) \\ \hat{A}_r &= -\frac{j\omega \mu \hat{Q}}{4\pi k^2 r^2} (1 + jkr) e^{-jkr} \\ &= -\frac{j\omega \mu \hat{Q}}{4\pi \omega \epsilon \mu r^2} (1 + jkr) e^{-jkr} \\ &= -\frac{j\hat{Q}}{4\pi \omega \epsilon r^2} (1 + jkr) e^{-jkr} \end{aligned}$$

3. Solutions for Volume Charge and Current Distributions

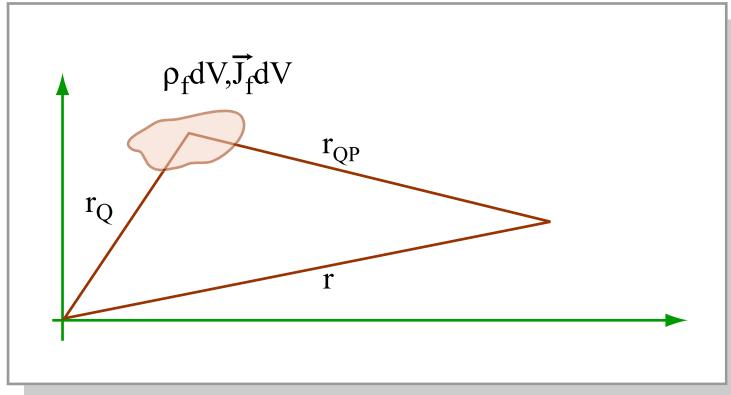


Image by MIT OpenCourseWare.

$$\Phi(r, \theta, \phi, t) = \int_{\text{all charge}} dV \frac{\rho_f (t - \frac{r_{QP}}{c})}{4\pi \epsilon r_{QP}}$$

$$\nabla^2 \bar{A} - \frac{1}{c^2} \frac{\partial^2 \bar{A}}{\partial t^2} = -\mu \bar{J}_f \Rightarrow \bar{A} = \int_{\text{all current}} dV \frac{\mu \bar{J}_f (t - \frac{r_{QP}}{c})}{4\pi r_{QP}}$$

II. Radiation From a Point Electric Dipole

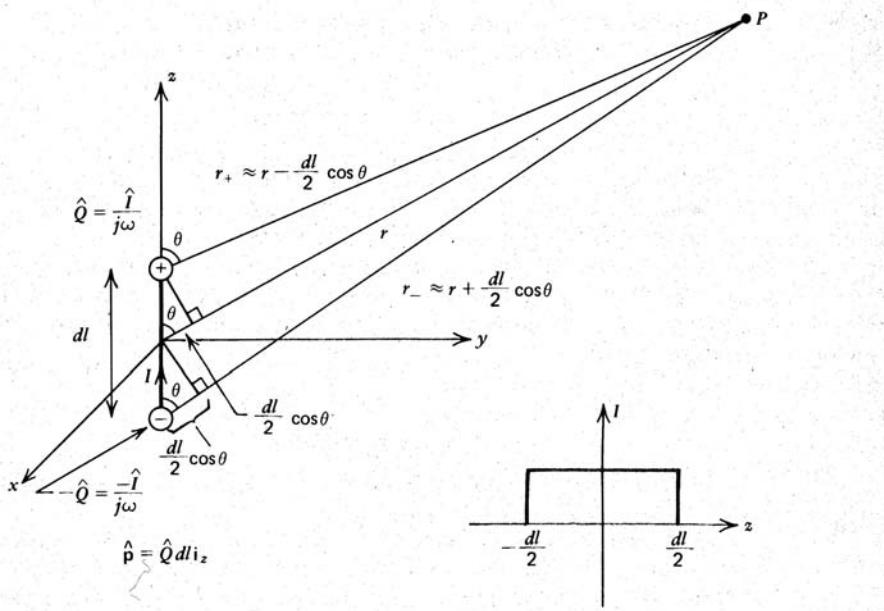


Figure 9-1 A point dipole antenna is composed of a very short uniformly distributed current-carrying wire. Because the current is discontinuous at the ends, equal magnitude but opposite polarity charges accumulate there forming an electric dipole.

From *Electromagnetic Field Theory: A Problem Solving Approach*, by Markus Zahn, 1987. Used with permission.

$$\begin{aligned}\bar{i}(t) &= \operatorname{Re} \left[\hat{I} e^{j\omega t} \bar{i}_z \right] \\ \hat{Q} &= \frac{\hat{I}}{j\omega} \text{ at } z = \pm \frac{dl}{2} \\ \hat{p} &= \hat{Q} dl \bar{i}_z \quad (\text{dipole moment})\end{aligned}$$

A. Vector Potential

$$\begin{aligned}A_z(r, \theta, t) &= \operatorname{Re} \left[\hat{A}_z(r, \theta) e^{j\omega t} \right], \frac{\partial}{\partial \phi} = 0 \\ A_z(r, t) &= \operatorname{Re} \left[\int_{-\frac{dl}{2}}^{+\frac{dl}{2}} dz \frac{\mu \hat{I} e^{j\omega(t - \frac{r_{QP}}{c})}}{4\pi r_{QP}} \right] \\ &= \operatorname{Re} \left[\frac{\mu \hat{I}}{4\pi r} e^{j\omega(t - \frac{r}{c})} \right] \\ \hat{A}_z(r) &= \frac{\mu \hat{I} dl}{4\pi r} e^{-jkr}, k = \frac{\omega}{c} \\ A_z(r, t) &= \operatorname{Re} \left[\hat{A}_z(r) e^{j\omega t} \right] \\ \bar{i}_z &= \bar{i}_r \cos(\theta) - \bar{i}_\theta \sin(\theta) \\ \bar{A} &= A_z \bar{i}_z = A_z (\bar{i}_r \cos(\theta) - \bar{i}_\theta \sin(\theta)) \\ \hat{A}_r &= \hat{A}_z \cos(\theta), \hat{A}_\theta = -\hat{A}_z \sin(\theta)\end{aligned}$$

B. Scalar Potential

$$\begin{aligned}
\hat{\Phi} &= \frac{jc^2}{\omega} \nabla \cdot \hat{A} = \frac{jc^2}{\omega} \left[\frac{1}{r^2} \frac{\partial}{\partial r} (r^2 \hat{A}_r) + \frac{1}{r \sin(\theta)} \frac{\partial}{\partial \theta} (\sin(\theta) \hat{A}_\theta) \right] \\
&= \frac{j}{\omega \epsilon \mu} \frac{\hat{I} dl}{4\pi} \left[\frac{\cos(\theta)}{r^2} \frac{d}{dr} \left(\frac{r^2 e^{-jkr}}{r} \right) - \frac{\sin(\theta)}{r \sin(\theta)} \frac{\partial}{\partial \theta} \left(\frac{\sin(\theta)}{r} e^{-jkr} \right) \right] \\
&= \frac{-j \hat{I} dl (1 + jkr) e^{-jkr}}{4\pi \omega \epsilon r^2} \cos(\theta) \\
&= \frac{\hat{Q} dl}{4\pi \epsilon r^2} (1 + jkr) e^{-jkr} \cos(\theta)
\end{aligned}$$

C. Magnetic Field

$$\begin{aligned}
\hat{H} &= \frac{1}{\mu} \nabla \times \hat{A} = \frac{1}{\mu} \left[\frac{1}{r \sin(\theta)} \left(\frac{\partial}{\partial \theta} \left(\hat{A}_\phi \overset{0}{\cancel{\sin}}(\theta) \right) - \frac{\partial \hat{A}_\theta}{\partial \phi} \right) \bar{i}_r \right. \\
&\quad \left. + \frac{\bar{i}_\theta}{r} \left(\frac{1}{\sin(\theta)} \frac{\partial \hat{A}_r}{\partial \phi} \overset{0}{\cancel{\phi}} - \frac{\partial}{\partial r} \left(r \hat{A}_\phi \overset{0}{\cancel{\sin}} \right) \right) + \frac{\bar{i}_\phi}{r} \left(\frac{\partial}{\partial r} (r A_\theta) - \frac{\partial A_r}{\partial \theta} \right) \right] \\
&\quad (\text{with } A_\phi = 0, \frac{\partial}{\partial \phi} = 0) \\
\bar{H} &= \frac{1}{\mu} \nabla \times \bar{A} = \bar{i}_\phi \frac{1}{\mu r} \left(\frac{\partial}{\partial r} (r A_\theta) - \frac{\partial A_r}{\partial \theta} \right) \\
\hat{H}_\phi &= -\bar{i}_\phi \frac{\hat{I} dl k^2}{4\pi} \sin(\theta) \left(\frac{1}{jkr} + \frac{1}{(jkr)^2} \right) e^{-jkr}
\end{aligned}$$

D. Electric Field

$$\begin{aligned}
\hat{E} &= \frac{1}{j\omega\epsilon} \nabla \times \hat{H} = \frac{1}{j\omega\epsilon} \left[\frac{1}{r \sin(\theta)} \frac{\partial}{\partial \theta} (\hat{H}_\phi \sin(\theta)) \bar{i}_r - \frac{1}{r} \frac{\partial}{\partial r} (r \hat{H}_\phi) \bar{i}_\theta \right] \\
&= -\frac{\hat{I} dl k^2}{4\pi} \sqrt{\frac{\mu}{\epsilon}} \left[\bar{i}_r \cdot 2 \cos(\theta) \left(\frac{1}{(jkr)^2} + \frac{1}{(jkr)^3} \right) \right. \\
&\quad \left. + \bar{i}_\theta \sin(\theta) \left(\frac{1}{jkr} + \frac{1}{(jkr)^2} + \frac{1}{(jkr)^3} \right) \right] e^{-jkr}
\end{aligned}$$

E. Far Field Limit ($kr \gg 1$)

$$\lim_{kr \gg 1} \underbrace{\hat{E}_\theta}_{\substack{\text{Same relationship} \\ \text{as plane waves}}} = \sqrt{\frac{\mu}{\epsilon}} \hat{H}_\phi = \frac{\hat{E}_0}{jkr} \sin(\theta) e^{-jkr}, \hat{E}_0 = -\frac{\hat{I} dl k^2}{4\pi} \sqrt{\frac{\mu}{\epsilon}}$$

III. Power Density

$$\begin{aligned}
\langle \bar{S} \rangle &= \frac{1}{2} \operatorname{Re} \left(\hat{\bar{E}} \times \hat{\bar{H}}^* \right) \\
\hat{\bar{E}} &= \hat{E}_0 \left[2 \cos(\theta) \bar{i}_r \left(\frac{1}{(jkr)^2} + \frac{1}{(jkr)^3} \right) + \sin(\theta) \bar{i}_\theta \left(\frac{1}{jkr} + \frac{1}{(jkr)^2} + \frac{1}{(jkr)^3} \right) \right] e^{-jkr} \\
\hat{\bar{H}} &= \frac{\hat{E}_0}{\eta} \bar{i}_\phi \sin(\theta) \left(\frac{1}{jkr} + \frac{1}{(jkr)^2} \right) e^{-jkr} \\
\langle \bar{S} \rangle &= \operatorname{Re} \left\{ \frac{|\hat{E}_0|^2}{2\eta} \sin(\theta) \left[2 \cos(\theta) \left(\frac{1}{(jkr)^2} + \frac{1}{(jkr)^3} \right) \right] \left[-\frac{1}{jkr} + \frac{1}{(jkr)^2} \right] \underbrace{\bar{i}_r \times \bar{i}_\phi}_{-\bar{i}_\theta} \right. \\
&\quad \left. + \sin(\theta) \left(\frac{1}{jkr} + \frac{1}{(jkr)^2} + \frac{1}{(jkr)^3} \right) \left(-\frac{1}{jkr} + \frac{1}{(jkr)^2} \right) \underbrace{\bar{i}_\theta \times \bar{i}_\phi}_{\bar{i}_r} \right\} \\
\langle \bar{S} \rangle &= \frac{|\hat{E}_0|^2}{2\eta} \sin(\theta) \operatorname{Re} \left\{ -2 \cos(\theta) \left(\underbrace{-\frac{1}{(jkr)^3}}_{\text{imag}} - \underbrace{\frac{1}{(jkr)^4}}_{\text{cancel}} + \underbrace{\frac{1}{(jkr)^4}}_{\text{cancel}} + \underbrace{\frac{1}{(jkr)^5}}_{\text{imag}} \right) \bar{i}_\theta \right. \\
&\quad \left. + \sin(\theta) \left(\underbrace{-\frac{1}{(jkr)^2}}_{\text{real}} + \underbrace{\frac{1}{(jkr)^3}}_{\text{imaginary/cancel}} - \underbrace{\frac{1}{(jkr)^3}}_{\text{cancel}} + \underbrace{\frac{1}{(jkr)^4}}_{\text{cancel}} - \underbrace{\frac{1}{(jkr)^4}}_{\text{cancel}} + \underbrace{\frac{1}{(jkr)^5}}_{\text{imag}} \right) \bar{i}_r \right\} \\
&= \frac{|\hat{E}_0|^2 \sin^2(\theta)}{2\eta} \bar{i}_r, \hat{E}_0 = -\frac{\hat{I} dl k^2}{4\pi} \eta, \eta = \sqrt{\frac{\mu}{\epsilon}}
\end{aligned}$$

IV. Total Time Average Radiated Power

A. Power

$$\begin{aligned}
\langle P \rangle &= \int_0^\pi d\theta \int_0^{2\pi} d\phi \langle S_r \rangle r^2 \sin(\theta) \\
&= \frac{|\hat{E}_0|^2}{2\eta k^2} \int_0^\pi d\theta \int_0^{2\pi} d\phi \sin^3(\theta) \\
&= \frac{|\hat{E}_0|^2 2\pi}{2\eta k^2} \int_0^\pi d\theta \sin^3(\theta) \\
&= \frac{\pi |\hat{E}_0|^2}{\eta k^2} \underbrace{\left[-\frac{1}{3} \cos(\theta) (\sin^2(\theta) + 2) \right]}_{\frac{4}{3}} \Big|_0^\pi \\
&= \frac{4\pi}{3\eta k^2} |\hat{E}_0|^2 \\
&= \frac{4\pi}{3\eta k^2} \frac{|\hat{I} dl|^2 k^4 2}{4 \cdot 16\pi^2} \eta^2 \\
&= \frac{|\hat{I} dl|^2 \eta k^2}{12\pi}
\end{aligned}$$

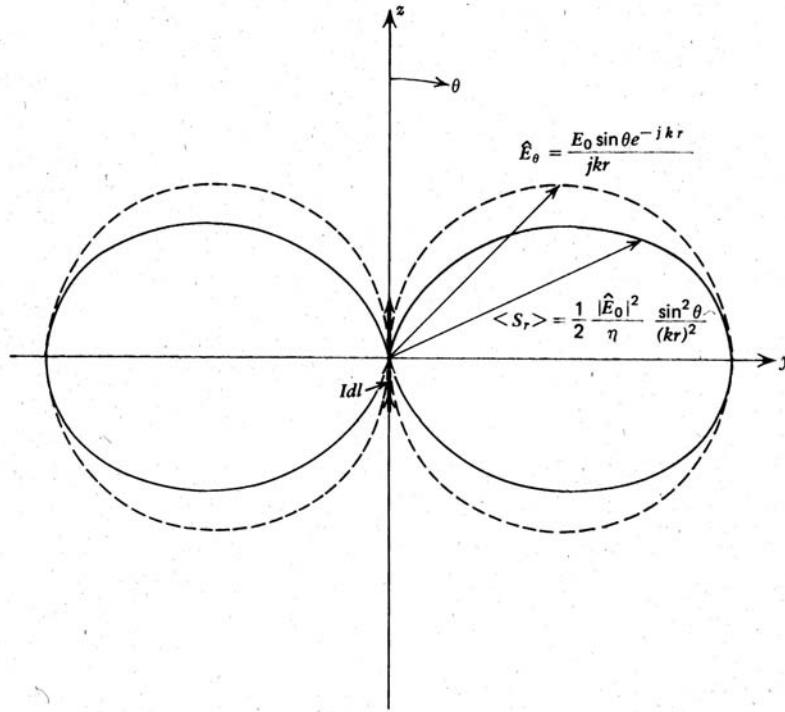


Figure 9-3 The strength of the electric field and power density due to a z-directed point dipole as a function of angle θ is proportional to the length of the vector from the origin to the radiation pattern.

From *Electromagnetic Field Theory: A Problem Solving Approach*, by Markus Zahn, 1987. Used with permission.

Antenna Gain:

$$\begin{aligned}
 G(\theta, \phi) &= \frac{\langle S \rangle_r}{\langle P \rangle / 4\pi r^2} \\
 &= \frac{|\hat{E}_0|^2 \sin^2(\theta) \cdot 3\eta k^2 \cdot 4\pi r^2}{2\eta k^2 r^2 \cdot 4\pi |\hat{E}_0|^2} \\
 &= \frac{3}{2} \sin^2(\theta)
 \end{aligned}$$

B. Radiation Resistance R

$$\begin{aligned}
 \langle P \rangle &= \frac{|\hat{I}dl|^2 \eta k^2}{12\pi} = \frac{1}{2} |\hat{I}|^2 R \\
 R &= \frac{\eta}{6\pi} (kdl)^2 = \frac{\eta}{6\pi} 4\pi \left(\frac{dl}{\lambda} \right)^2 \\
 &= \frac{2}{3} \pi \eta \left(\frac{dl}{\lambda} \right)^2
 \end{aligned}$$

In free space: $\eta_0 = \sqrt{\frac{\mu_0}{\epsilon_0}} \approx 120\pi \Rightarrow R = 80\pi^2 \left(\frac{dl}{\lambda} \right)^2$

C. Effective Length of Antenna: $\hat{I}(z)$

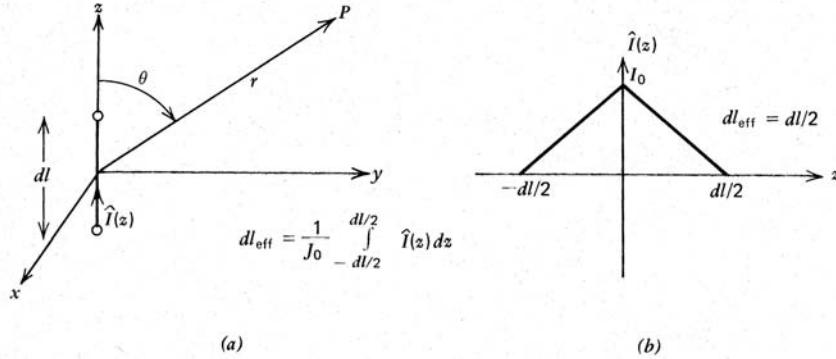


Figure 9-4 (a) If a point electric dipole has a nonuniform current distribution, the solutions are of the same form if we replace the actual dipole length dl by an effective length dl_{eff} . (b) For a triangular current distribution the effective length is half the true length.

From *Electromagnetic Field Theory: A Problem Solving Approach*, by Markus Zahn, 1987. Used with permission.

If $\frac{dl}{\lambda} \ll 1$, but current distribution depends on z : $\hat{I}(z)$

$$\tilde{A}_z(r) = \int_{-\frac{dl}{2}}^{+\frac{dl}{2}} dz \frac{\mu \hat{I}(z) e^{-jkr_{QP}}}{4\pi r_{QP}} \approx \frac{\mu e^{-jkr}}{4\pi r} \int_{-\frac{dl}{2}}^{+\frac{dl}{2}} dz \hat{I}(z)$$

$$dl_{\text{eff}} = \frac{1}{\hat{I}(z=0)} \int_{-\frac{dl}{2}}^{+\frac{dl}{2}} dz \hat{I}(z)$$

Example:

$$\hat{I}(z) = \begin{cases} I_0 \left(1 - \frac{2z}{dl}\right) & 0 < z < \frac{dl}{2} \\ I_0 \left(1 + \frac{2z}{dl}\right) & -\frac{dl}{2} < z < 0 \end{cases}$$

$$\hat{I}(z=0) = I_0$$

$$dl_{\text{eff}} = \frac{1}{I_0} \int_{-\frac{dl}{2}}^{+\frac{dl}{2}} dz \hat{I}(z) = \frac{1}{I_0} \left(\frac{J_0 dl}{2} \right) = \frac{dl}{2}$$

$$R = \frac{2\pi\eta}{3} \left(\frac{dl_{\text{eff}}}{\lambda} \right)^2 = 20\pi^2 \sqrt{\frac{\mu_r}{\epsilon_r}} \left(\frac{dl}{\lambda} \right)^2$$

D. Rayleigh Scattering (or, why is the sky blue?)

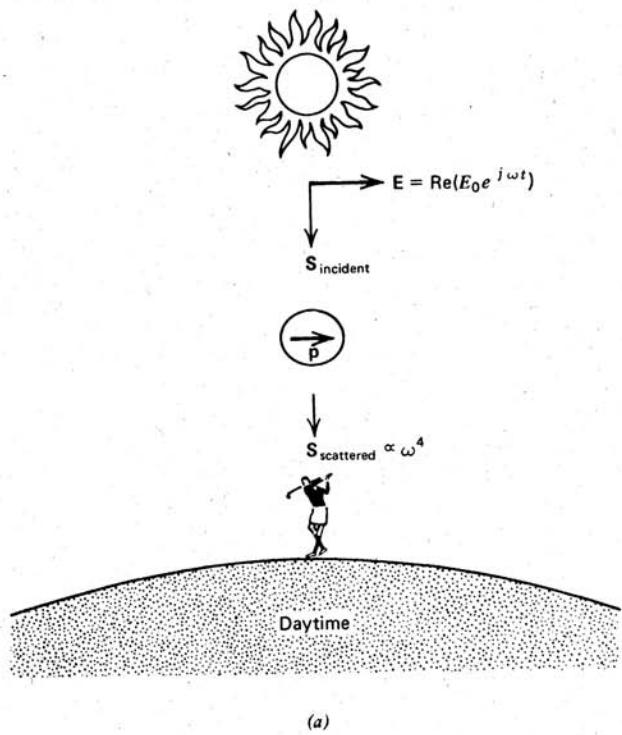
Electric field of light from the sun polarizes particles with dipole moment \bar{p} in direction of \bar{E} .

$$\hat{p} = \hat{Q} dl = \frac{\hat{I} dl}{j\omega} \Rightarrow |\hat{I} dl| = |\omega \hat{p}|$$

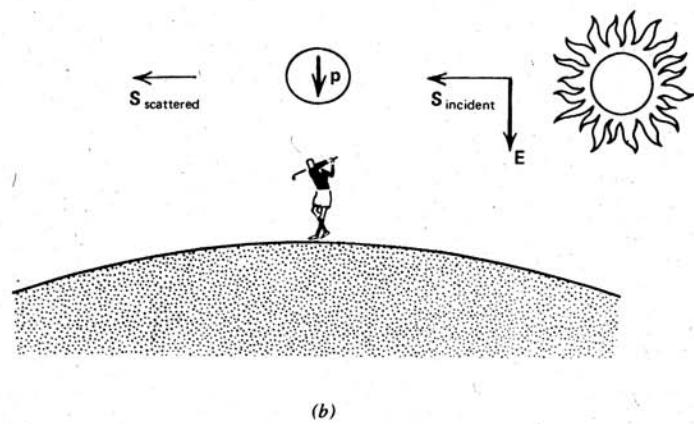
$$\langle P \rangle = \frac{|\hat{I} dl|^2 \eta k^2}{12\pi}$$

$$= |\omega \hat{p}|^2 \frac{\eta \omega^2 \epsilon \mu}{12\pi}$$

$$= \omega^4 |\hat{p}|^2 \frac{\eta \epsilon \mu}{12\pi}$$



(a)



(b)

Figure 9-5 An incident electric field polarizes dipoles that then re-radiate their energy primarily perpendicular to the polarizing electric field. The time-average scattered power increases with the fourth power of frequency so shorter wavelengths of light are scattered more than longer wavelengths. (a) During the daytime an earth observer sees more of the blue scattered light so the sky looks blue (short wavelengths). (b) Near sunset the light reaching the observer lacks blue so the sky appears reddish (long wavelength).

From *Electromagnetic Field Theory: A Problem Solving Approach*, by Markus Zahn, 1987. Used with permission.

Higher frequencies, and thus shorter wavelengths, are scattered in direction of S_{incident} :

$$\text{red: } \lambda \approx 633 \text{ nm}$$

$$\text{blue: } \lambda \approx 450 \text{ nm}$$

Blue light is thus scattered more than red.