

# THE ENVIRONMENT OF SPACE

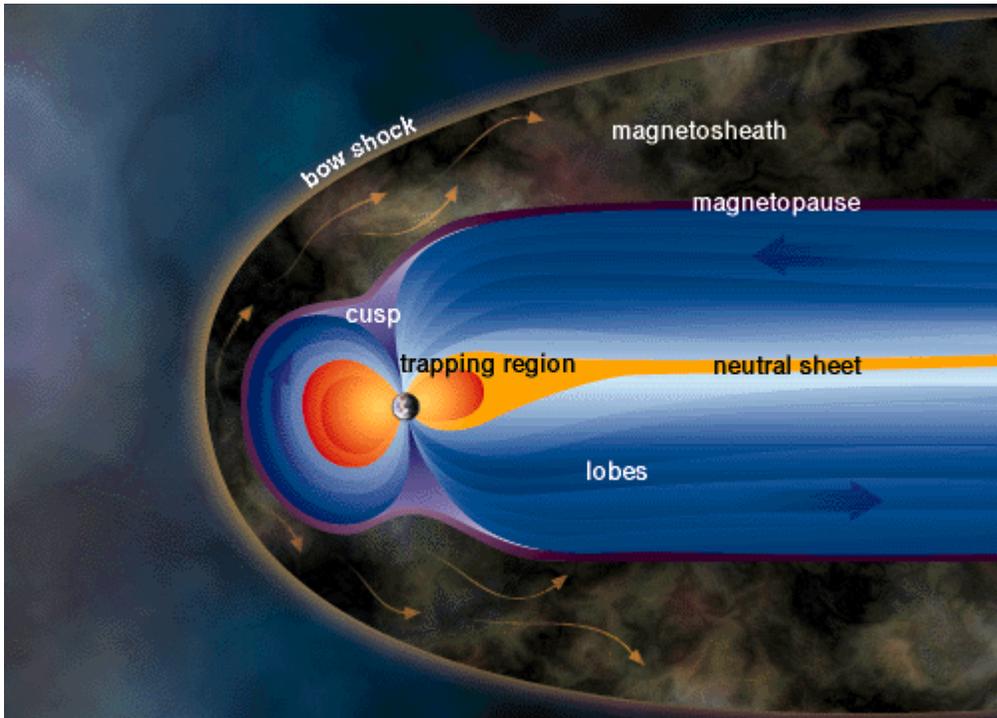


Image courtesy of NASA.

Col. John Keesee

# OUTLINE

- Overview of effects
- Solar Cycle
- Gravity
- Neutral Atmosphere
- Ionosphere
- GeoMagnetic Field
- Plasma
- Radiation

# OVERVIEW OF THE EFFECTS OF THE SPACE ENVIRONMENT

- Outgassing in near vacuum
- Atmospheric drag
- Chemical reactions
- Plasma-induced charging
- Radiation damage of microcircuits, solar arrays, and sensors
- Single event upsets in digital devices
- Hyper-velocity impacts

# Solar Cycle

- Solar Cycle affects all space environments.
- Solar intensity is highly variable
- Variability caused by distortions in magnetic field caused by differential rotation
- Indicators are sunspots and flares

# LONG TERM SOLAR CYCLE INDICES

- Sunspot number  $R$   
 $10$  (solar min)  $\leq R \leq 150$  (solar max)
- Solar flux  $F_{10.7}$   
Radio emission line of Fe (2800 MHz)  
Related to variation in EUV  
Measures effect of sun on our atmosphere  
Measured in solar flux units ( $10^{-22}$  w/m<sup>2</sup>)  
 $50$  (solar min)  $\leq F_{10.7} \leq 240$  (solar max)

# SHORT TERM SOLAR CYCLE INDEX

- Geomagnetic Index  $A_p$ 
  - Daily average of maximum variation in the earth's surface magnetic field at mid latitude (units of  $2 \times 10^{-9}$  T)

$A_p = 0$  quiet

$A_p = 15$  to  $30$  active

$A_p > 50$  major solar storm

# GRAVITY

$$\text{force} = -G \frac{m_1 m_2}{r^2} \hat{r}$$
$$G = 6.672 \times 10^{-11} \text{m}^3 \text{kg}^{-1}$$

At surface of earth

$$f_g = -m \frac{G m_e}{R_E^2} \quad g = \frac{G m_e}{R_E^2} \approx 9.8 \frac{\text{m}}{\text{sec}^2}$$

# MICROGRAVITY

- Satellites in orbit are in free fall - accelerating radially toward earth at the rate of free fall.

- Deviations from zero-g

- Atmospheric drag  $\ddot{x} = 0.5 \left( \frac{C_D A}{m} \right) \rho a^2 \omega^2$

- Gravity gradient

$$\ddot{x} = -x\omega^2 \quad \ddot{y} = -y\omega^2 \quad \ddot{z} = 2z\omega^2$$

- Spacecraft rotation

$$\ddot{x} = x\omega^2 \quad \ddot{z} = z\omega^2 \quad (\text{rotation about Y axis})$$

- Coriolis forces

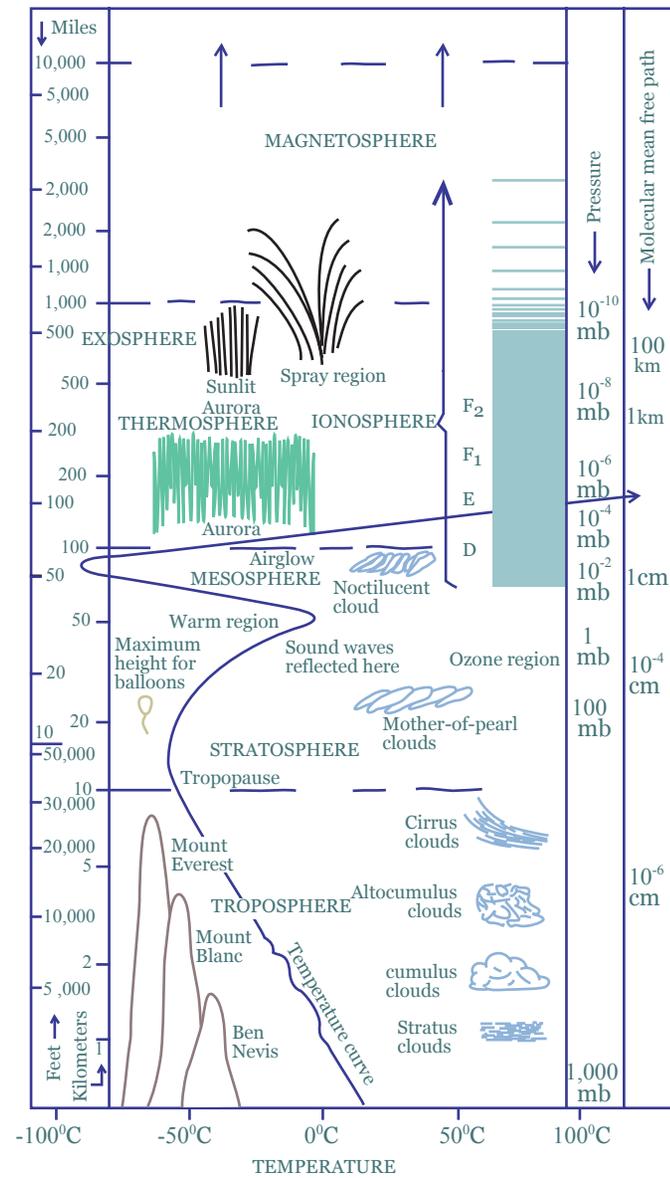
$$\ddot{x} = 2z\omega \quad \ddot{y} = 0 \quad \ddot{z} = -2x\omega$$

# ATMOSPHERIC MODEL

## NEUTRAL ATMOSPHERE

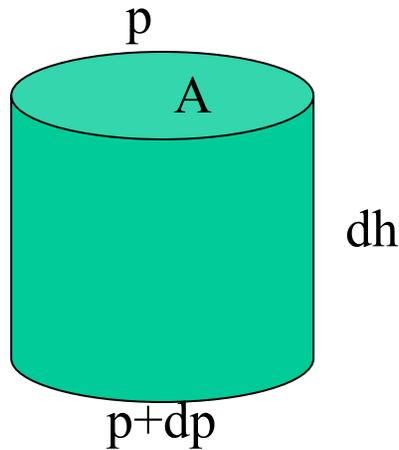
- Turbo sphere (0 ~ 120Km) is well mixed (78% N<sub>2</sub>, 21% O<sub>2</sub>)
  - Troposphere (0 ~ 10Km) warmed by earth as heated by sun
  - Stratosphere (10 ~ 50 Km) heated from above by absorption of UV by O<sub>3</sub>
  - Mesosphere (50 ~ 90Km) heated by radiation from stratosphere, cooled by radiation into space
  - Thermosphere (90 ~ 600Km) very sensitive to solar cycle, heated by absorption of EUV.
- Neutral atmosphere varies with season and time of day

# Layers of the Earth's Atmosphere



# DENSITY ALTITUDE MODEL

Assume perfect gas and constant temperature



$$p = n k T \quad \frac{dp}{dh} = \frac{d(nkT)}{dh}$$

$n$  is number density (number/m<sup>3</sup>)  $dpA - n m g A dh = 0$

$k$  is Boltzmann's constant

$M$  is average molecular mass

$H \sim 8.4\text{km}$   $h \sim 120\text{km}$

$H \equiv kT/mg$  (scale height)

$$\frac{dp}{dh} = nMg = \frac{d(nkT)}{dh}$$

$$\frac{dn}{n} = \frac{Mg}{KT} dh$$

$$n = n_0 \exp(-h/H)$$

# Atmospheric Gases

- At higher altitudes O<sub>2</sub> breaks down into O by UV
- Primarily O from 80 - 90 km to 500 km
- Hydrogen and Helium beyond 500 km
- Kinetic energy of O atom at 7.8 km/s  $\sim 5\text{eV}$  (enough to break molecular bonds  $\sim 1 - 2\text{eV}$ )
- O is highly reactive and destructive to spacecraft
- Temperature at LEO increases with altitude
- Atmosphere expands when heated by high UV (solar max)
- LEO densities  $\sim 10^8$  particles/cm<sup>3</sup>

# ATMOSPHERIC MODEL

Most common Mass Spectrometer and Incoherent Scatter model - 1986 (MSIS - 1986)

- Based on measured data
- Requires  $A_p$ ,  $F_{10.7}$ , month as input
- Gives average values of  $n$ ,  $n_o$ ,  $T$ , atomic mass as function of altitude
- Instantaneous values can vary by factor of 10

<http://nssdc.gsfc.nasa.gov/space/model/atmos/msis.html>

# AERODYNAMIC DRAG

Drag

$$\bar{D} = -\frac{1}{2} \rho \bar{v} \cdot \bar{v} \left( \frac{\bar{v}}{|\bar{v}|} \right) C_D A$$

$$\bar{D} = m \frac{d\bar{v}}{dt}$$

$$\Delta v = \frac{1}{2} \rho v^2 \left[ \frac{C_D A}{m} \right] \Delta t$$

Ballistic coefficient  $\beta = \left[ \frac{m}{C_D A} \right]$

$\rho$  = density of the atmosphere =  $m_o n_o$

$$= 16 \times 1.67 \times 10^{-27} \times 10^{13} = 2.67 \times 10^{-13} \text{ kg/m}^3$$

$V = 7.8 \text{ km/s}$

$C_D$  - Drag coefficient

$A$  - Cross sectional area

# DRAG COEFFICIENTS

Derived from Newtonian Aerodynamics. Depends on what air molecule does at impact

– Reflected  $C_D = 4$

– Absorbed  $C_D = 2$

Since  $F = d(mv)/dt$

$$D = -F = -d(mv)/dt$$

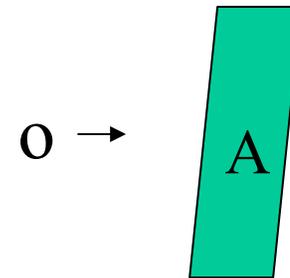
$$C_D = \frac{D}{\frac{1}{2}\rho V^2 A} = -\frac{m(V_f - V_i)}{\frac{1}{2}\rho V_i^2 A dt}$$

$$m = \frac{2}{\rho} A v_i dt$$

$$C_D = -2 (v_f - v_i)/v_i$$

$$= 2 \text{ if } v_f = 0$$

$$= 4 \text{ if } v_f = -v_i$$



in rarefied atmosphere

# TYPICAL DRAG PARAMETERS

	$\beta$ (kg/m <sup>2</sup> )	$C_D$	
LANDSAT	25 - 123	3.4 - 4	
ERS - 1	12 - 135	4	
Hubble	29 - 192	3.3 - 4	90,000Kg
Echo 1	0.515	2	

Typically  $C_D \sim 2.2 - 4$  for spacecraft. (see SMAD Table 8.3)

$\Delta V$  over one year ( $\beta = 100$  kg/m<sup>2</sup>)

<u>h (km)</u>	<u><math>\Delta V</math> /year (m/s)</u>	
100	$10^7$	
200	$2 - 5 \times 10^3$	solar (min - max)
300	40 - 600	
400	3 - 200	

# SATELLITE LIFETIMES

Large variation depending on initial altitude and solar min/max condition (see SMAD Fig. 8 - 4)

At LEO, design must compensate for effects of drag.

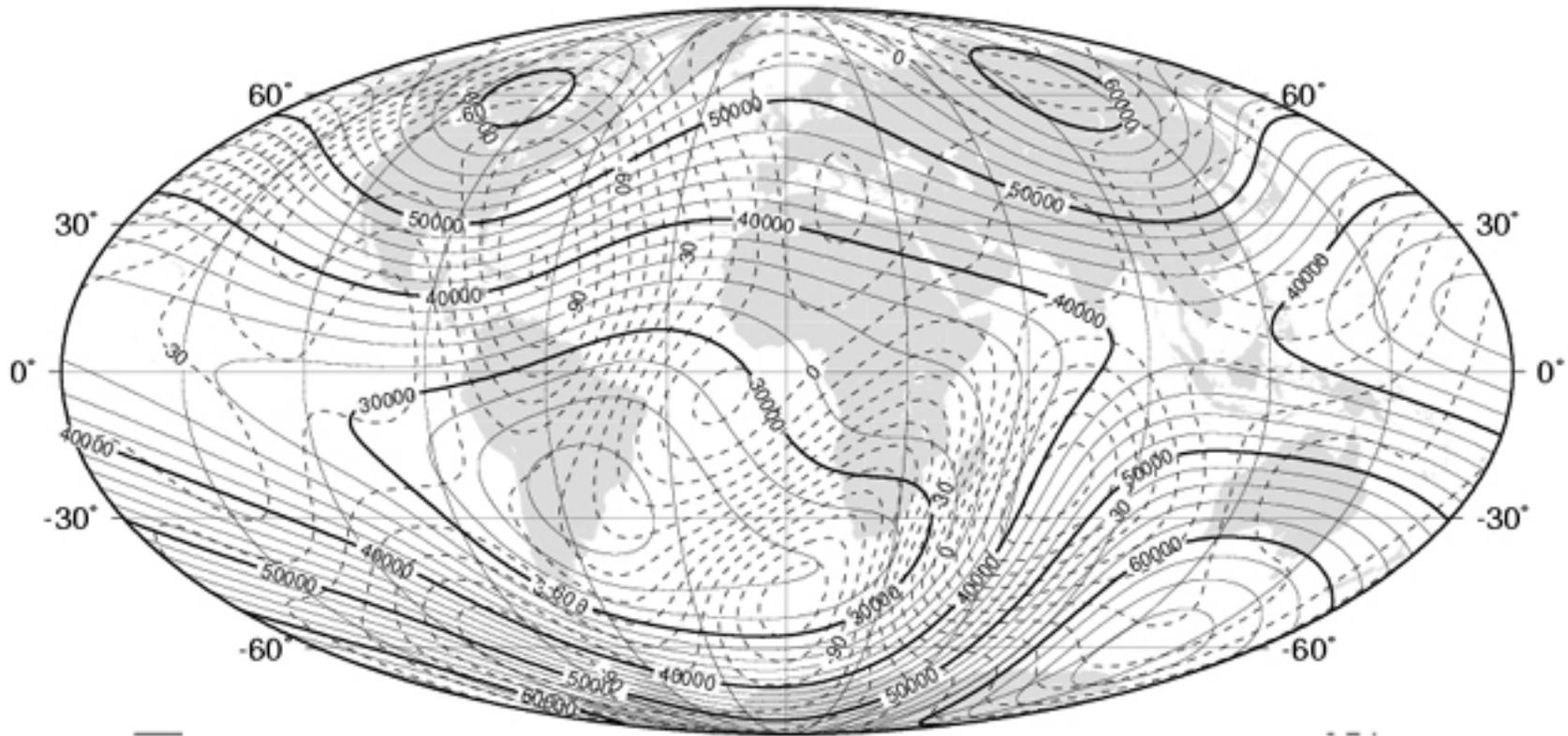
# MAGNETIC FIELD EFFECTS

- Deflects charged particles/solar wind.
  - South Atlantic Anomaly
- Creates the structure of the ionosphere/plasmasphere
  - Magnetosphere
  - Van Allen radiation belts
- Direct effects on Spacecraft systems
  - Avionics - induced potential effects
  - Power - induced potential effects
  - GN&C - magnetic torquer performance, sizing
  - Structures - induced currents
  - TT&C - location of SAA

# GEOMAGNETIC FIELD

- Earth's Magnetic field comes from three sources
  - internal field (99%)
    - currents inside the Earth
    - residual magnetism of elements contained in crust
  - External field 1%
    - Currents in the magnetosphere
- $B_i$  internal field varies slowly on the order of 100 years (0.05%/year.)
- Poles of magnetic field lie in Siberia and South Australia.

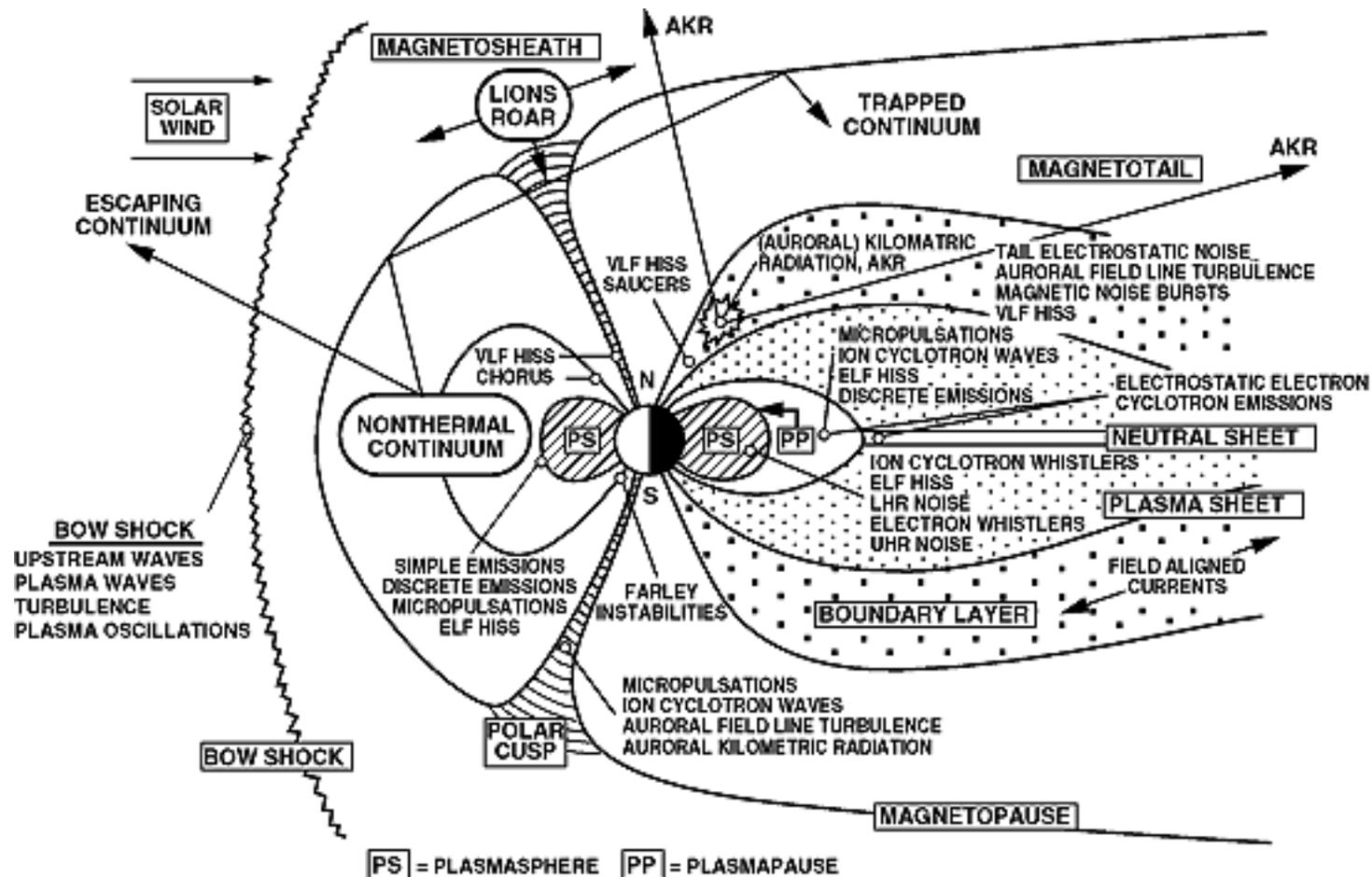
# GEOMAGNETIC FIELD



Units (Total Intensity) : nanoTeslas  
Contour Interval : 2000 nanoTeslas  
Map Projection : Mollweide

Units (Annual Change) : nanoTeslas/yr  
Contour Interval : 10 nanoTeslas/yr  
Map Projection : Mollweide

# MAGNETOSPHERE



# Magnetosphere (continued)

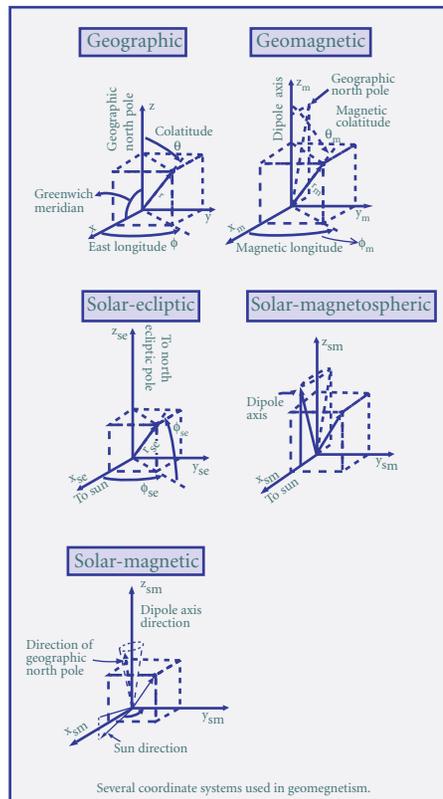
- Earth's field extends 10 Earth Radii ( $R_E$ ) toward the sun
  - terminates at magneto pause
- Earth's field slows and deflects solar wind
  - Compressed, heated, turbulent
  - Bow shock at about  $14 R_E$
- Polar field lines are swept back in night-side tail
  - Does not close
  - Neutral sheet
- Surface of discontinuity in magnetic field implies current flow in the surface
  - Sunward magnetopause - eastward current flow across sub-solar point.
  - Neutral sheet current flow is westward across the tail

# EXTERNAL MAGNETIC FIELD

- $B_e$  generated by ring currents and solar wind. Large variation with time
  - Milliseconds to 11-year cycle scales.
- Variations caused by
  - Magnetosphere fluctuations (geomagnetic storms)
  - Solar activity
- Geomagnetic storms dump large numbers of charged particles from magnetosphere into atmosphere
  - Ionizes and heats the atmosphere
  - Altitudes from 300 km to over 1000 km
  - Persist 8-12 hours after storm subsides

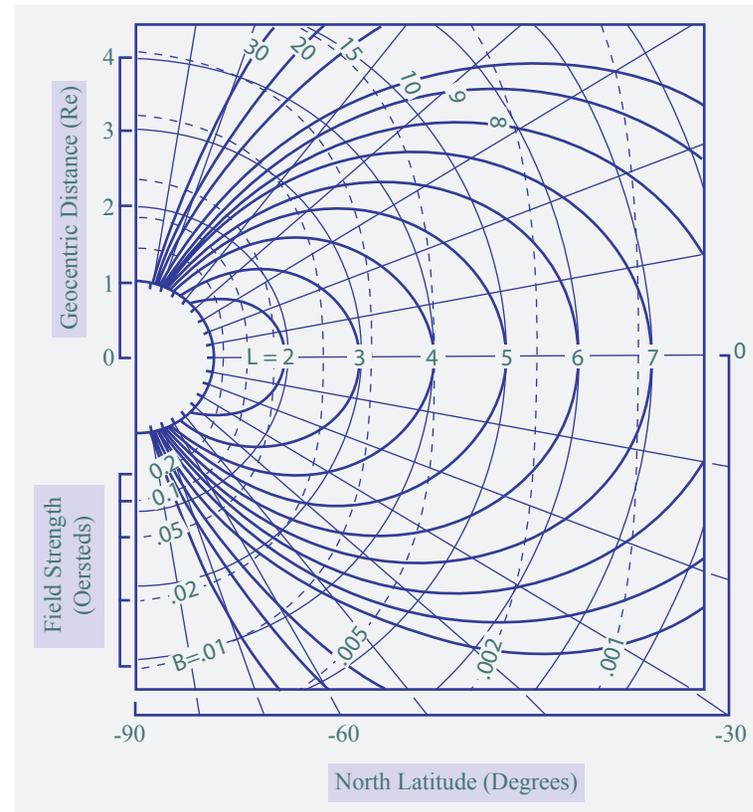
# GEOMAGNETIC COORDINATE SYSTEMS

## Geomagnetic



B

## B - L



L=8

# GEOMAGNETIC FIELD

## Magnitude Formula/Models

Tilted dipole ( $11^\circ$  from geographic north)

$$B_i(r, \theta_m, \phi_m) = -\frac{M}{r^3} (3 \cos^2(\theta_m) + 1)^{1/2} \quad \text{at LEO}$$

where

$$M = 0.311 \times 10^{-4} \quad \text{T-R}_e^3$$

$$= 7.9 \times 10^{15} \quad \text{T - m}^3$$

$$B_r = -\frac{M}{r^3} 2 \cos \theta_m$$

$$B_{\theta_m} = -\frac{M}{r^3} \sin \theta_m$$

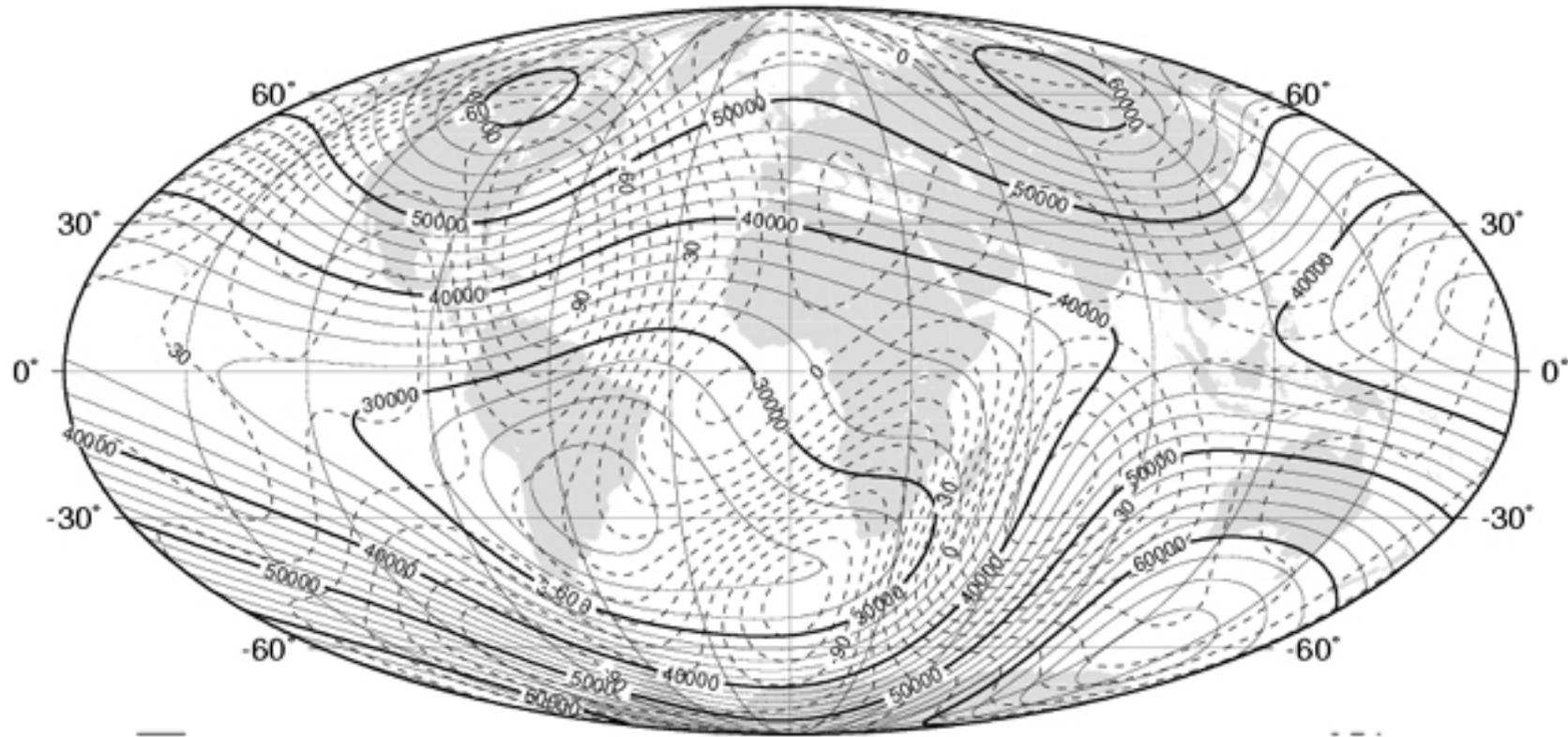
$$B_{\phi_m} = 0$$

International Geomagnetic Reference Field  
1987 (IGRF1987)

# FIELD VALUES

- Minimum (near equator) =  $0.25 \times 10^{-4}$  T
- Maximum (near polar caps) =  $0.50 \times 10^{-4}$  T
- Two peaks near north pole
- Two minimum near equator
- Largest minima is known as South Atlantic Anomaly
  - Much higher radiation exposure at LEO
- Geomagnetic storms impose variations of  $0.01 \times 10^{-4}$  T

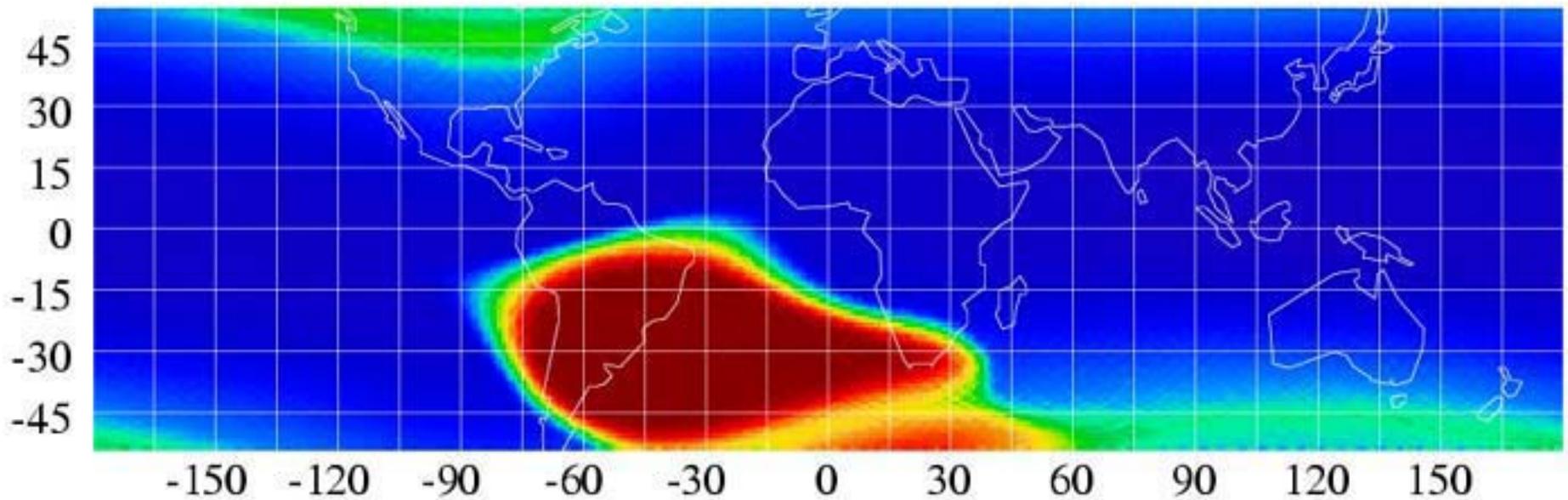
# TOTAL FIELD INTENSITY



Units (Total Intensity) : nanoTeslas  
Contour Interval : 2000 nanoTeslas  
Map Projection : Mollweide

Units (Annual Change) : nanoTeslas/yr  
Contour Interval : 10 nanoTeslas/yr  
Map Projection : Mollweide

# SOUTH ATLANTIC ANOMALY



Reduced protection in SAA allows greater effect of high energy particles - electronic upsets, instrument interference.<sup>28</sup>

# PLASMA EFFECTS OVERVIEW

- Plasma is a gas made up of ions and free electrons in roughly equal numbers.
- Causes
  - Electromagnetic Interference
  - Spacecraft charging & arcing
  - Material effects
- Effects
  - Avionics - Upsets from EMI
  - Power - floating potential, contaminated solar arrays, current losses
  - GN & C - torques from induced potential
  - Materials - sputtering, contamination effects on surface materials

# PLASMA EFFECTS (cont.)

- Effects continued

Optics systems - contamination changes properties of surface materials.

Propulsion - Thruster firings change/shift the floating potential by contacting the plasma.

# PLASMA GENERALIZATION

- Plasma is caused by UV, EUV, X-ray photoelectric effect on atmospheric molecules.
  - Breaks diatomic molecule bonds.
  - Ejects electrons from outer shells.
- As UV, EUV, X-ray penetrate the atmosphere, ion density increases with atmospheric density until most UV, EUV have been absorbed (>60 Km altitude).  
Varies dramatically with altitude, latitude, magnetic field strength, time of day and solar activity.
- Electrically charged region of atmosphere is called the ionosphere.
- Gas in ionosphere is called ionospheric plasma.

# LEO PLASMA ENVIRONMENT

- Balance between increasing density and increasing absorption leads to formation of ionization layers.

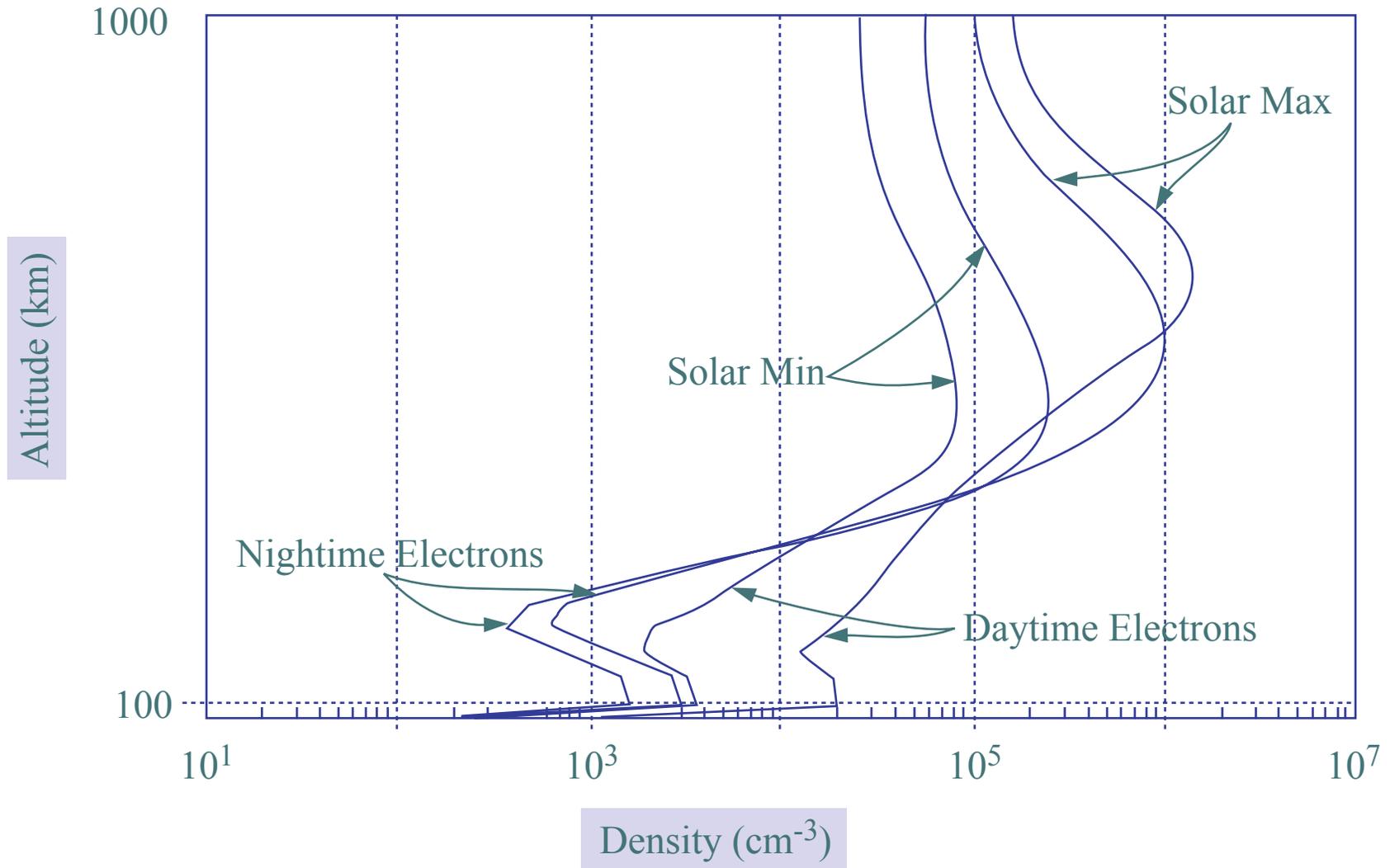
F - layer                      150 km - 1000 km

E - layer                      100 km - 150 km

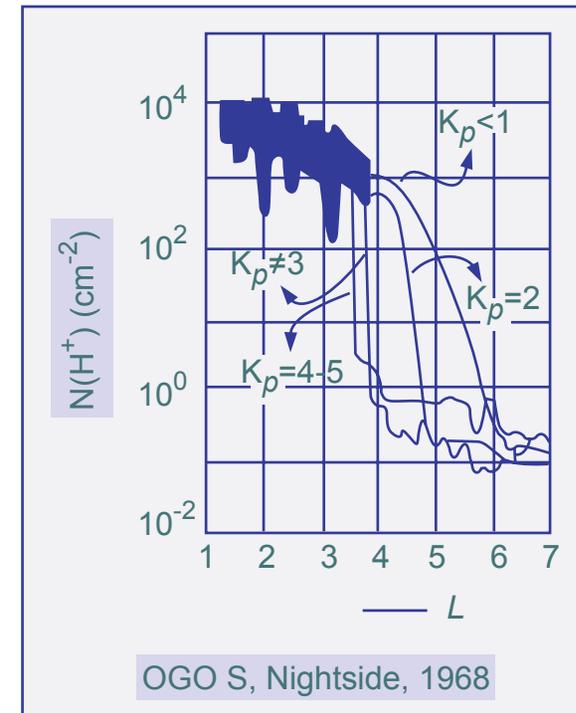
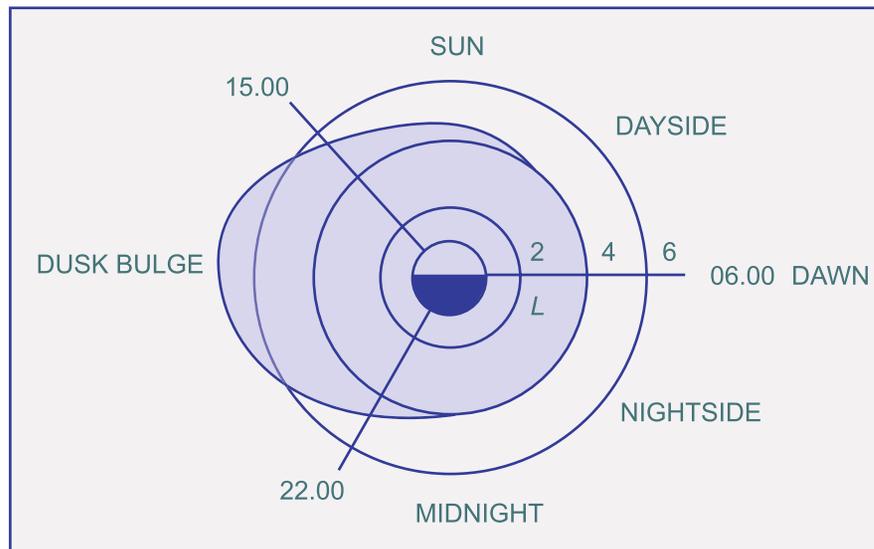
D - layer                      60 km - 100 km

- Transition region from ion-free atmosphere to fully ionized region called the plasmasphere.
- Plasmasphere ion densities peak at  $10^{10}/\text{m}^3$  to  $10^{11}/\text{m}^3$  at 1000 km
  - Drops to  $10^9/\text{m}^3$  at its boundary
- Outer boundary called plasmopause
  - Density drops to  $10^5/\text{m}^3$  to  $10^6/\text{m}^3$
  - Height is  $\sim 4 R_E$  between 0000 and 1800 hours
  - Expands to  $\sim 7 R_E$  during the local dusk (dusk bulge)

# ELECTRON DENSITY



# PLASMAPAUSE HEIGHT VS LOCAL TIME



$K_p$  is Magnetic Activity Index

# ION CONCENTRATIONS

– Similar to neutral atmosphere

D - layer       $\text{NO}^+/\text{O}^+$

E - layer       $\text{O}^+$

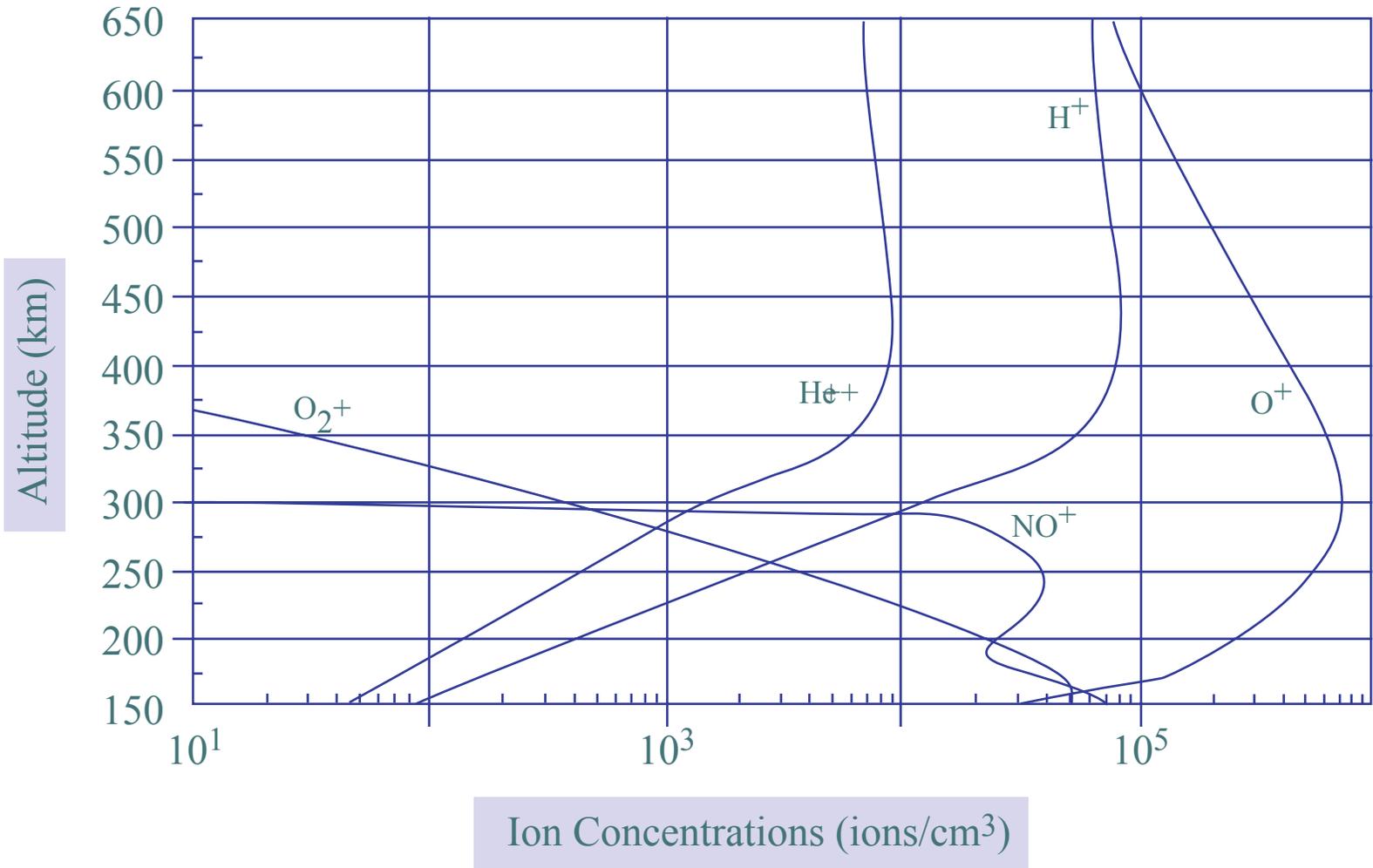
F - layer       $\text{O}^+/\text{H}^+$

-Daytime F layer density peaks at  $10^{12}/\text{m}^3$  (300 km)

-Nighttime F-layer density drops to  $10^{11}/\text{m}^3$  (500 km)

– Composition transitions from  $\text{O}^+$  to  $\text{H}^+$

# ION CONCENTRATIONS (cont.)



# PLASMA TEMPERATURES

Increases from  $\sim 100\text{K}$  at 50 - 60 km to

2000 - 3000K above 500 km

Electron temperature  $T_e = 4000\text{K} - 6000\text{K}$

Ion temperature  $T_i = 2000\text{K} - 3000\text{K}$

Density much higher at solar maximum due to higher UV/EUV fluxes.

# LEO PLASMA ENVIRONMENT MODELS

International Reference Ionosphere (IRI)

-Outputs - electron density  $n_e$

- ion composition  $n_i$

- Temperature  $T_e$ ,  $T_i$

-Inputs (latitude, longitude, altitude, solar activity (R), time).

Available at :

<http://nssdc.gsfc.nasa.gov/space/model/ionos/iri.html>

“Ionospheric models” Carlson, Schunk, Heelis, Basu

# RADIO FREQUENCY TRANSMISSIVITY

- Plasma transitions from a perfect conductor to perfect dielectric as a function of frequency.
- Plasma frequency  $\omega_{pe} = \left( \frac{n_e e^2}{\epsilon_0 m} \right)^{\frac{1}{2}}$
- Dielectric constant  $\epsilon = \epsilon_0 \left( 1 - \left( \frac{\omega_{pe}}{\omega} \right)^2 \right)$
- For  $\omega \gg \omega_{pe}$  the plasma appears like free space
- For  $\omega \sim \omega_{pe}$  electromagnetic waves cannot propagate
  - Transmissions from below are reflected
  - Transmissions from within are absorbed
- For  $\omega \gtrsim \omega_{pe}$  random variations in  $n_e$  can cause random delays and phase shifts

# SPACECRAFT CHARGING

- At LEO spacecraft become negatively charged
  - Plasma is dense but low energy
  - Orbital velocity is higher than ion thermal velocity
  - Lower than electron thermal velocity
  - Electrons impact all surfaces
  - Ions impact ram surfaces only
- Geo spacecraft charge during magnetospheric substorms between longitudes corresponding to midnight and dawn
- Biased surfaces (solar arrays) influence the floating potential

# CHARGING EFFECTS

- Instrument reading bias
- Arcing-induced EMI, electronics upsets
- Increased current collection
- Re-attraction of contaminants
- Ion sputtering, accelerated erosion of materials

Spacecraft must be designed to keep differential charging below the breakdown voltages or must tolerate the effects of discharges.

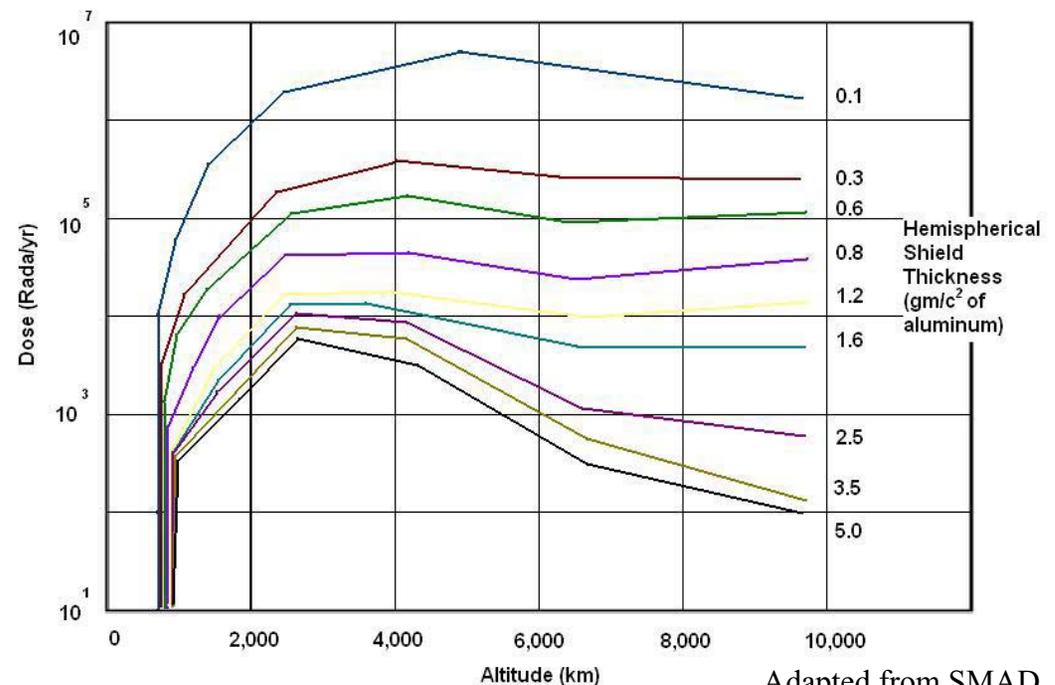
# RADIATION

- Most radiation effects occur by energy deposition
  - Function of both energy, type of particle and material into which energy is deposited.
- Definitions

1 rad (Si) = 100 ergs/gm into Silicon

1 Cray (Si) = 1 J/kg into Si

1 rad (Si) =  $10^{-4}$  Cray



Adapted from SMAD.

# RADIATION DAMAGE THRESHOLDS

In many materials the total dose of radiation is the most critical issue. In other circumstances the time over which the dose is received is equally important.

<u>Material</u>	<u>Damage Threshold (rad)</u>
Biological Matter	$10^1 - 10^2$
Electrical Matter	$10^2 - 10^4$
Lubricants, hydraulic fluid	$10^5 - 10^7$
Ceramics, glasses	$10^6 - 10^8$
Polymeric materials	$10^7 - 10^9$
Structural metals	$10^9 - 10^{11}$

# SPACECRAFT EFFECTS

- High energy particles travel through spacecraft material and deposit kinetic energy
  - Displaces atoms.
  - Leaves a stream of charged atoms in their wake.
- Reduces power output of solar arrays
- Causes sensitive electronics to fail
- Increases sensor background noise
- Radiation exposure to crews

# HIGH ENERGY RADIATION

- Definition

For Electrons  $E > 100 \text{ keV}$

For protons and heavy ions  $E > 1 \text{ MeV}$

- Sources

- Van Allen Belt → (electrons and protons) (trapped radiation)
- Galactic cosmic rays → interplanetary protons and ionized heavy nuclei
- Protons associated with solar proton events

# VAN ALLEN BELTS

- Torodial belts around the earth made up of electrons and ions (primarily protons) with energies  $> 30$  keV.
- Two big zones
  - Inner belt  $\sim 1000$  Km  $\rightarrow$  6000 km altitude
    - Protons  $E > 10$ 's of MeV
    - Electrons  $E \sim 1 - 10$  MeV
  - Outer belt 10,000 - 60,000 km
    - Electrons  $E \sim 0.04 - 4.5$  MeV

# VAN ALLEN BELTS (cont.)

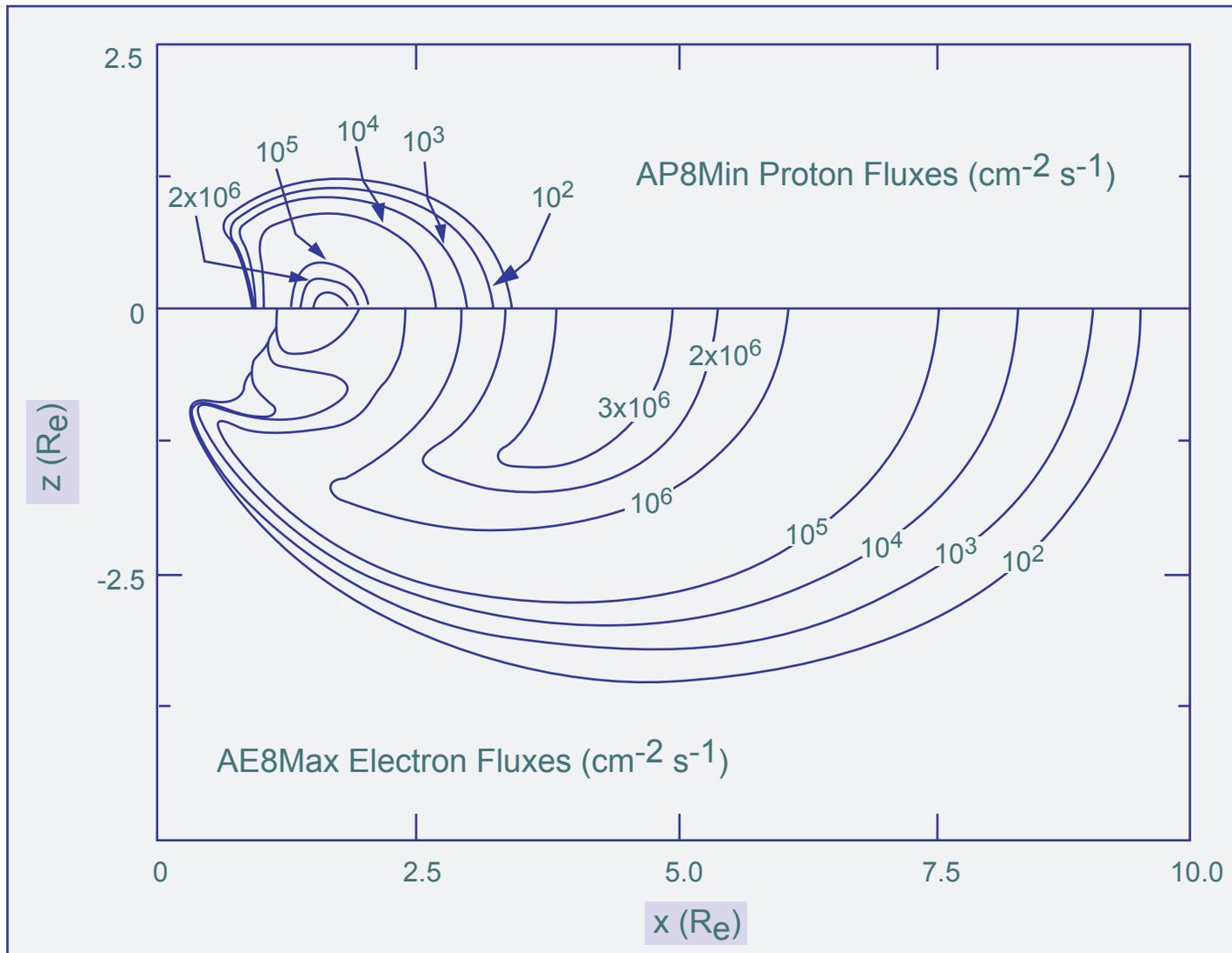
- Sources
  - acceleration of lower-energy particles by magnetic storm activity
  - trapping of decay products produced by cosmic ray collisions with the atmosphere
  - solar flares

# CONCENTRATION MECHANISM

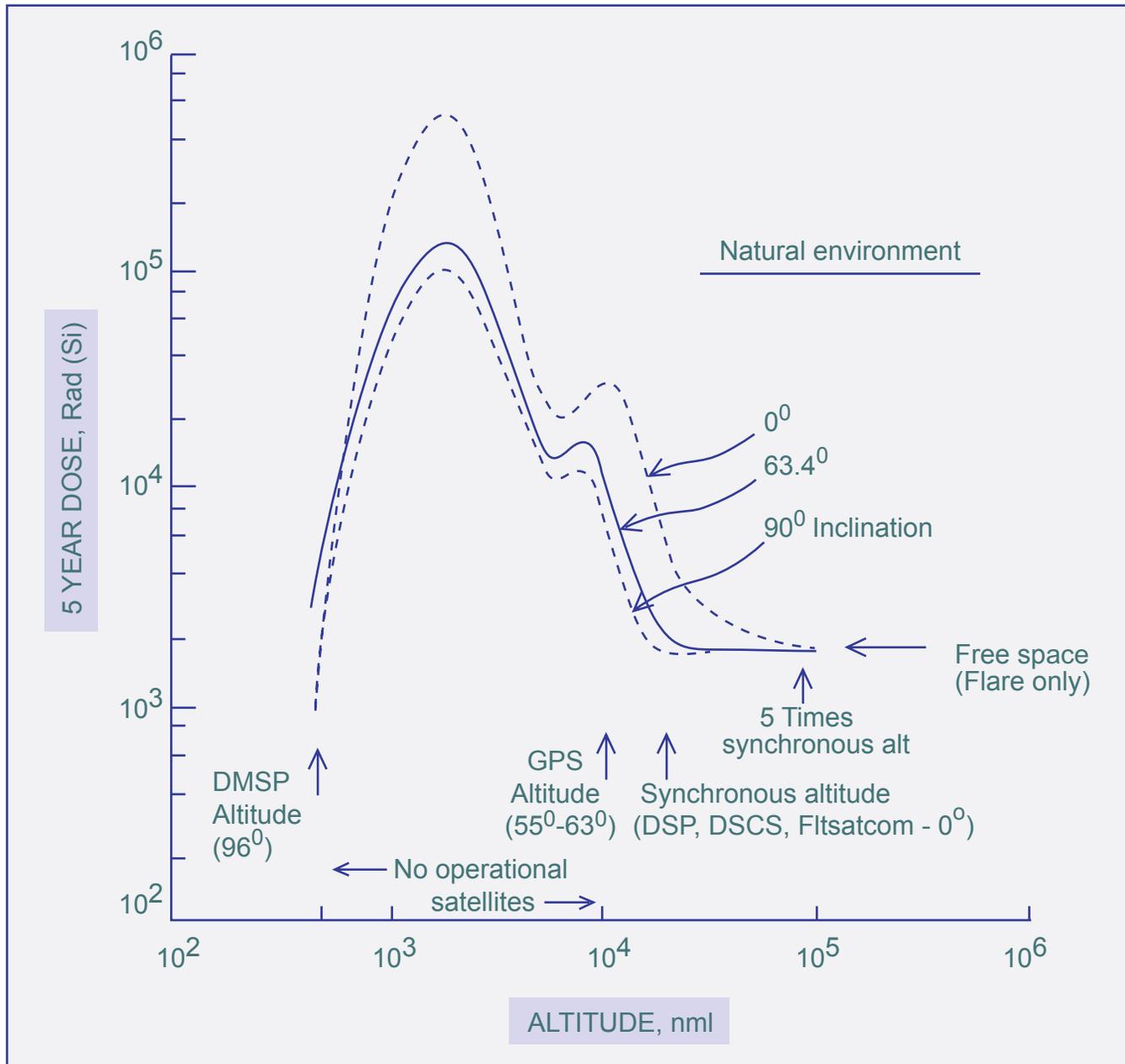
- Earth's magnetic field concentrates on large fluxes of electrons, protons and some heavy ions.
- Radiation belt particles spiral back and forth along magnetic field lines.
  - Ionizing radiation belts reach lowest altitude of the eastern coast of the eastern coast of South America (SAA).

(Image removed due to copyright considerations.)

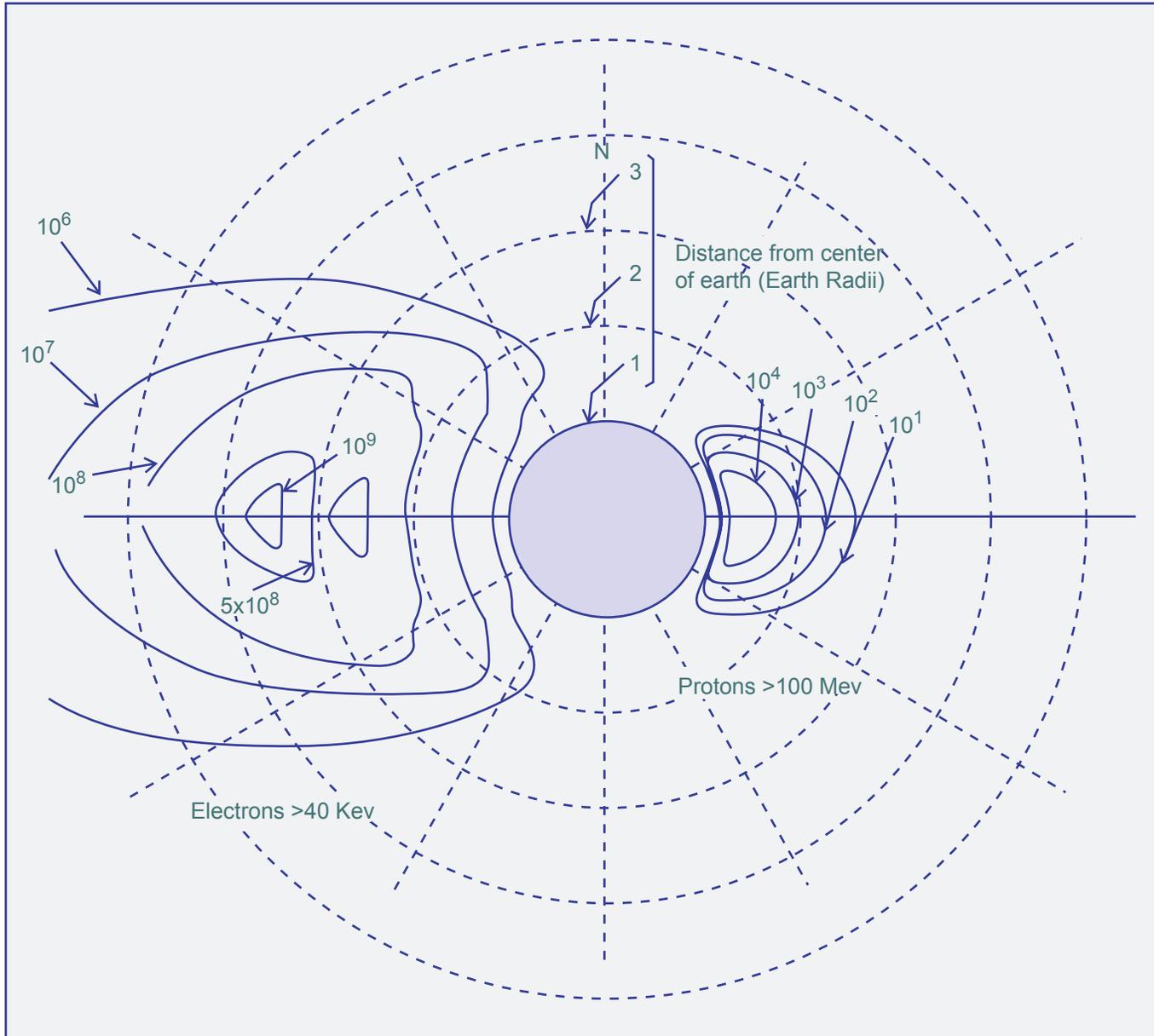
# ELECTRON AND PROTON FLUXES



# 5 YEAR DOSE



# TRAPPED RADIATION BELTS

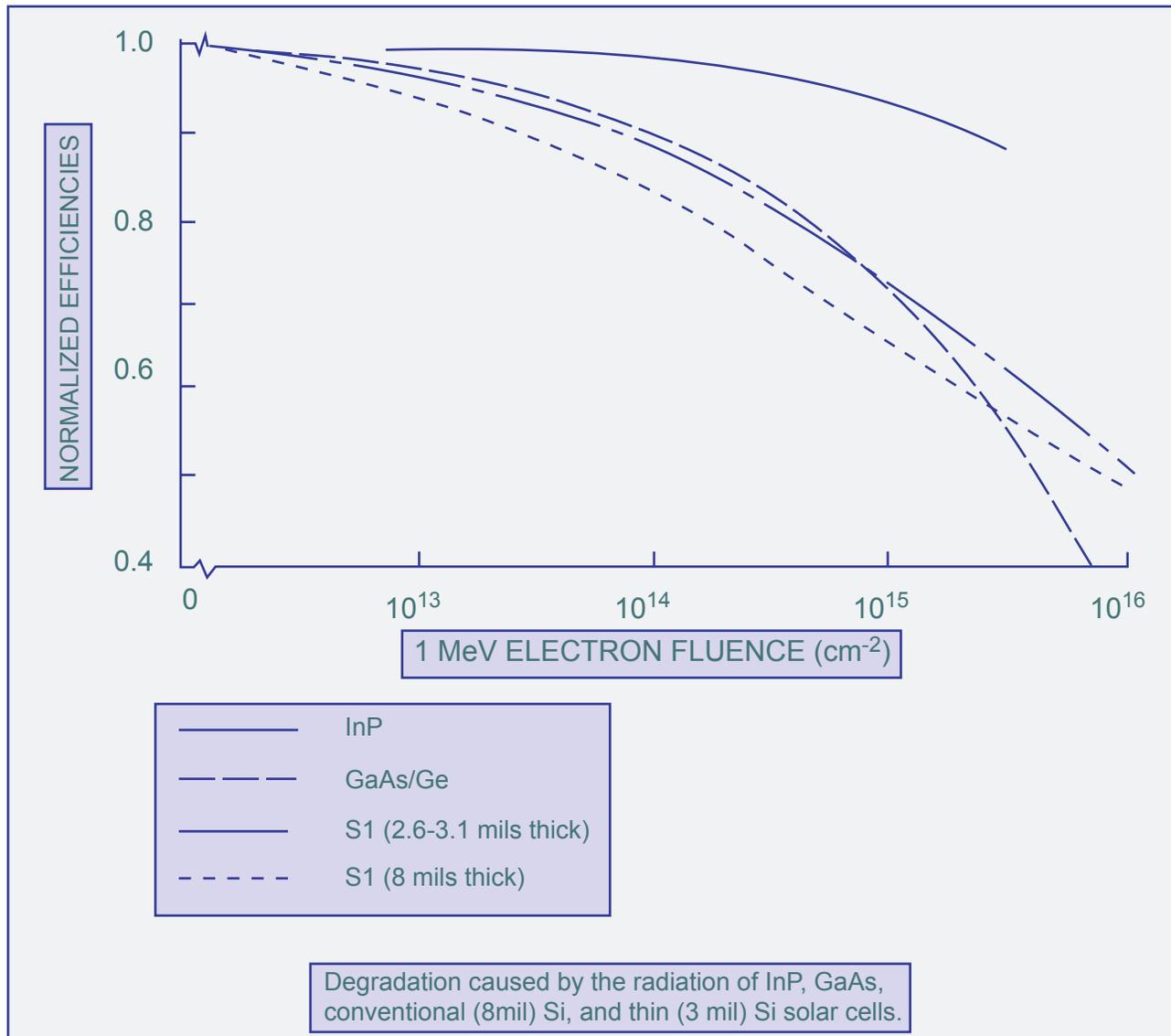


# VAN ALLEN BELT

## RADIATION STABILITY

- Inner belt
  - Fairly stable with changes in solar cycle
  - May change by a factor of three as a result of geomagnetic storms loading in high energy electrons.
- Outer belt
  - Electron concentrations may change by a factor of 1000 during geomagnetic storms.
- Standard Models (AP8 protons) and (AE8 electrons)
  - Require B, L and whether solar min/solar max
  - Provide omni-directional fluxes of protons  $50 \text{ keV} < E < 500 \text{ MeV}$  and electrons  $50 \text{ keV} < 7 \text{ MeV}$

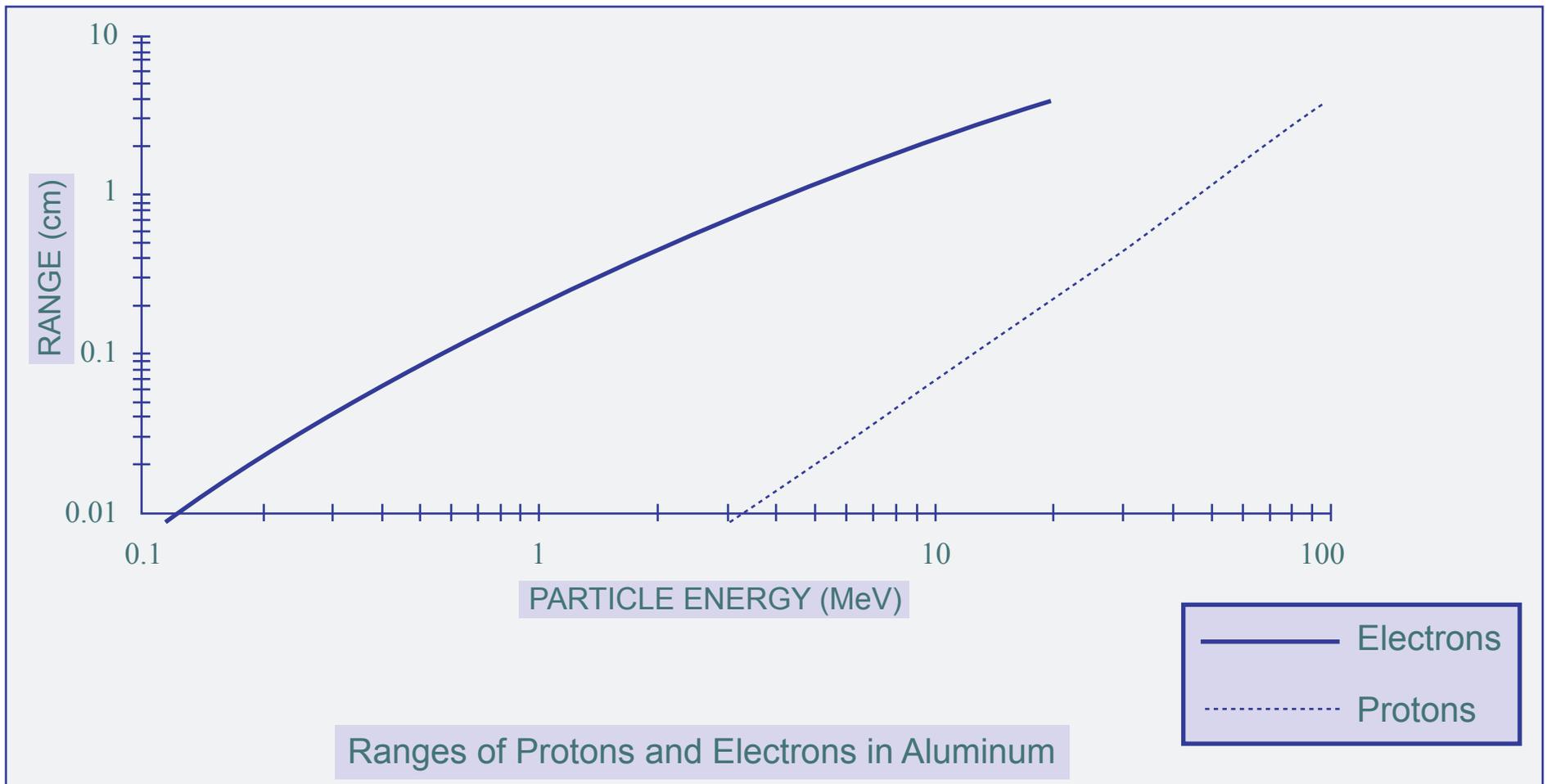
# SOLAR CELL DEGRADATION



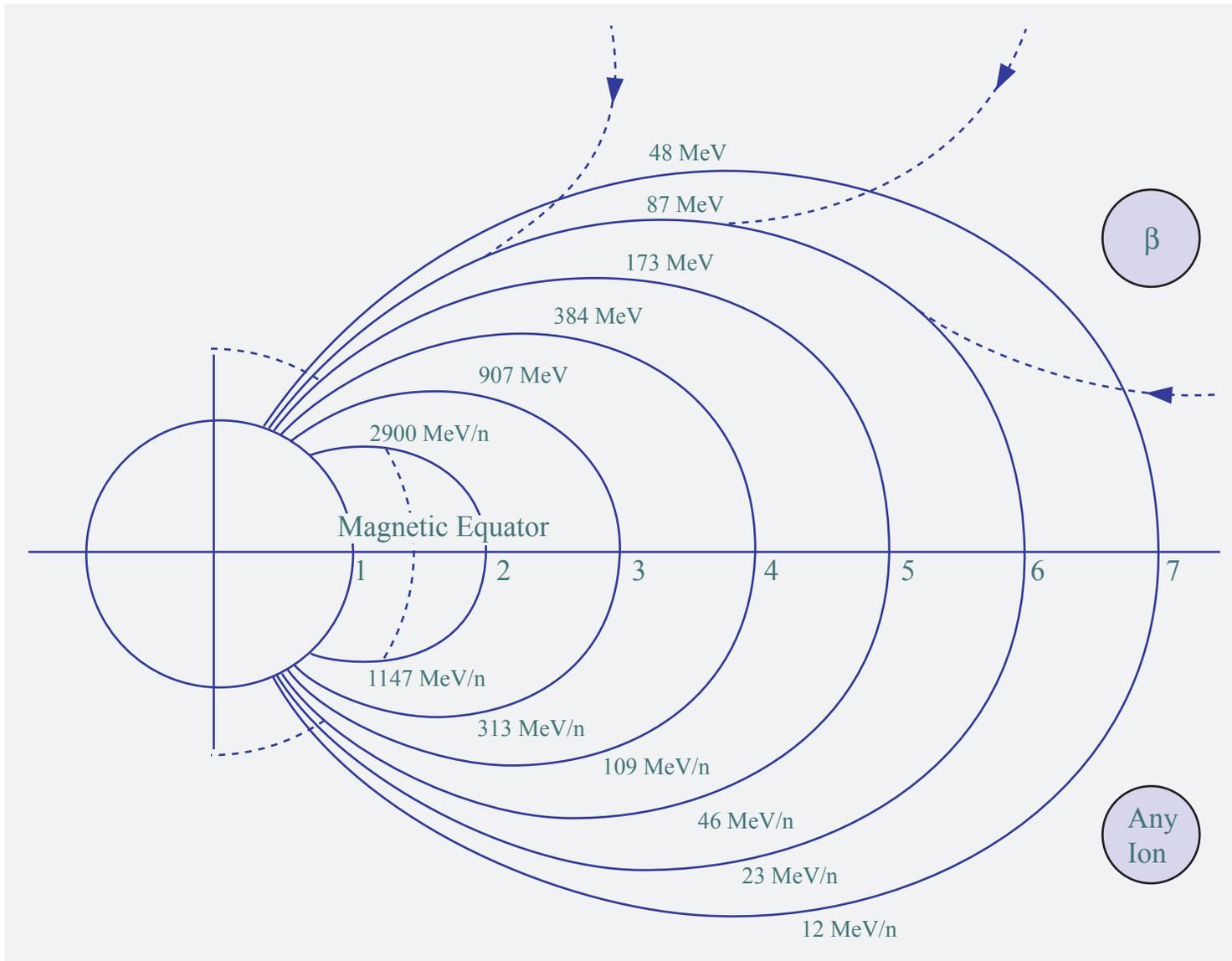
# GALACTIC COSMIC RAYS

- Primarily interplanetary protons and ionized heavy nuclei
  - $1 \text{ MeV} < E < 1 \text{ GeV}$  per nucleon
  - Cause Single Event Upsets (SEU)
- Sources are outside the solar system
  - other solar flares
  - nova and supernova explosions
  - quasars

# PARTICLE RANGE



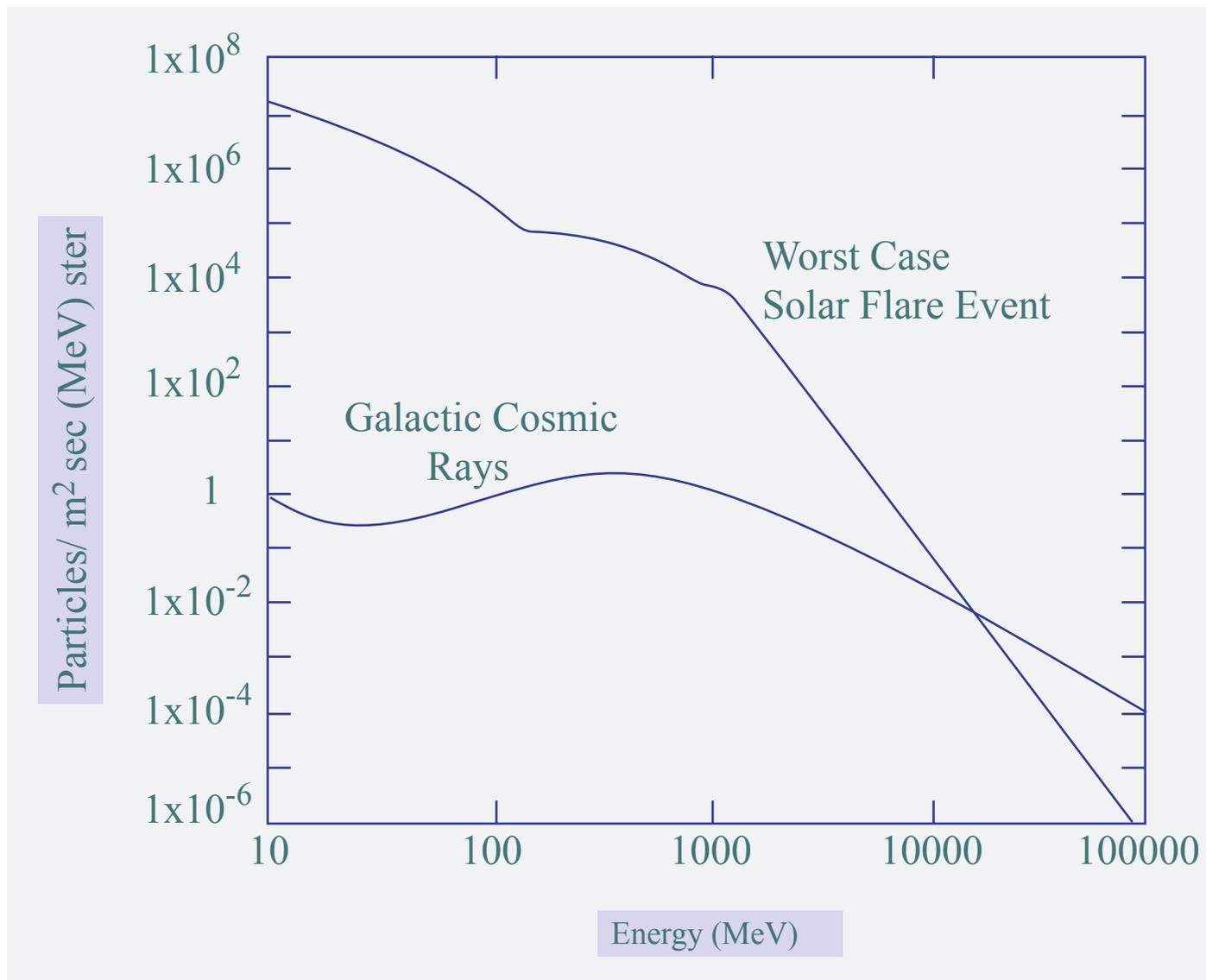
# MAGNETIC SHIELDING



# SOLAR PROTON EFFECTS

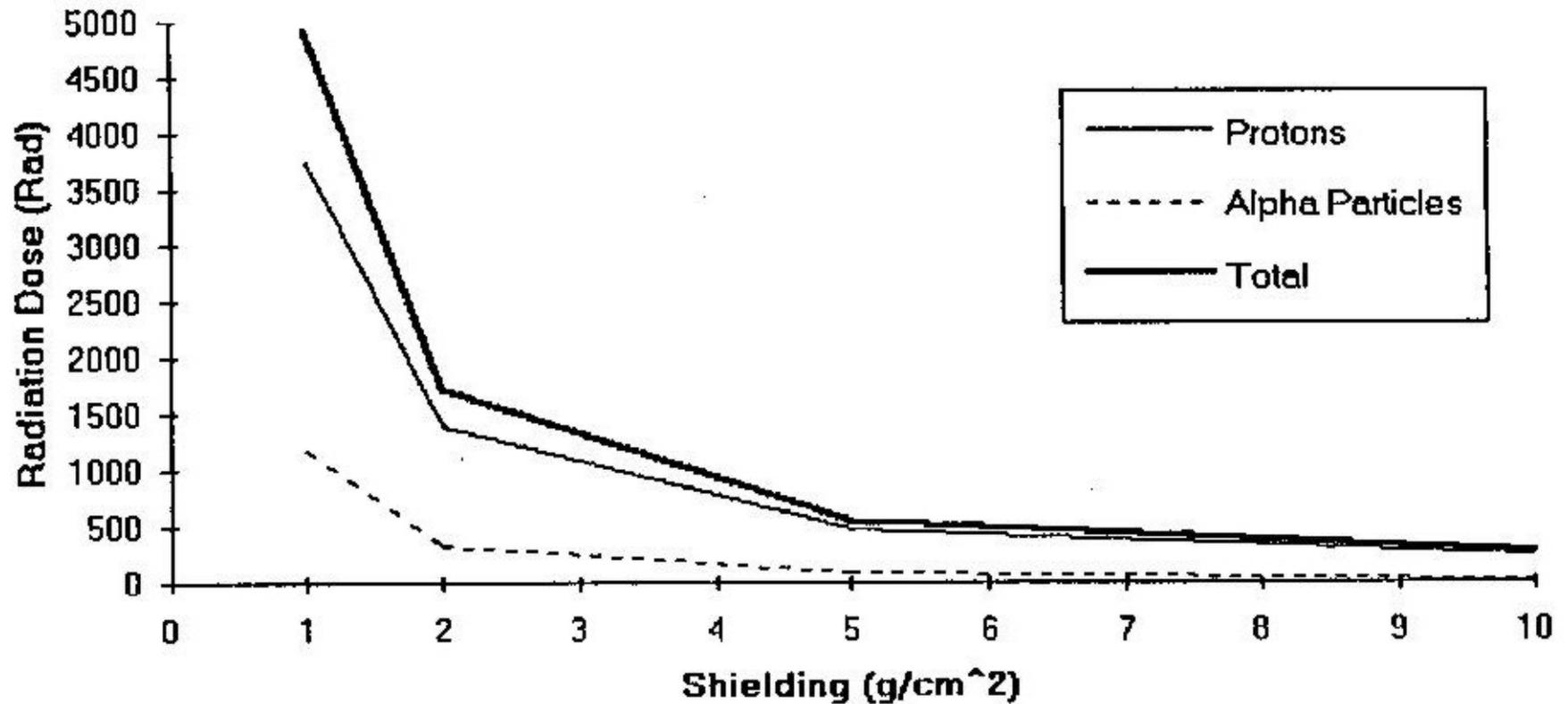
- Solar flares often eject high energy hydrogen and other nuclei
  - $1 \text{ MeV} < E < 10 \text{ GeV/nucleon}$
  - At low energies the number can be much greater than galactic cosmic radiation level
- Solar events are sporadic but correlate somewhat with the solar cycle
- These events make a Mars Mission hazardous

# PARTICLE ENERGY

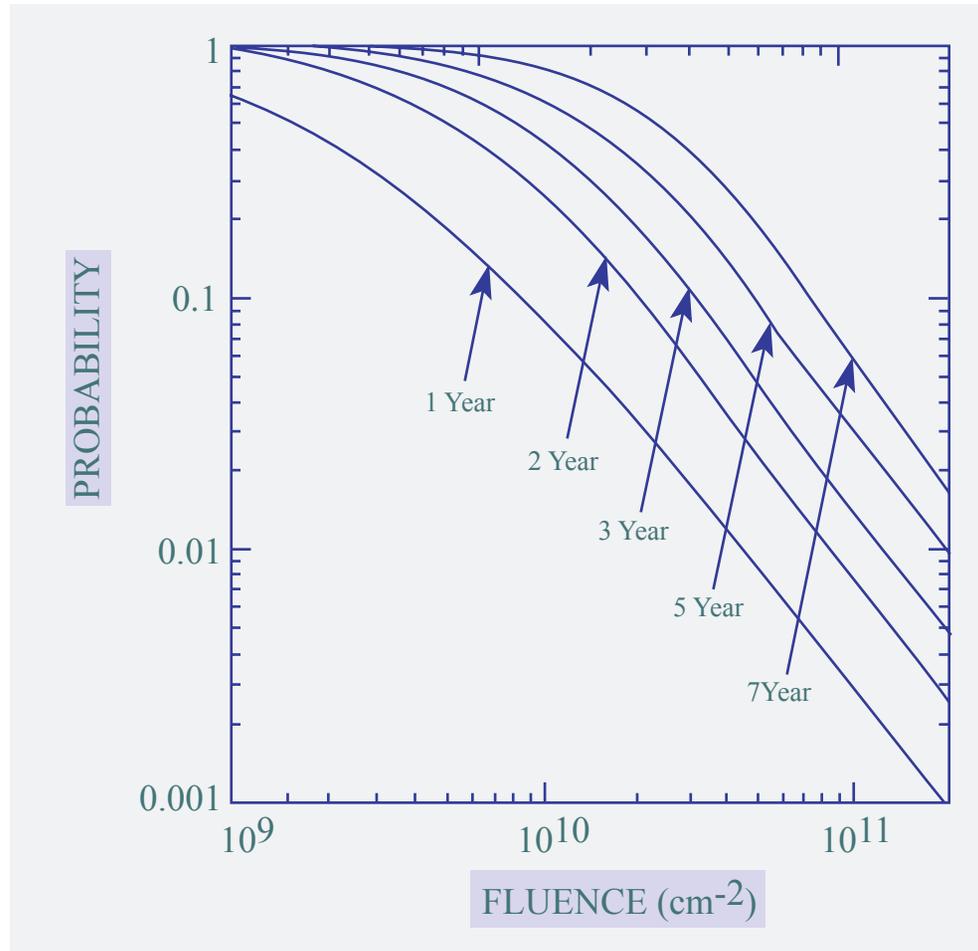


# SOLAR PROTON DOSE

Average Radiation Dose From Large Solar Proton Event



# FEYNMAN MODEL



Based on data from 1963 to 1991

# ELECTROMAGNETIC RADIATION

- Radio
  - 1 - 10 MHz galactic electromagnetic radiation
  - terminal noise
  - not significant for single event environment
- Visible/IR
  - solar flux
  - heating
- UV/EUV/X-ray
  - EUV @ 100 to 1000 Å is significant for surface chemistry

# References

- Wertz, James R. and Wiley JH. Larson, Space Mission Analysis and Design, Third edition, Microcosm Press, El Segundo CA 1999
- Pisacane, Vincenti and Robert C. Moore, Fundamentals of Space Systems, Oxford University Press, NY, 1994.
- [http://nssdc.gsfc.nasa.gov/space/model/models\\_home.html](http://nssdc.gsfc.nasa.gov/space/model/models_home.html)
- <http://nssdc.gsfc.nasa.gov/space/model/magnetos/igrf.html>